

Heavy-Quark Interactions in the Quark-Gluon Plasma and Single-Electron Spectra at RHIC

Hendrik van Hees and Ralf Rapp

Heavy quarks and quarkonia are believed to be valuable probes of the medium produced in ultrarelativistic heavy-ion collisions (URHICs). Recent spectra of semileptonic decay electrons associated with charm and bottom hadrons in Au-Au collisions at the Relativistic Heavy-Ion Collider (RHIC) have shown a surprisingly large suppression (small nuclear modification factor, R_{AA}) [1-3] and elliptic flow (v_2) [4-6], indicating significantly stronger interactions of heavy quarks in the quark-gluon plasma (QGP) than expected within perturbative QCD (pQCD).

To study microscopic reaction mechanisms underlying the behavior of heavy quarks in a strongly interacting QGP (sQGP), we have introduced D - and B -meson like resonance states [7] mediating elastic rescattering of c - and b -quarks. Here, we study the consequences of this conjecture for the thermalization and flow of the heavy quarks in the sQGP formed in Au-Au collisions at RHIC, employing a Fokker-Planck approach, followed by a combined quark-coalescence/fragmentation model for the hadronization of the heavy quarks. The resulting transverse-momentum (p_T) spectra and v_2 of the single decay-electrons are confronted with RHIC data [8].

The drag and diffusion coefficients for heavy-quarks, entering the Fokker-Planck equation, are calculated from elastic rescattering off light antiquarks in the sQGP via D - and B -meson resonances assumed to survive above the critical temperature. This is motivated by recent lattice QCD (lQCD) computations of hadronic correlators and lQCD-based potential models which indicate colorless resonances in both the light- and heavy-quark sector [9]. The resonant $Q\bar{q}$ cross sections are supplemented with leading-order pQCD elastic scattering [10]. We find that the resulting drag coefficients imply thermalization times which are lower by a factor ~ 3 compared to pQCD scattering alone [7].

The coefficients are used in a relativistic Langevin simulation for heavy-quark interactions in an isentropically expanding, elliptic QGP fireball corresponding to impact parameter $b=7\text{fm}$ 200 AGeV Au-Au collisions at RHIC. The expansion parameters are adjusted to resemble the time evolution of radial and elliptic flow of the bulk matter in hydrodynamic simulations [11]. The proper thermal equilibrium limit in the Langevin process is implemented via the Hänggi-Klimontovich realization [12], cf. also Ref. [13].

The initial heavy-quark p_T -distributions and the relative magnitude of c - and b -quark spectra are determined by fitting experimental D - and D^* - spectra in d-Au collisions [14] and attributing the missing yield of the corresponding semileptonic electrons [15] at higher p_T to B -meson decays. This leads to a cross-section ratio for b - and c -quark pair production of $4.9 \cdot 10^{-3}$ and a crossing of D - and B -decay electrons at $p_T \sim 5$ GeV, in line with expectations from pQCD.

The c - and b -quark output spectra from the Langevin simulation are subjected to coalescence with light antiquarks following the model of Ref. [16]. Conservation of c - and b -quark number is ensured by hadronizing unpaired heavy quarks via δ -function fragmentation. Finally, the single-electron spectra are

obtained from D - and B -meson three-body decays. Fig. 1 shows that resonance scattering leads to a substantial decrease in R_{AA} and increase in v_2 compared to elastic pQCD rescattering. While coalescence with light antiquarks amplifies v_2 , it also leads to *harder* D - and B -meson momentum spectra, which increases the electron R_{AA} . The approach to thermalization of c -quarks leads to a strong quenching of their electron-decay spectra, entailing that B -meson decay contributions become prevalent for electron momenta of $p_T > 2.5$ GeV. This substantially reduces the effects in the single-electron spectra, since b -quarks (due to their large mass) are much less affected in the sQGP than c -quarks.

In conclusion, resonance formation in the sQGP could play an essential role in understanding current observations on semileptonic single-electron spectra at RHIC. Future calculations should aim at a consistent inclusion of radiative energy-loss which is expected to become the dominant effect at (very) high p_T .

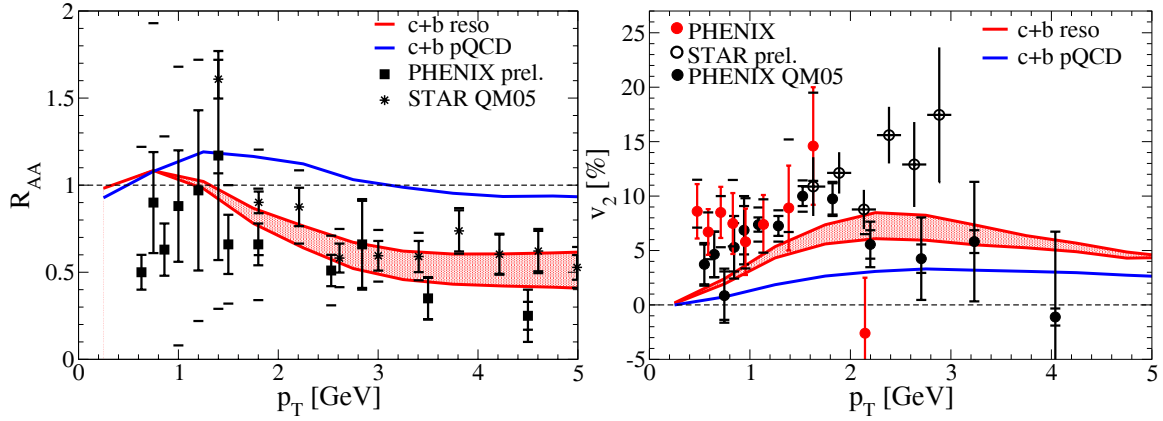


Figure 1. Nuclear Modification factor, R_{AA} (left panel), and elliptic flow, v_2 (right panel), of semileptonic D - and B -meson decay electrons in $b=7$ fm 200 AGeV Au-Au collisions assuming different elastic heavy-quark interactions in the QGP with subsequent hadronization via coalescence and fragmentation, compared to PHENIX and STAR data [1,2,5,6].

- [1] S. S. Adler *et al.* (PHENIX Collaboration), Phys. Rev. Lett. **94**, 082301 (2005).
- [2] B. Jacak *et al.* (PHENIX Collaboration), nucl-ex/0508036.
- [3] J. Bielcik *et al.* (STAR Collaboration), nucl-ex/0511005.
- [4] S. S. Adler *et al.* (PHENIX Collaboration), Phys. Rev. C **72**, 024901 (2005).
- [5] F. Laue *et al.* (STAR Collaboration), J. Phys. G **31**, S27 (2005).
- [6] Y. Akiba *et al.* (PHENIX Collaboration), nucl-ex/0510008.
- [7] H. van Hees and R. Rapp, Phys. Rev. C **71**, 034907 (2005).
- [8] H. van Hees, V. Greco and R. Rapp, Phys. Rev. C. **73**, 034913 (2006).
- [9] M. Asakawa and T. Hatsuda, Phys. Rev. Lett. **92**, 012001 (2004); F. Karsch and E. Laermann, hep-lat/0305025; R. Morrin *et al.*, PoS LAT2005 (2005) 176; E.V. Shuryak and I. Zahed, Phys. Rev. C **70**, 021901(R) (2004); C. Y. Wong, Phys. Rev. C **72**, 034906 (2005); M. Mannarelli and R. Rapp, Phys. Rev. C **72**, 064905 (2005); Á. Mócsy and P. Petretczky, Phys. Rev. D **73**, 074007 (2006).

- [10] B.L. Combridge, Nucl. Phys. **B151**, 429 (1979).
- [11] P.F. Kolb, J. Sollfrank, and U. Heinz, Phys. Rev. C **62**, 054909 (2000).
- [12] J. Dunkel and P. Hänggi, Phys. Rev. E **71**, 016124 (2005).
- [13] G.D. Moore and D. Teaney, Phys. Rev. C **71**, 064904 (2005).
- [14] J. Adams *et al.* (STAR Collaboration), Phys. Rev. Lett. **94**, 062301 (2005); A. Tai *et al.* (STAR Collaboration), J. Phys. G **30**, S809 (2004).
- [15] A.A. P. Suaide *et al.* (STAR Collaboration), J. Phys. G **30**, S1179 (2004).
- [16] V. Greco, C. M. Ko, and R. Rapp, Phys. Lett. **B595**, 202 (2004).