

Systematic investigation of the **Hoyle-analog states** in light nuclei.

V.S. Vasilevsky ¹, K. Katō ², N. Takibayev ³

¹Bogolyubov Institute for Theoretical Physics, Kiev, Ukraine

²Nuclear Reaction Data Centre, Faculty of Science, Hokkaido University,
Sapporo, Japan

³Al-Farabi Kazakh National University, Almaty, Kazakhstan

4th International Workshop
“State of the Art in Nuclear Cluster Physics”,
Galveston, Texas, USA, May 13-18, 2018.

1 Introduction

- Motivations
- Hoyle state
- Three-cluster AMHHB Model

2 Theoretical analysis

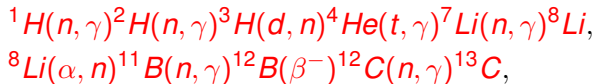
- Hoyle states in light nuclei
- Hoyle-analog states in ^{11}B and ^{11}C
- Hoyle states in ^{10}B
- Hoyle states in ^9Be and ^9B
- Shape of triangles
- Hoyle-analog states in light nuclei

Motivations

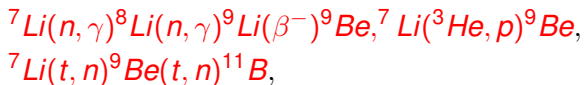
- Synthesis of light nuclei in stellar environment
 - Resonance states in three-cluster continuum
 - Nature of resonance states in three-cluster continuum
 - **Hoyle-analog states**

Inhomogeneous BBN reactions

Inhomogeneous BBN reactions involve reaction chains such as

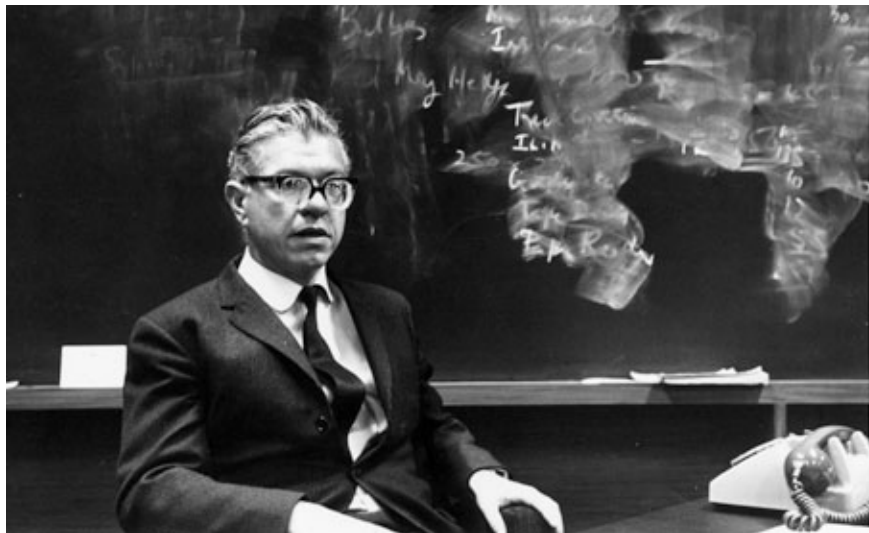


and



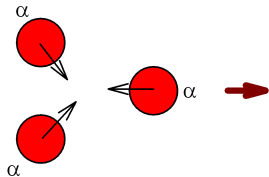
C.A. Bertulani, T. Kajino. "Frontiers in nuclear astrophysics", Progr. Part. Nucl. Phys., **89**, 56, 2016.

Fred Hoyle and Hoyle state

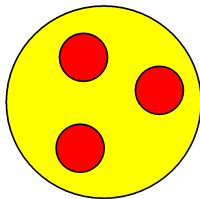


Synthesis of ^{12}C . Hoyle state

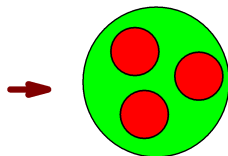
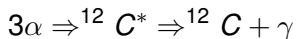
Triple collision



Hoyle state

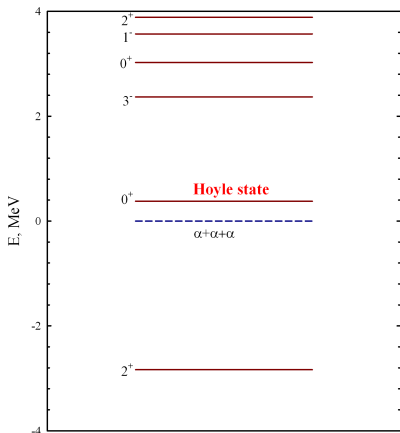
 $^{12}\text{C}^*$

Ground state

 ^{12}C 

F. Hoyle. "On Nuclear Reactions Occurring in Very Hot Stars. I. The Synthesis of Elements from Carbon to Nickel." *Astrophys. J. Suppl.*, **1**, 121, 1954.

^{12}C . Hoyle state



- Hoyle state is the 0^+ resonance state in ^{12}C
- Energy of the Hoyle state is $E = 0.4 \text{ MeV}$ above the $\alpha + \alpha + \alpha$ threshold or $E = 7.65 \text{ MeV}$ above the ^{12}C ground state
- Very small width of the Hoyle state $\Gamma = 8.5 \text{ eV}$, compare with width of the 0^+ resonance state in ^8Be $\Gamma = 5.57 \text{ eV}$

- F. Hoyle. "On Nuclear Reactions Occurring in Very Hot Stars. I. The Synthesis of Elements from Carbon to Nickel." *Astrophysical Journal Supplement*, vol. 1, p.121 (1954).
- M. Freer, H. O. U. Fynbo, "The Hoyle state in ^{12}C ", *Prog. Part. Nucl. Phys.*, **78**, pp. 1-23, 2014

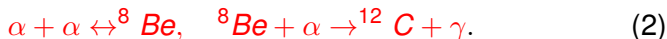
¹²C. Hoyle state

Two quotations from F. Hoyle paper [1].

- 1 It was pointed out some years ago by **Bethe** [2] that effective element-building inside stars must proceed, in the absence of hydrogen, by triple collisions as a starting point:



- 2 It is convenient to replace reaction (1) by



This is a permissible step, since the lifetime of the unstable ⁸Be is appreciably longer than the time required for nuclear collision of two α particles; that is, longer than the α particle radius divided by the relative velocity.

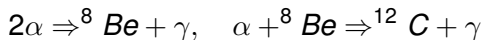
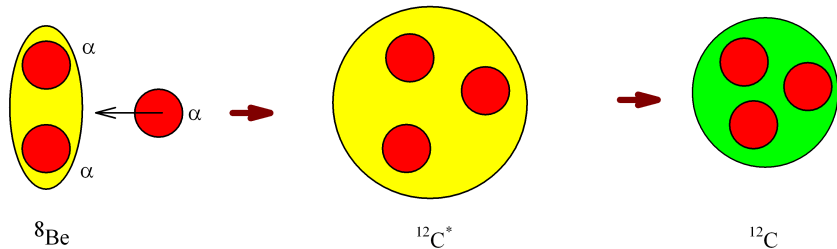
- 1 F. Hoyle. "On Nuclear Reactions Occurring in Very Hot Stars. I. The Synthesis of Elements from Carbon to Nickel." *Astrophys. J. Suppl.*, **1**, 121, 1954.
- 2 H.A. Bethe. *Phys. Rev.*, "Energy Production in Stars." **55**, 434, 1939.

Synthesis of ^{12}C . Hoyle state

Two-body collision

Hoyle state

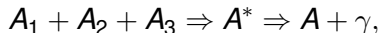
Ground state



The 0^+ resonance in ^8Be : $E=91.8$ keV, $\Gamma = 5.57 \pm 0.25$ eV, $\tau \approx 10^{-16}$ sec.

Synthesis of light nuclei in triple collision

General scheme



where A^* denotes excited (**resonance**) state of a compound system.
First step, excitation of resonance state



Second step, transition from resonance to bound state $A^* \Rightarrow A + \gamma$.

- ${}^9\text{Be} = \alpha + \alpha + n$
- ${}^9\text{B} = \alpha + \alpha + p$
- ${}^{10}\text{B} = \alpha + \alpha + d$
- ${}^{11}\text{B} = \alpha + \alpha + {}^3\text{H}$
- ${}^{11}\text{C} = \alpha + \alpha + {}^3\text{He}$

The Algebraic Model for scattering with the Hyperspherical Harmonics Basis (AMHHB)

- V. Vasilevsky, A. V. Nesterov, F. Arickx, J. Broeckhove. *Phys. Rev. C*, vol. **63**, 034606 (16 pp), 2001.
- V. Vasilevsky, A. V. Nesterov, F. Arickx, J. Broeckhove. *Phys. Rev. C*, vol. **63**, 034607, 2001.
- V. Vasilevsky, A. V. Nesterov, F. Arickx, and J. Broeckhove. *Phys. Rev. C*, vol. **63**, 064604, 2001.
- J. Broeckhove, F. Arickx, P. Hellinckx, V. S. Vasilevsky, A. V. Nesterov. *J. Phys. G Nucl. Phys.*, vol. **34**, 1955, 2007.
- V. S. Vasilevsky, K. Katō, N. Zh. Takibayev. *Phys. Rev. C*, **96**, 034322, 2017.

Alternative methods

- Complex Scaling Method (K. Katō, et al)

T. Myo, Y. Kikuchi, H. Masui and K. Katō, "Recent development of complex scaling method for many-body resonances and continua in light nuclei", Prog. Part. Nucl. Phys., 79, 1, 2014.

- Antisymmetrized Molecular Dynamics (H. Horiuchi, Y. Kanada-En'yo, et al)

- Fermionic Molecular Dynamics (H. Feldmeier, T. Neff)

- Microscopic R-Matrix Method (D. Baye, P. Descouvemont)

- No-Core Shell Model (P. Navrátil, S. Quaglioni)

Three-cluster model and Hyperspherical Harmonics Method

- V. Vasilevsky, A. V. Nesterov, F. Arickx, J. Broeckhove. *Phys. Rev. C*, vol. **63**, 034606, 2001.
- Korenov, S. and Descouvemont, P. *Nucl. Phys. A*, **740**, 249, 2004.
- S. Quaglioni, C. Romero-Redondo, and P. Navrátil, *Phys. Rev. C*, **88**, 034320, 2013.

Quest for the Hoyle-analog states

- 1 Kanada-En'yo, Y. "Negative parity states in ^{11}B and ^{11}C and the similarity with ^{12}C ", *Phys. Rev. C*, **75**, 024302, 2007.
- 2 Kanada-En'yo, Y. and Suhara, T. and Kobayashi, F., "Cluster states in ^{11}B , ^{11}C and ^8He and their similarity to ^{12}C ", *J. Phys. Confe. Ser.*, **321**, 012009, 2011.
- 3 Yamada, T. and Funaki, Y. " $\alpha + \alpha + t$ cluster structures and $^{12}\text{C}(0_2^+)$ -analog states in ^{11}B ", *Phys. Rev. C*, **82**, 064315, 2010.
- 4 Yamada, T. and Funaki, Y. "Three-Body Cluster Structures and Hoyle-Analogue States in ^{11}B ", *Progr. Theor. Phys. Suppl.*, **196**, 388, 2012.
- 5 Y. Chiba, M. Kimura "Hoyle-analogue state in ^{13}C studied with Antisymmetrized Molecular Dynamics", arXiv:1801.00562, 2018.

Three-cluster wave function

Wave function for three s -shell clusters in the LS coupling scheme:

$$\psi^J = \hat{\mathcal{A}} \{ [\Phi_1(A_1) \Phi_2(A_2) \Phi_3(A_3)]_S f_L(\mathbf{x}, \mathbf{y}) \}_J$$

where

- $\Phi_\nu(A_\nu)$ is a shell-model wave function for the internal motion of α -particle ($\nu = 1, 2, 3$) (fixed)
- $f(\mathbf{x}, \mathbf{y})$ is a wave function of inter-cluster motion (to be determined)
- $\hat{\mathcal{A}}$ is the antisymmetrization operator.

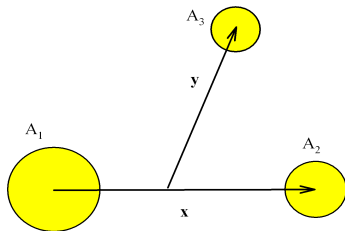
Hyperspherical coordinates

Hyperspherical coordinates: $\{\mathbf{x}, \mathbf{y}\} \Rightarrow \{\rho, \Omega\}$, $\Omega = \{\theta, \hat{\mathbf{x}}, \hat{\mathbf{y}}\}$

$$\rho = \sqrt{\mathbf{x}^2 + \mathbf{y}^2} \quad \text{hyper-radius,}$$

$$\hat{\mathbf{x}} = \mathbf{x}/|\mathbf{x}|, \quad \hat{\mathbf{y}} = \mathbf{y}/|\mathbf{y}| \quad \text{unit vectors}$$

$$|\mathbf{x}| = \rho \cos \theta, \quad |\mathbf{y}| = \rho \sin \theta$$



Expansion of wave function

$$f_L(\mathbf{x}, \mathbf{y}) = \sum_{l_x, l_y} f_{l_x, l_y; L}(x, y) \{ Y_{l_x}(\hat{\mathbf{x}}) Y_{l_y}(\hat{\mathbf{y}}) \}_L$$

$$= \sum_c \psi_c(\rho) \mathcal{Y}_c(\Omega) = \sum_c \sum_{n_\rho=0}^{\infty} C_{n_\rho, c} R_{n_\rho, K}(\rho, b) \mathcal{Y}_c(\Omega)$$

Hyperspherical Harmonics

Here $c = \{K, l_x, l_y, L\}$, $\mathcal{Y}_c(\Omega)$ is a hyperspherical harmonic

$$\mathcal{Y}_c(\Omega) = \chi_{K, l_x, l_y}(\theta) \cdot \{Y_{l_x}(\hat{\mathbf{x}}) Y_{l_y}(\hat{\mathbf{y}})\}_L$$

and $R_{n_\rho, K}(\rho, b)$ is a six-dimensional oscillator function

$$R_{n_\rho, K}(\rho, b) = (-1)^{n_\rho} \mathcal{N}_{n_\rho, K} r^K \exp\left\{-\frac{1}{2}r^2\right\} L_{n_\rho}^{K+3}(r^2),$$

$$r = \rho/b, \quad \mathcal{N}_{n_\rho, K} = b^{-3} \sqrt{\frac{2\Gamma(n_\rho + 1)}{\Gamma(n_\rho + K + 3)}},$$

and b is an oscillator length.

Hyperspherical Harmonics

Basis of the Hyperspherical Harmonics

$$\begin{aligned}\Psi_J &= \sum_{\mathbf{c}} \hat{\mathcal{A}} \{ \Phi_1(\mathbf{A}_1) \Phi_2(\mathbf{A}_2) \Phi_3(\mathbf{A}_3) \psi_{\mathbf{c}}(\rho) \mathcal{Y}_{\mathbf{c}}(\Omega) \} \\ &= \sum_{\mathbf{c}} \sum_{n_{\rho}=0}^{\infty} C_{n_{\rho}, \mathbf{c}} |n_{\rho}, \mathbf{c}\rangle\end{aligned}$$

where $\{C_{n_{\rho}, \mathbf{c}}\}$ is a wave function in the oscillator representation, $|n_{\rho}, \mathbf{c}\rangle$ is a three-cluster oscillator function

$$|n_{\rho}, \mathbf{c}\rangle = \hat{\mathcal{A}} \{ \Phi_1(\mathbf{A}_1) \Phi_2(\mathbf{A}_2) \Phi_3(\mathbf{A}_3) R_{n_{\rho}, \mathbf{c}}(\rho, \mathbf{b}) \mathcal{Y}_{\mathbf{c}}(\Omega) \}$$

Oscillator basis and linear algebraic equations

The many-particle Schrödinger equation

$$(\hat{H} - E) \Psi = 0$$

in oscillator representation

$$\sum_{\tilde{c}} \sum_{\tilde{n}_\rho=0}^{\infty} \left[\langle n_\rho, c | \hat{H} | \tilde{n}_\rho, \tilde{c} \rangle - E \langle n_\rho, c | \tilde{n}_\rho, \tilde{c} \rangle \right] C_{\tilde{n}_\rho, \tilde{c}} = 0$$

Numerical realization

$$\sum_{\tilde{c}=1}^{N_c} \sum_{\tilde{n}_\rho=0}^N \left[\langle n_\rho, c | \hat{H} | \tilde{n}_\rho, \tilde{c} \rangle - E \langle n_\rho, c | \tilde{n}_\rho, \tilde{c} \rangle \right] C_{\tilde{n}_\rho, \tilde{c}} = 0$$

plus proper boundary conditions (Details in V. Vasilevsky et al. *Phys. Rev. C*, vol. **63**, 034606, 2001).

Main stages of our calculations

- 1 By solving the system of linear equations, we obtain the scattering S matrix

$$\|S_{c,\tilde{c}}\|,$$

where $c, \tilde{c} = 1, 2, \dots, N_{ch}$

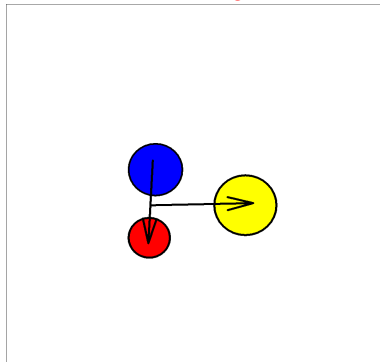
- 2 By analyzing behavior of the S matrix as a function of energy, we extract parameters of resonance states.
- 3 We study resonance wave functions in coordinate and oscillator representations

We are going to search for the Hoyle-analog states in light nuclei

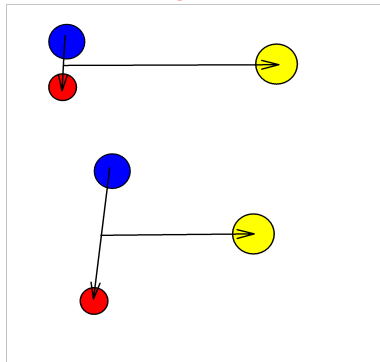
Reference	Nucleus	Configuration
Phys. Atom. Nucl., 77, 555, 2014	${}^9\text{Be}$	$\alpha + \alpha + n$
	${}^9\text{B}$	$\alpha + \alpha + p$
Phys. Rev., C96, 034322, 2017	${}^9\text{Be}$	$\alpha + \alpha + n$
	${}^9\text{B}$	$\alpha + \alpha + p$
Ukr. J. Phys., 59, 1065, 2014	${}^{10}\text{B}$	$\alpha + \alpha + d$
Ukr. J. Phys., 58, 544, 2013	${}^{11}\text{B}$	$\alpha + \alpha + {}^3\text{H}$
	${}^{11}\text{C}$	$\alpha + \alpha + {}^3\text{He}$
Phys. Rev., C85, 4607, 2012	${}^{12}\text{C}$	$\alpha + \alpha + \alpha$

Wave function in Oscillator representation.

Small N_{sh}

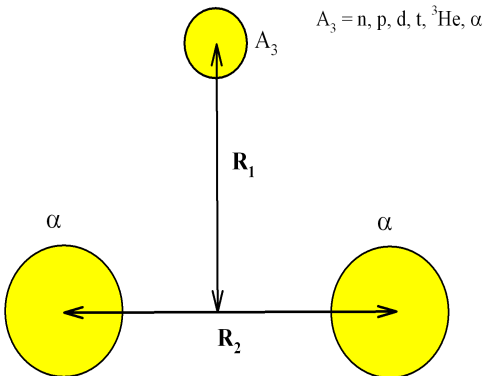


Large N_{sh}



Condensed information about wave function

$$W_{sh}(N_{sh}) = \sum_{n_\rho, K \in N_{sh}} |C_{n_\rho, K}|^2, \quad N_{sh} = 0, 1, 2, \dots$$

Inter-cluster distances R_1 and R_2 

$$R_1^2 = \int d\mathbf{x}d\mathbf{y} |f(\mathbf{x}, \mathbf{y})|^2 \cdot \mathbf{y}^2$$

$$R_2^2 = \int d\mathbf{x}d\mathbf{y} |f(\mathbf{x}, \mathbf{y})|^2 \cdot \mathbf{x}^2$$

Normalization condition:

$$\langle \Psi_{E,J} | \Psi_{\tilde{E},J} \rangle = \sum_{n_\rho, c} C_{n_\rho, c}^{E,J} C_{n_\rho, c}^{\tilde{E},J} = \delta(E - \tilde{E}).$$

Theoretical set-up

Input parameters

NN potential

Minnesota potential, modified Hasegawa-Nagata potential

Basis

Hypermomentum: $K_{\max} = 14$ for even parity states

Hypermomentum: $K_{\max} = 13$ for odd parity states

Hyperradial excitations: $n_\rho \leq 100$ ($\sim 200 \hbar\Omega$)

Adjustable parameters

Oscillator length b : $E_{th}(b) = \min\{E(A_1) + E(A_2) + E(A_3)\}$

Majorana parameter m of NN potential: $E_{GS}(m) = E_{GS}^{(Exp.)}$

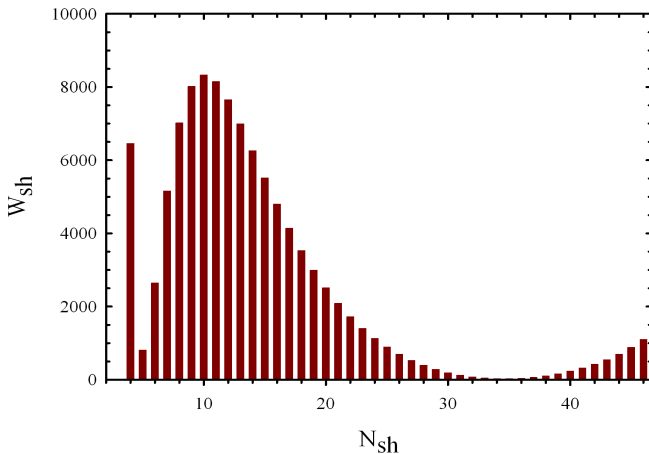
Hoyle state and other resonance states in ^{12}C .

J^π	CSM [1]		AMHHB [2]	
	E , MeV	Γ , keV	E , MeV	Γ , keV
0^+	0.76	2.4	0.68	2.9
	1.66	1480	5.16	534
2^+	2.28	1100	2.78	10
	5.14	1900	3.17	280
	6.82	240	5.60	0.6
1^-	3.65	0.30	3.52	0.21

1 C. Kurokawa, K. Katō. *Nucl. Phys. A*, **792**, 87, 2007.

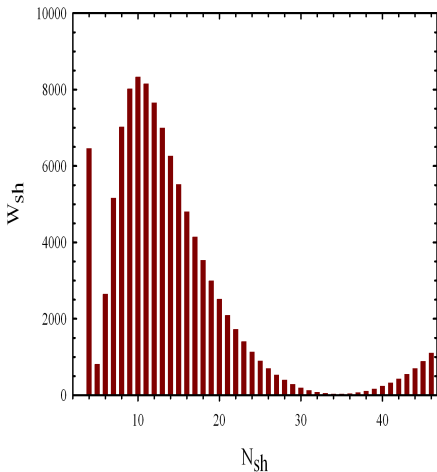
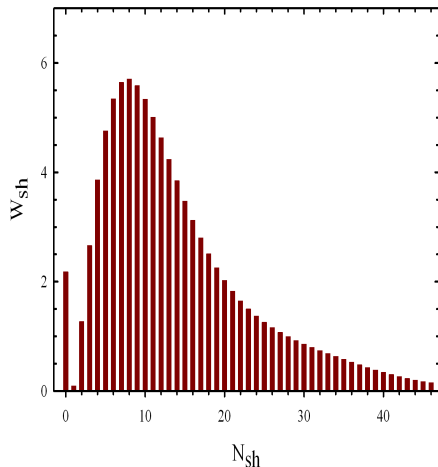
2 V. Vasilevsky, et al. *Phys. Rev. C*, **85**, 034318, 2012.

Wave function of the Hoyle state ($E = 0.39$ MeV, $\Gamma = 1$ eV) in ^{12}C .



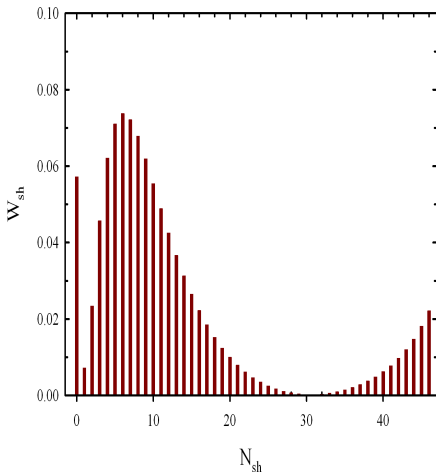
Similar results:

- 1 C. Kurokawa, K. Katō. *Nucl. Phys. A*, **792**, 87, 2007.
- 2 T. Neff, H. Feldmeier. *Few-Body Syst.*, **45**, 145, 2009.

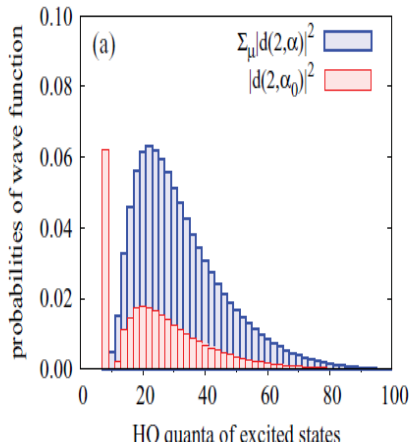
Wave functions of the 0^+ resonance states in ^{12}C First resonance, $\Gamma=1$ eVSecond resonance, $\Gamma=692$ keV

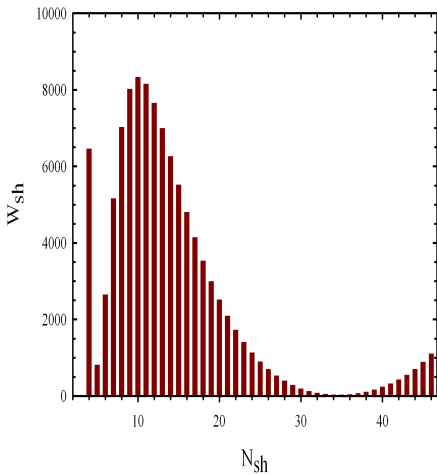
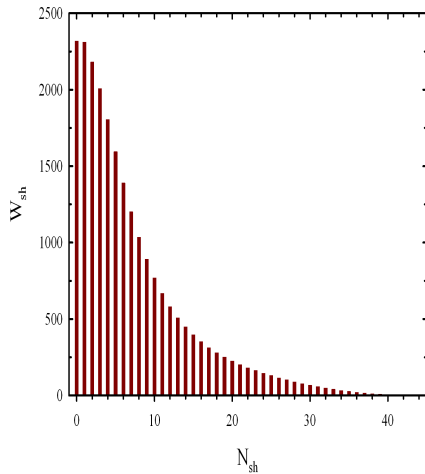
Wave functions of the 0^+ resonance state in ^{12}C

AMHHB

CSM: C. Kurokawa, K. Katō. *Nucl. Phys. A*, **792**, 87, 2007.

CSM



Wave functions of the 0^+ and 1^- resonance states in ^{12}C 0^+ state, $\Gamma=1$ eV 1^- state, $\Gamma=979$ eV

Candidates for the Hoyle state

The Hoyle state

Criteria

- 1 Resonance state which is close to three-cluster threshold (E is small)
- 2 Resonance state with very small width (Γ is small)
- 3 Resonance state with zero values of the total orbital momentum ($L=0$)
- 4 Large values of W_{sh} in the internal region $0 \leq N_{sh} \leq 45$

Candidates for the Hoyle state in ^{11}B

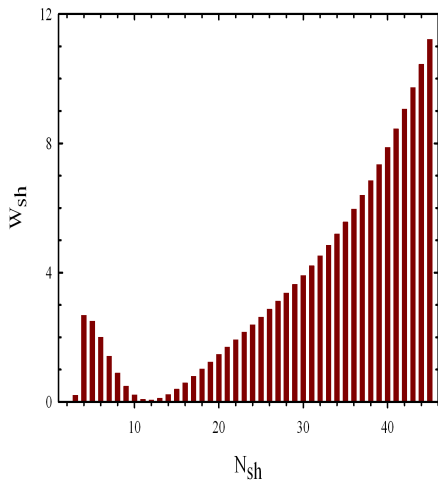
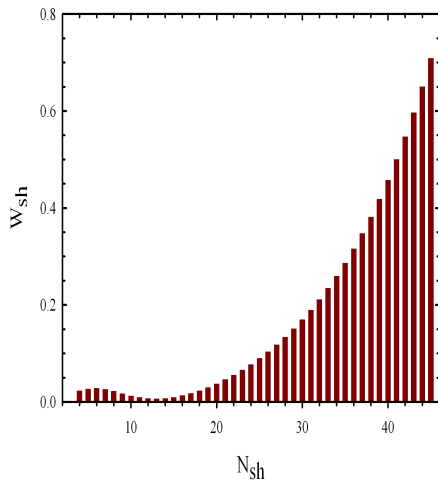
J^π	E , MeV	Γ , keV	J^π	E , MeV	Γ , keV
$\frac{3}{2}^-$	0.755	0.58	$\frac{1}{2}^+$	0.437	15.26
	1.402	185.18		0.702	12.30
	1.756	143.72		1.597	15.95
$\frac{1}{2}^-$	1.436	374.64	$\frac{3}{2}^+$	1.147	1.50
	1.895	100.95		1.367	8.58
	2.404	450.07		1.715	41.24
$\frac{5}{2}^-$	0.583	$5.14 \cdot 10^{-4}$	$\frac{5}{2}^+$	1.047	1.54
	1.990	32.63		1.951	40.20
	2.251	138.87		2.265	54.73
	2.905	120.46		2.748	167.61

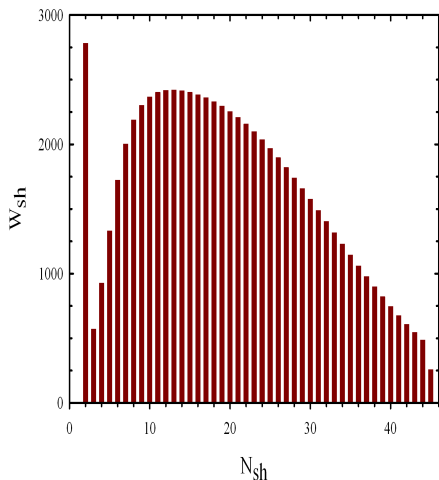
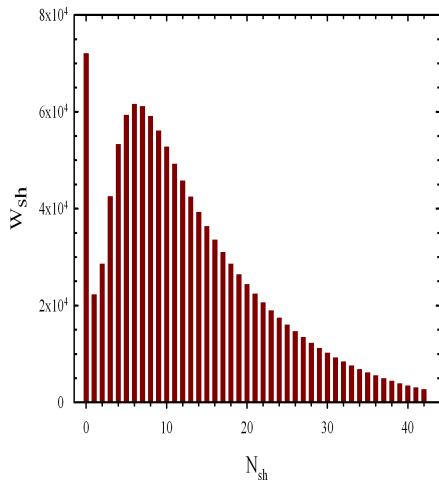
V. S. Vasilevsky. "Microscopic three-cluster description of ^{11}B and ^{11}C nuclei", Ukr. J Phys., 58, pp. 544-553, 2013.

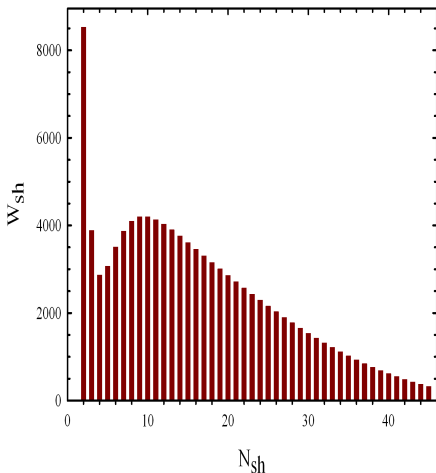
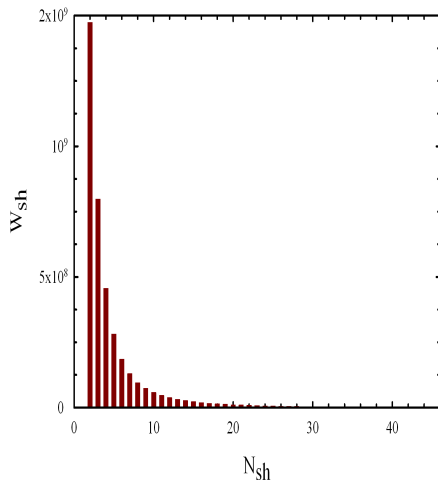
Candidates for the Hoyle state in ^{11}C

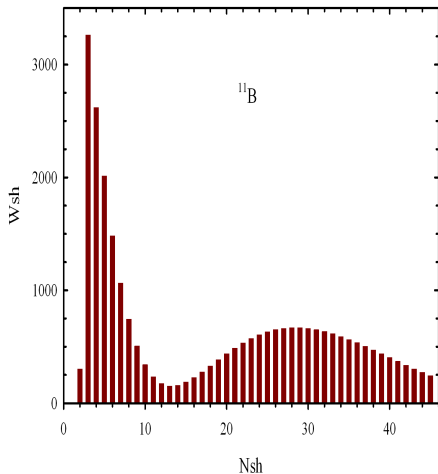
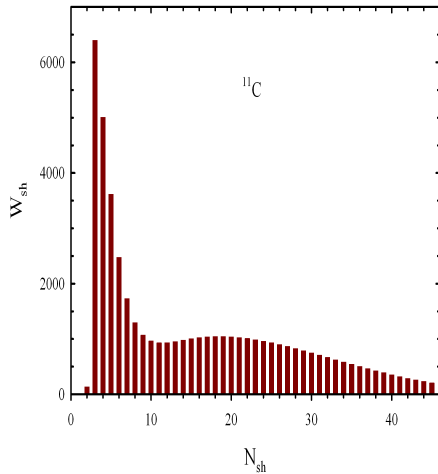
J^π	E , MeV	Γ , keV	J^π	E , MeV	Γ , keV
$\frac{3}{2}^-$	0.805	$9.93 \cdot 10^{-3}$	$\frac{1}{2}^+$	0.906	162.94
	1.920	105.08		1.930	59.88
	2.324	619.76		2.679	86.69
$\frac{1}{2}^-$	1.142	0.71	$\frac{3}{2}^+$	2.268	34.25
	2.266	790.98		2.478	159.28
	3.014	366.15		2.850	115.19
$\frac{5}{2}^-$	0.783	$9.65 \cdot 10^{-5}$	$\frac{5}{2}^+$	1.460	0.90
	1.897	5.77		2.346	82.72
	3.026	182.69		3.179	122.75

V. S. Vasilevsky. "Microscopic three-cluster description of ^{11}B and ^{11}C nuclei", Ukr. J Phys., 58, pp. 544-553, 2013.

Wave function of $1/2^+$ state ^{11}B , $E=0.44$ MeV, $\Gamma=15$ keV ^{11}C , $E=0.91$ MeV, $\Gamma=163$ keV

Wave function of $3/2^-$ state ^{11}B , $E=0.76$ MeV, $\Gamma=580$ eV ^{11}C , $E=0.81$ MeV, $\Gamma=9$ eV

Wave function of $5/2^-$ states ^{11}B , $E=0.58$ MeV, $\Gamma=0.5$ eV ^{11}C , $E=0.78$ MeV, $\Gamma=0.1$ eV

Wave function of $5/2^+$ states ^{11}B , $E=1.05$ MeV, $\Gamma=1.5$ keV ^{11}C , $E=1.46$ MeV, $\Gamma=0.9$ keV

OCM+CSM versus AMHHB: ^{11}B

J^π	OCM+CSM [1,2]			AMHHB		
	E , MeV	Γ , keV	HAS	E , MeV	Γ , keV	HAS
$1/2^+$	0.75	190	Y	0.44	15	N
$3/2_3^-$	-2.90	-	-	-0.59	-	-
$3/2_4^-$				0.76	0.58	Y

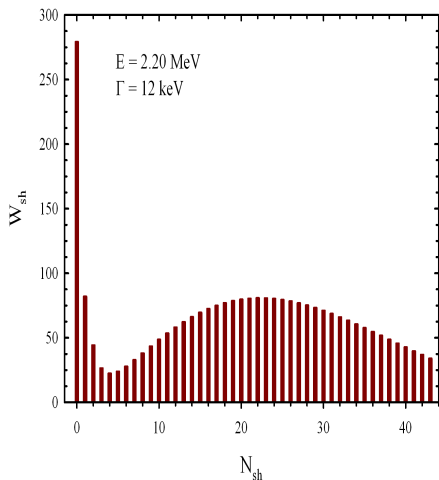
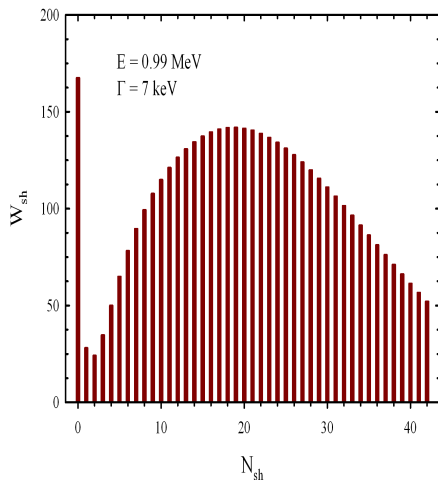
- 1 Yamada, T. and Funaki, Y. " $\alpha + \alpha + t$ cluster structures and $^{12}\text{C}(0_2^+)$ -analog states in ^{11}B ", *Phys. Rev. C*, **82**, 064315, 2010.
- 2 Yamada, T. and Funaki, Y. "Three-Body Cluster Structures and Hoyle-Analogue States in ^{11}B ", *Progr. Theor. Phys. Suppl.*, **196**, 388, 2012.

Candidates for the Hoyle state in ^{10}B

J^π	E , MeV	Γ , keV	Γ/E	R_1 , fm	R_2 , fm
1^+	0.604	232.30	0.384	6.67	10.67
	0.987	7.08	7.17×10^{-3}		
	1.536	196.36	0.128		
2^+	1.055	12.06	11.43×10^{-3}	6.64	10.83
	2.810	170.74	0.061		
3^+	1.062	11.73	11.05×10^{-3}	6.43	10.35
	2.202	526.47	0.239		
1^-	1.100	76.75	0.070	9.31	10.84
	1.820	562.71	0.309		

A. V. Nesterov, V. S. Vasilevsky, T. P. Kovalenko. "Spectrum of bound states of nucleus ^{10}B in a three-cluster microscopic model."

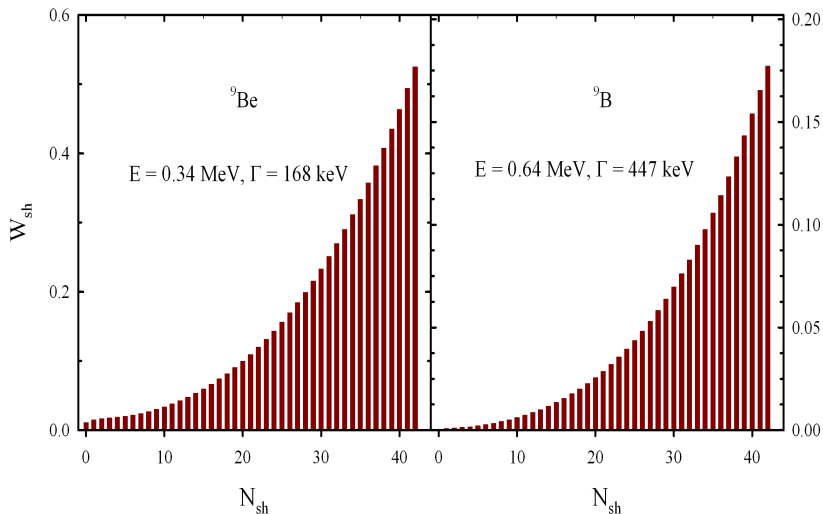
Ukr. J. Phys., **59**, 1065, 2014.

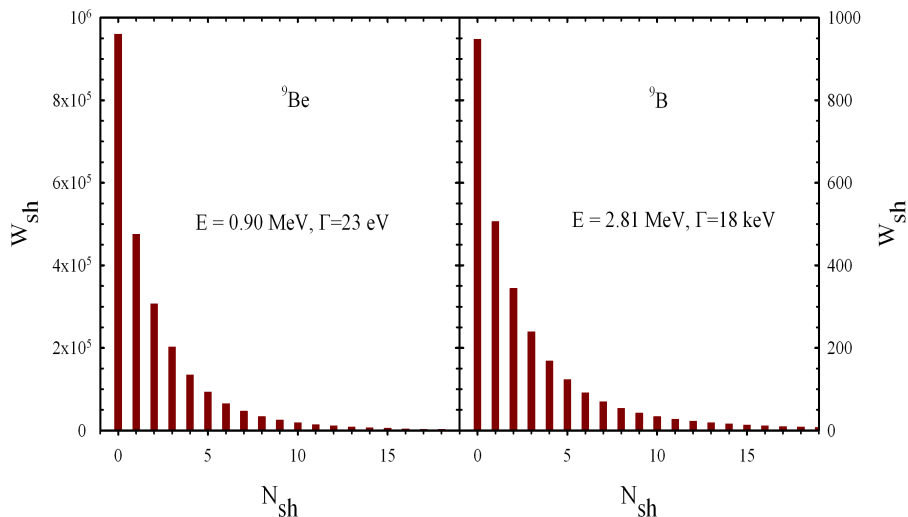
Wave function of 3^+ and 1^+ resonance states in ^{10}B . $^{10}\text{B}, J^\pi = 3^+$  $^{10}\text{B}, J^\pi = 1^+$ 

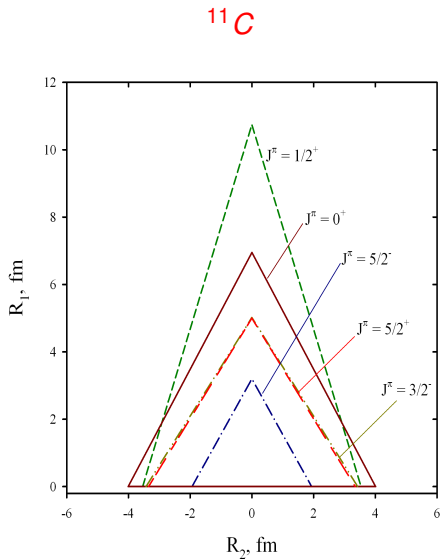
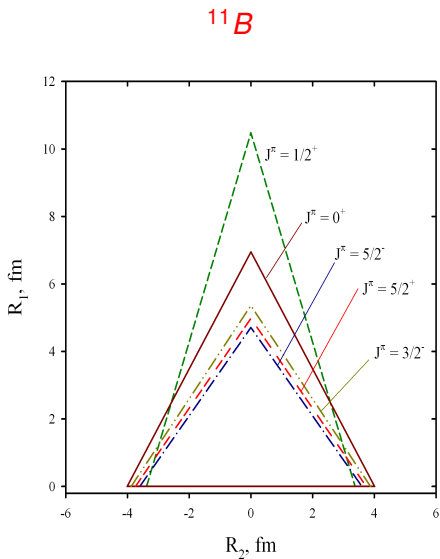
Candidates for the Hoyle state in ${}^9\text{Be}$ and ${}^9\text{B}$

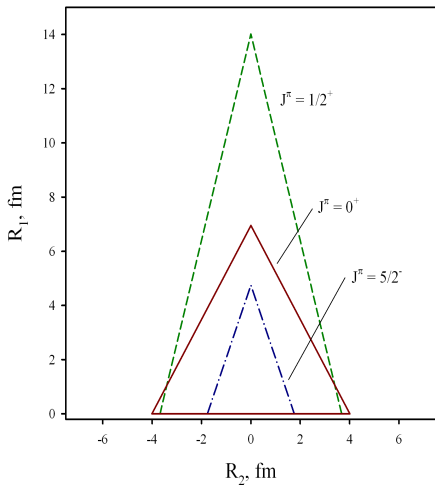
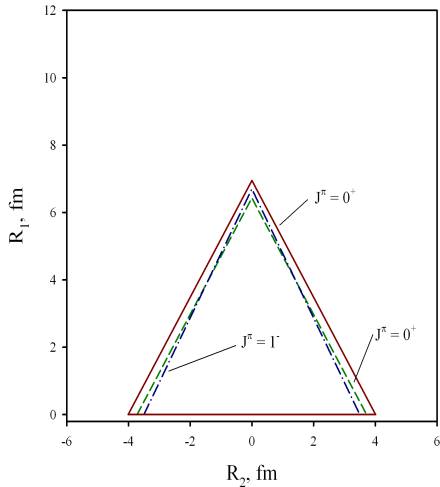
${}^9\text{Be}$			${}^9\text{B}$		
J^π	E , MeV	Γ , MeV	J^π	E , MeV	Γ , MeV
$3/2^-$	-1.574	-	$3/2^-$	0.379	1.076×10^{-6}
$1/2^+$	0.338	0.168	$1/2^+$	0.636	0.477
$5/2^-$	0.897	2.36×10^{-5}	$5/2^-$	2.805	0.018
$5/2^+$	2.086	0.112	$3/2^+$	2.338	2.796
$3/2_2^-$	2.704	2.534	$1/2^-$	3.398	3.428
$1/2^-$	2.866	1.597	$5/2^+$	3.670	0.415
$3/2^+$	4.062	1.224	$3/2_2^-$	3.420	3.361
$7/2^-$	4.766	4.041	$5/2_2^-$	5.697	5.146

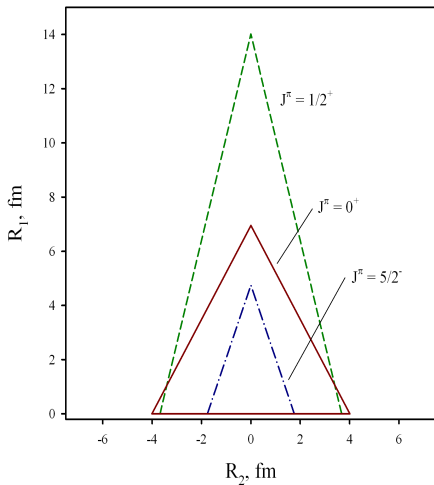
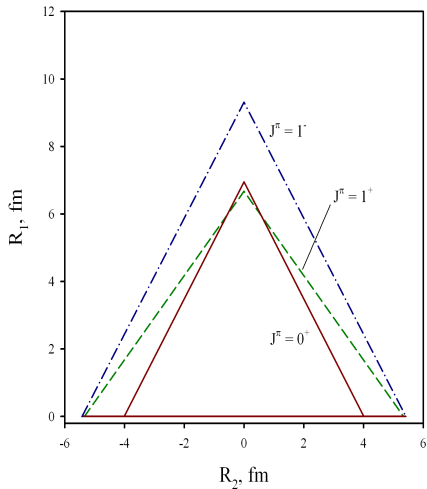
V. S. Vasilevsky, K. Katō, N. Takibayev. "Formation and decay of resonance states in ${}^9\text{Be}$ and ${}^9\text{B}$ nuclei. Microscopic three-cluster model investigations", Phys. Rev. C, **96**, 034322, 2017.

Wave function of the $1/2^+$ resonance states in ${}^9\text{Be}$ and ${}^9\text{B}$ 

Wave function of the $5/2^-$ resonance state in ${}^9\text{Be}$ and ${}^9\text{B}$.

Shape of triangles. Resonance states in ^{11}B and ^{11}C .

Shape of triangles. Resonance states in ${}^9\text{Be}$ and ${}^{12}\text{C}$. ${}^9\text{Be}$  ${}^{12}\text{C}$ 

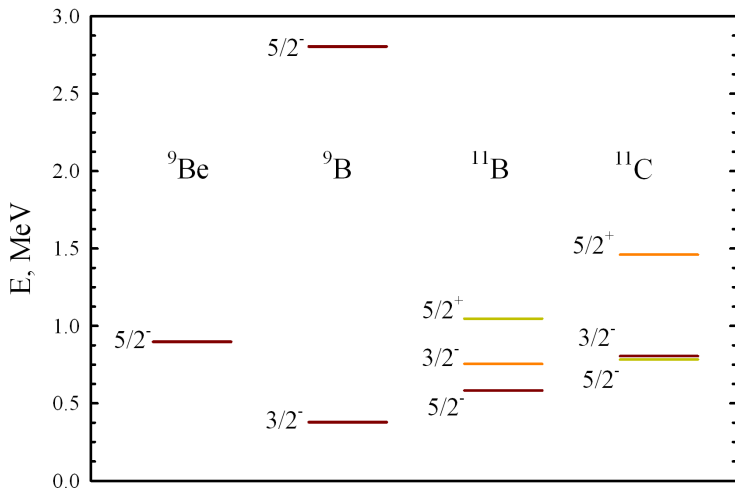
Shape of triangles. Resonance states in ${}^9\text{Be}$ and ${}^{10}\text{B}$. ${}^9\text{Be}$  ${}^{10}\text{B}$ 

Energy and width of the Hoyle-analog states

Nucleus	Partition	J^π	E , MeV	Γ , keV	Γ/E
${}^9\text{Be}$	$\alpha + \alpha + n$	$5/2^-$	0.897	$2.36 \cdot 10^{-2}$	$2.63 \cdot 10^{-5}$
${}^9\text{B}$	$\alpha + \alpha + n$	$3/2^-$	0.379	$1.08 \cdot 10^{-3}$	$2.84 \cdot 10^{-6}$
		$5/2^-$	2.805	$18.0 \cdot 10^{-3}$	$6.42 \cdot 10^{-6}$
${}^{11}\text{B}$	$\alpha + \alpha + {}^3\text{H}$	$5/2^-$	0.583	$5.14 \cdot 10^{-4}$	$8.87 \cdot 10^{-7}$
		$3/2^-$	0.755	0.58	7.70×10^{-4}
		$5/2^+$	1.047	1.54	1.47×10^{-3}
${}^{11}\text{C}$	$\alpha + \alpha + {}^3\text{He}$	$3/2^-$	0.805	$9.93 \cdot 10^{-3}$	$1.23 \cdot 10^{-5}$
		$5/2^-$	0.783	$9.64 \cdot 10^{-5}$	$1.23 \cdot 10^{-7}$
		$5/2^+$	1.460	0.90	6.16×10^{-4}
${}^{12}\text{C}$	$\alpha + \alpha + \alpha$	0^+	0.395	$1.14 \cdot 10^{-3}$	$2.88 \cdot 10^{-6}$

New criterion for the Hoyle-analog states: $\Gamma/E < 2 \cdot 10^{-3}$

Energy and width of the Hoyle-analog states



Summary

- Hoyle-analog states in light nuclei:
 - $5/2^-$ state in ${}^9\text{Be}$
 - $3/2^-$ $5/2^-$ states in ${}^9\text{B}$
 - $3/2^-$, $5/2^-$ and $5/2^+$ states in ${}^{11}\text{B}$
 - $3/2^-$, $5/2^-$ and $5/2^+$ states in ${}^{11}\text{C}$
- We did not find the Hoyle-analog state in ${}^{10}\text{B}$
- Hoyle-analog state is a very narrow resonance state with a fairly compact three-cluster configuration
- Synthesis of light nuclei ${}^9\text{Be}$, ${}^{11}\text{B}$ and ${}^{11}\text{C}$ is possible in a triple collision of $\alpha + \alpha + n$, $\alpha + \alpha + {}^3\text{H}$ and $\alpha + \alpha + {}^3\text{He}$, respectively

THANK YOU VERY MUCH!