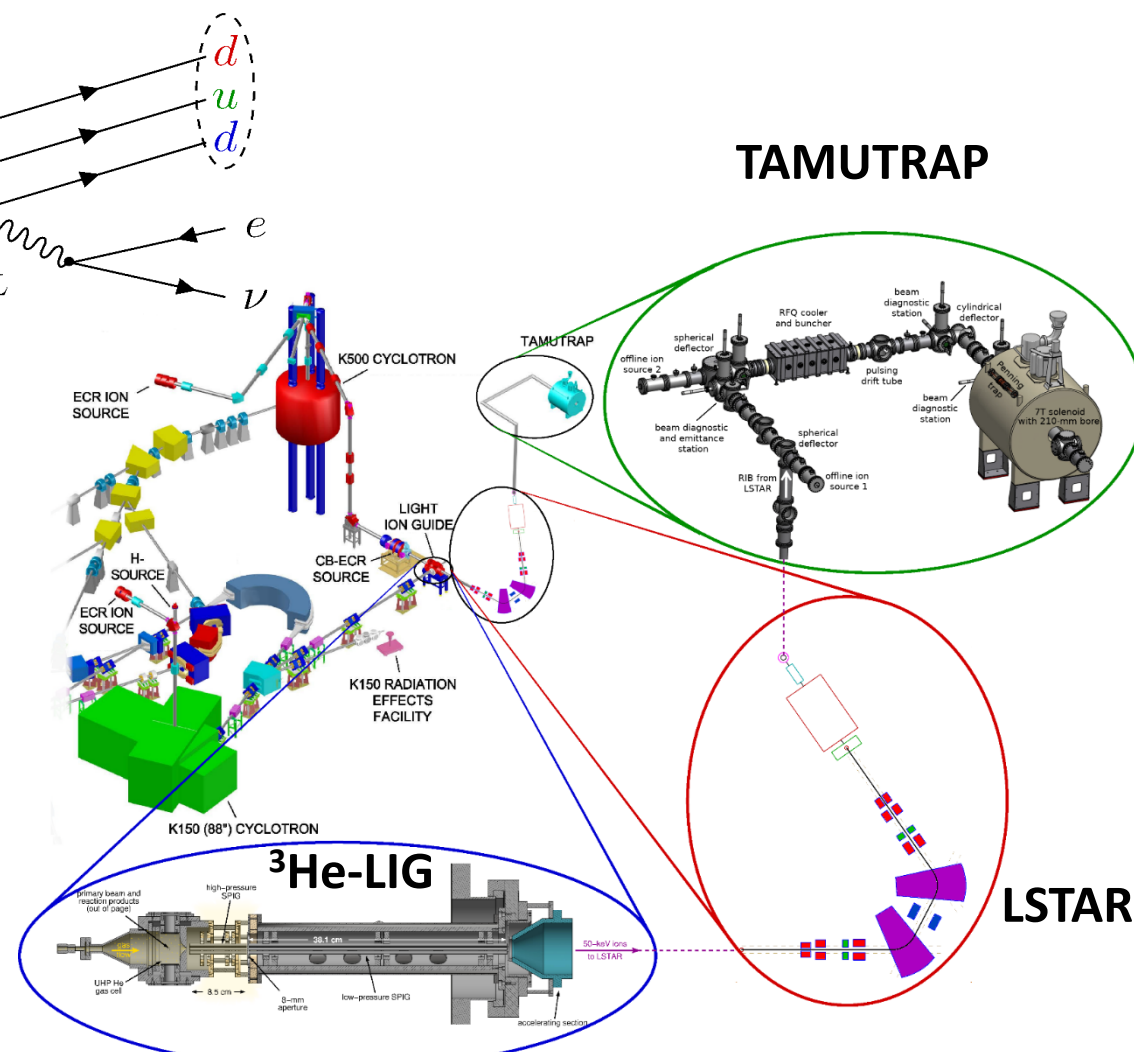
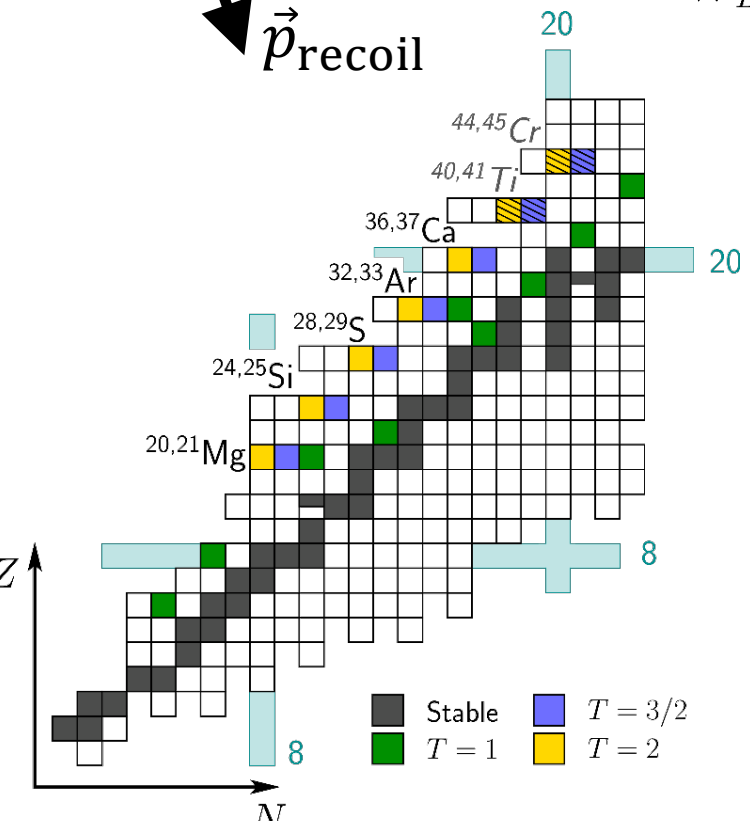
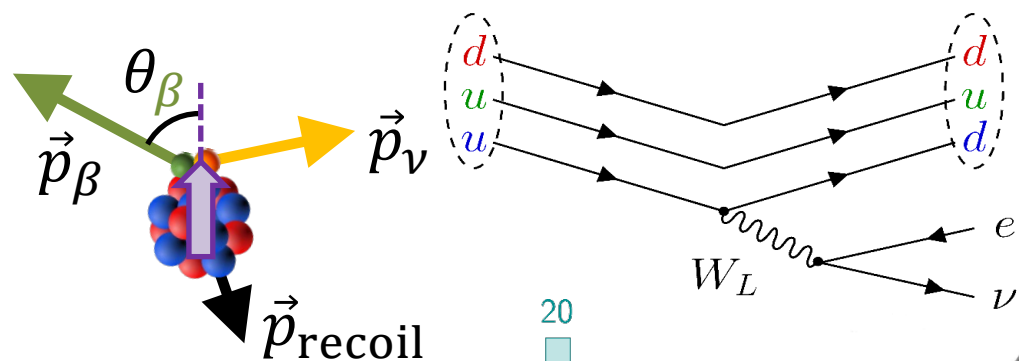




Expanding RIB Capabilities at the Cyclotron Institute: ^3He -LIG production with an Isobar Separator LSTAR




Outline




Motivation

-  Testing the standard model via the precision frontier
-  TAMUTRAP: Penning trap for β -delayed proton decay studies

^3He -driven light-ion guide

-  Initial tests with a prototype gas cell

Transporting the beam to TAMUTRAP

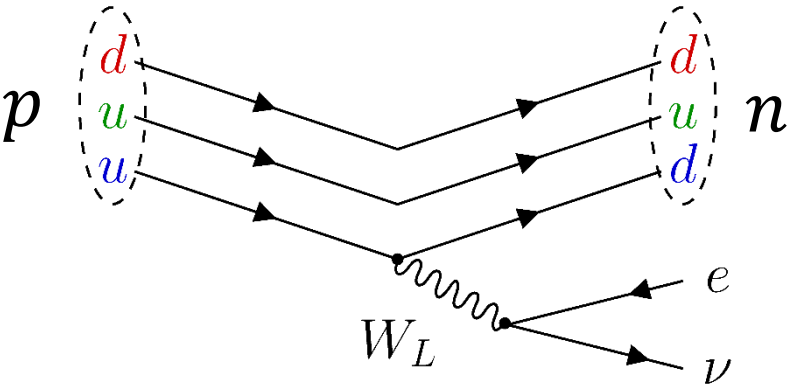
-  Design of the **L**ight-ion guide **S**eparator for **TAMU**'s K150 **RIBs** (LSTAR)
-  Expected performance of LSTAR
-  Looking forward

The standard model and beyond



This is the standard model:

pure $V - A$ interaction



$$H_\beta = \bar{p}\gamma_\mu n(C_V\bar{e}\gamma^\mu\nu + C'_V\bar{e}\gamma^\mu\gamma_5\nu) - \bar{p}\gamma_\mu\gamma_5 n(C_A\bar{e}\gamma^\mu\gamma_5\nu + C'_A\bar{e}\gamma^\mu\nu)$$

$$C_V = C'_V = 1$$

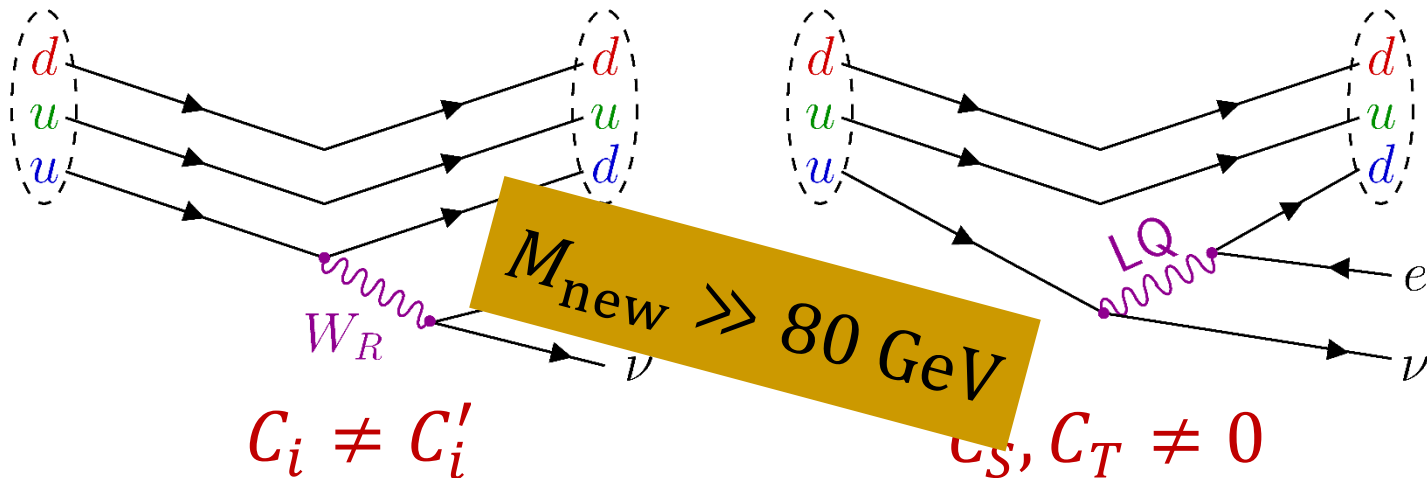
$$C_A = C'_A \approx 1.27$$

$$M_W = 80.385 \text{ GeV}$$



These are not:

Right-handed bosons, or scalar/tensor leptoquarks, or SUSY, or...



$$C_i \neq C'_i$$

$$C_S, C_T \neq 0$$

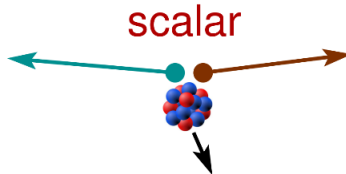
β decay and fundamental physics

The often-quoted angular distribution of β decay
(Jackson, Treiman and Wyld, Phys Rev **106** and Nucl Phys **4**, 1957)

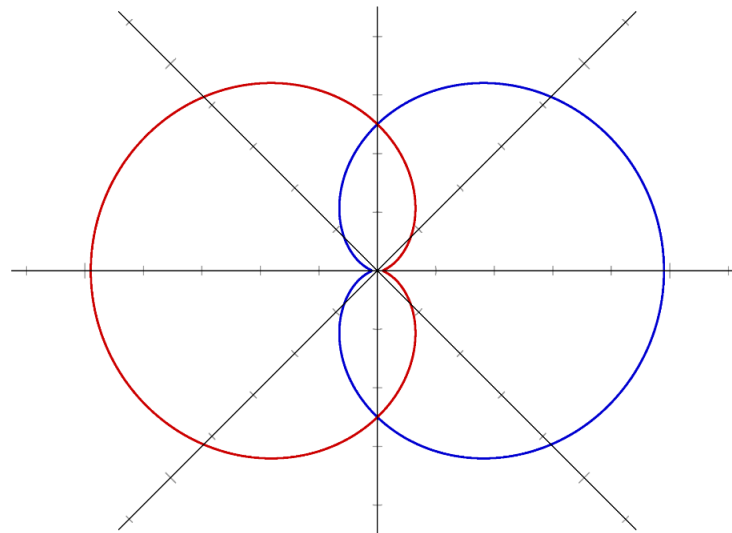
$$\frac{d^5W}{dE_e d\Omega_e d\Omega_{\nu_e}} = \overbrace{\frac{G_F^2 |V_{ud}|^2}{(2\pi)^5} p_e E_e (A_0 - E_e)^2}^{\text{basic decay rate}} \xi \left(1 + \overbrace{a_{\beta\nu} \frac{\vec{p}_e \cdot \vec{p}_{\nu_e}}{E_e E_{\nu_e}}}^{\beta-\nu \text{ correlation}} + \overbrace{b \frac{\Gamma m_e}{E_e}}^{\text{Fierz term}} \right)$$

$$a_{\beta\nu} = \frac{|C_V|^2 + |C'_V|^2 - |C_S|^2 - |C'_S|^2}{|C_V|^2 + |C'_V|^2 + |C_S|^2 + |C'_S|^2} = 1??$$

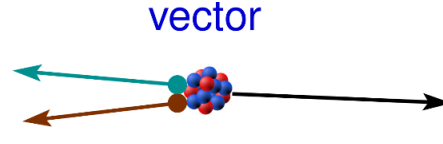
scalar



$$a_{\beta\nu} = \frac{-|C_S|^2 - |C'_S|^2}{|C_S|^2 + |C'_S|^2}$$



vector



$$a_{\beta\nu} = \frac{|C_V|^2 + |C'_V|^2}{|C_V|^2 + |C'_V|^2}$$

β decay and fundamental physics

- The often-quoted angular distribution of β decay (Jackson, Treiman and Wyld, Phys Rev **106** and Nucl Phys **4**, 1957)

$$\frac{d^5W}{dE_e d\Omega_e d\Omega_{\nu_e}} = \overbrace{\frac{G_F^2 |\mathbf{V}_{ud}|^2}{(2\pi)^5} p_e E_e (A_0 - E_e)^2}^{\text{basic decay rate}} \xi \left(1 + \overbrace{a_{\beta\nu} \frac{\vec{p}_e \cdot \vec{p}_{\nu_e}}{E_e E_{\nu_e}}}^{\beta-\nu \text{ correlation}} + \overbrace{b \frac{\Gamma m_e}{E_e}}^{\text{Fierz term}} \right)$$

$$a_{\beta\nu} = \frac{|C_V|^2 + |C'_V|^2 - |C_S|^2 - |C'_S|^2}{|C_V|^2 + |C'_V|^2 + |C_S|^2 + |C'_S|^2} = 1??$$

- Not as sensitive as the Fierz parameter, which is linear in the new couplings:

$$b = \frac{-2\Re(C_S^* C_V + C'_S{}^* C'_V)}{|C_V|^2 + |C'_V|^2 + |C_S|^2 + |C'_S|^2} = 0??$$

(see Falkowski, González-Alonso and Naviliat-Čunčić, JHEP **4**, 126 (2021))

β - ν correlation via β -delayed p decay

VOLUME 83, NUMBER 7

PHYSICAL REVIEW LETTERS

16 AUGUST 1999

Positron-Neutrino Correlation in the $0^+ \rightarrow 0^+$ Decay of ^{32}Ar

E. G. Adelberger,¹ C. Ortiz,² A. García,² H. E. Swanson,¹ M. Beck,¹ O. Tengblad,³ M. J. G. Borge,³ I. Martel,⁴
H. Bichsel,¹ and the ISOLDE Collaboration⁴

¹Department of Physics, University of Washington, Seattle, Washington 98195-1560

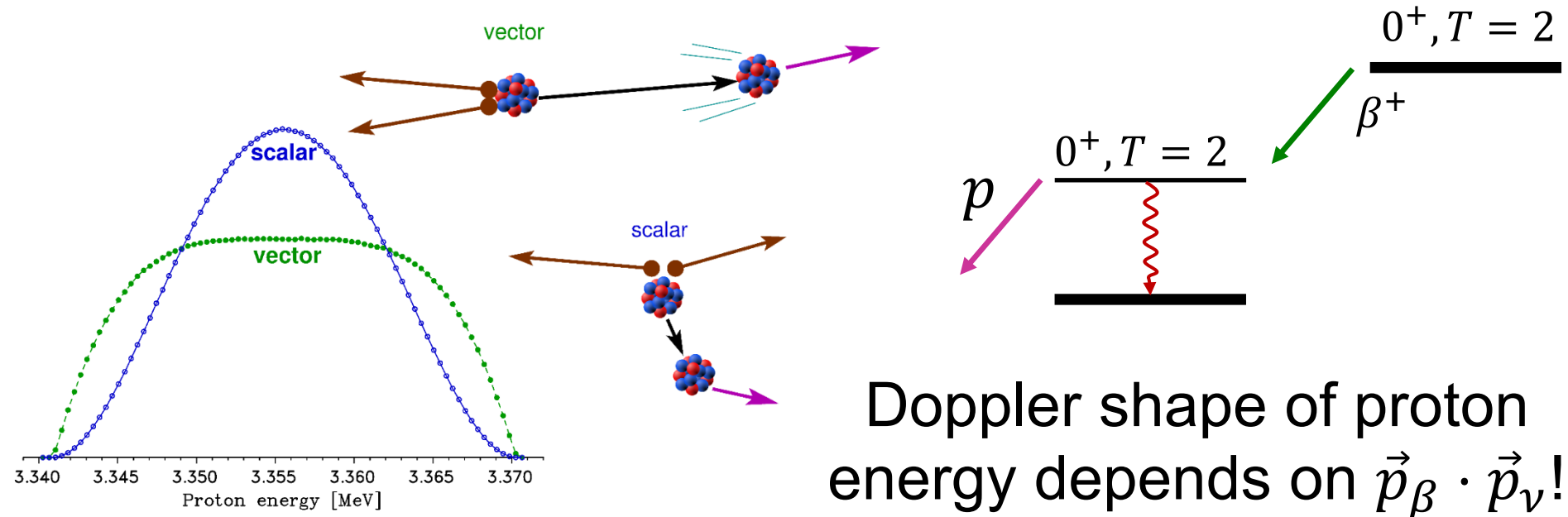
²Department of Physics, University of Notre Dame, Notre Dame, Indiana 46556

³Instituto de Estructura de la Materia, CSIC, E-28006 Madrid, Spain

⁴EP Division, CERN, Geneva, Switzerland CH-1211

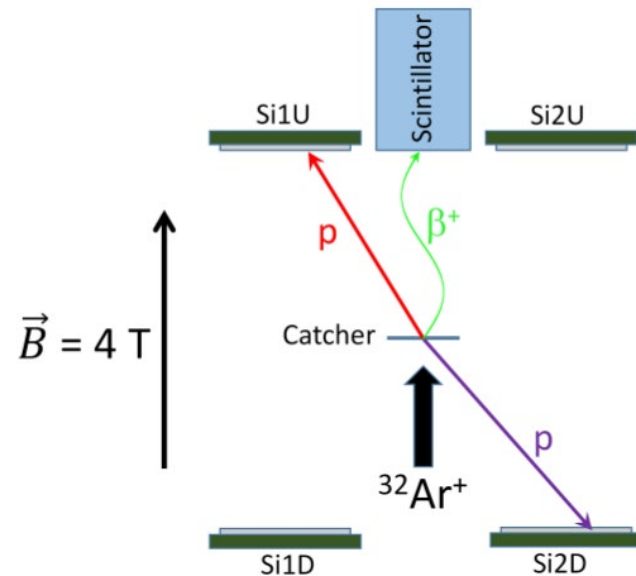
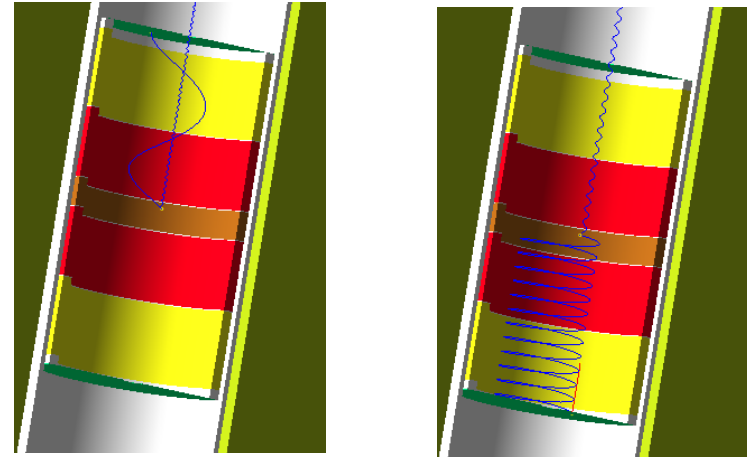
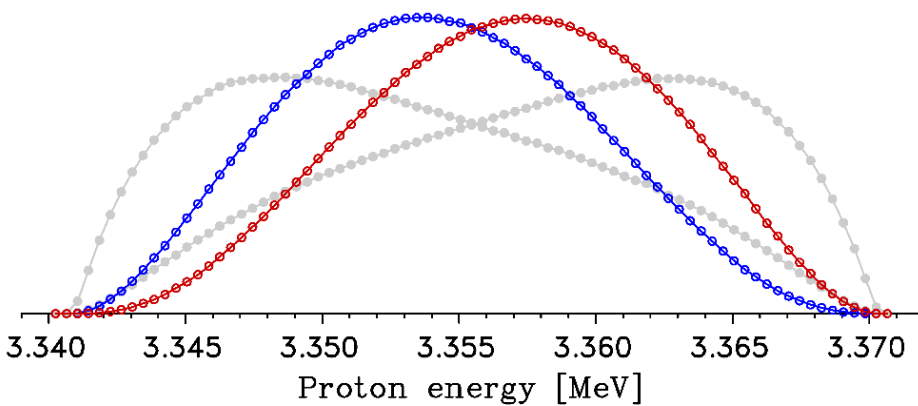
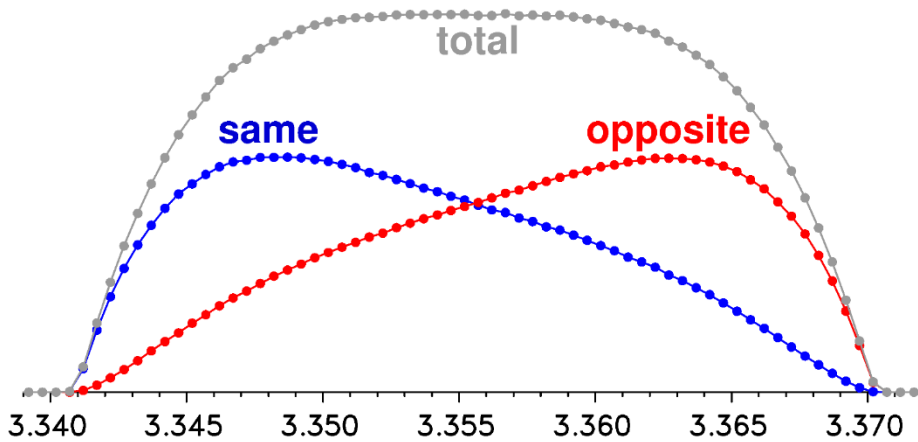
(Received 24 February 1999)

The positron-neutrino correlation in the $0^+ \rightarrow 0^+$ β decay of ^{32}Ar was measured at ISOLDE by analyzing the effect of lepton recoil on the shape of the narrow proton group following the superallowed decay. Our result is consistent with the standard model prediction. For vanishing Fierz interference we find $a = 0.9989 \pm 0.0052 \pm 0.0039$, which yields improved constraints on scalar weak interactions.



Doppler shape of proton
energy depends on $\vec{p}_\beta \cdot \vec{p}_\nu$!

Measure means instead of 2nd moments



WISArD collab, PRC 101, 055501 (2020)

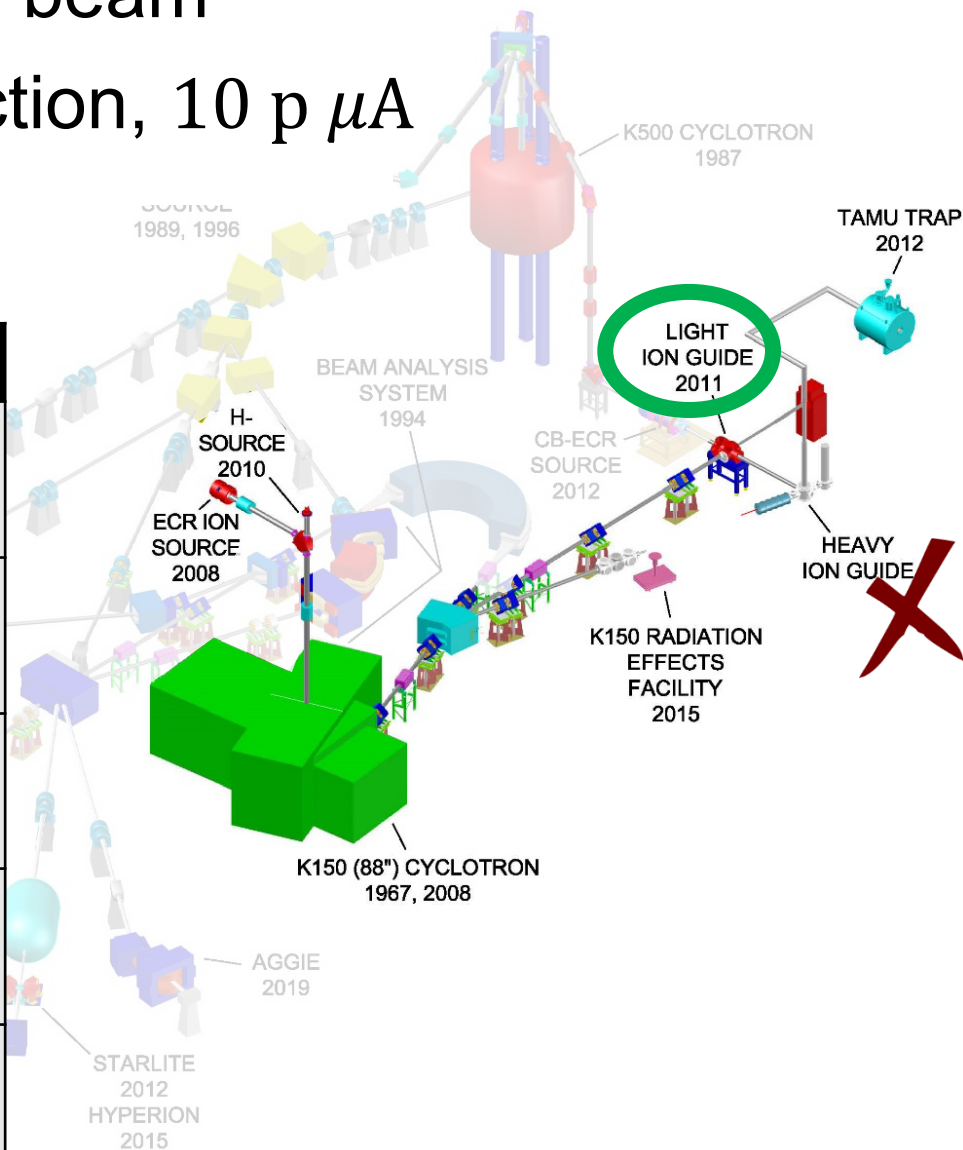
The proton-rich nuclei of interest

10-25 MeV/u ^3He primary beam

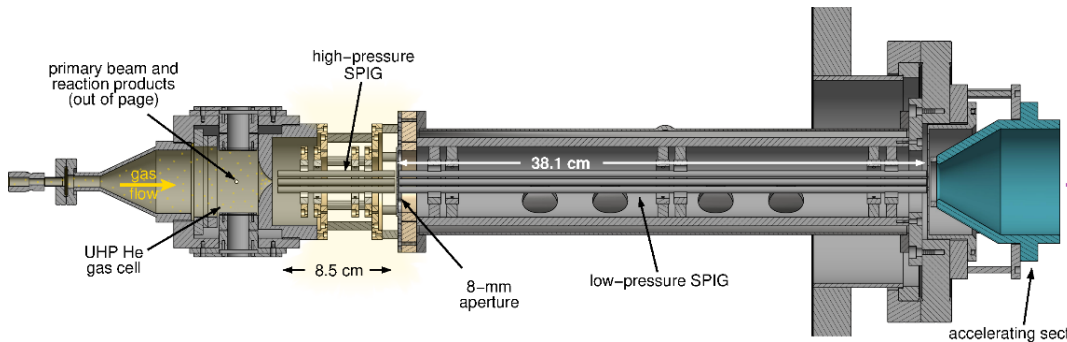
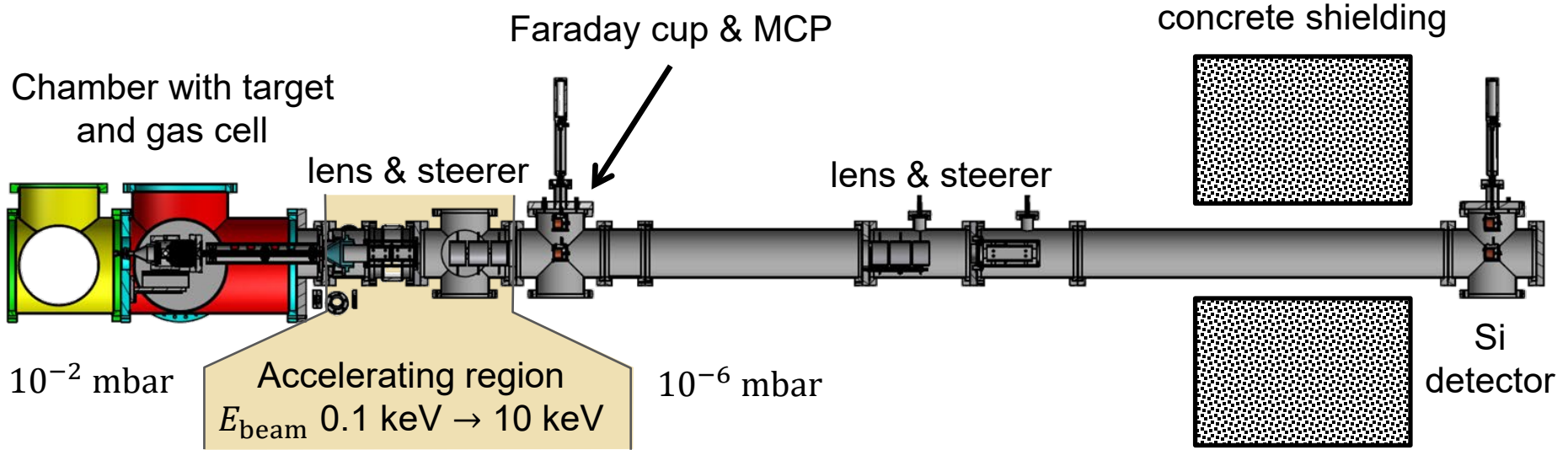
Fusion-evaporation reaction, $10 \text{ p } \mu\text{A}$

Target Product Production rate

^{20}Ne	^{20}Mg	4×10^3
	^{21}Mg	3×10^5
^{24}Mg	^{24}Si	3×10^3
	^{25}Si	2×10^5
^{28}Si	^{28}S	3×10^3
	^{29}S	8×10^4
^{32}S	^{32}Ar	0.9×10^3
	^{33}Ar	0.9×10^5
^{36}Ar	^{36}Ca	0.2×10^3
	^{37}Ca	0.2×10^5

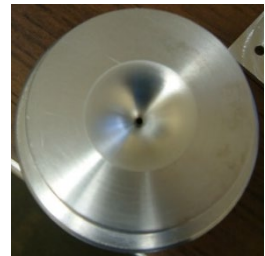
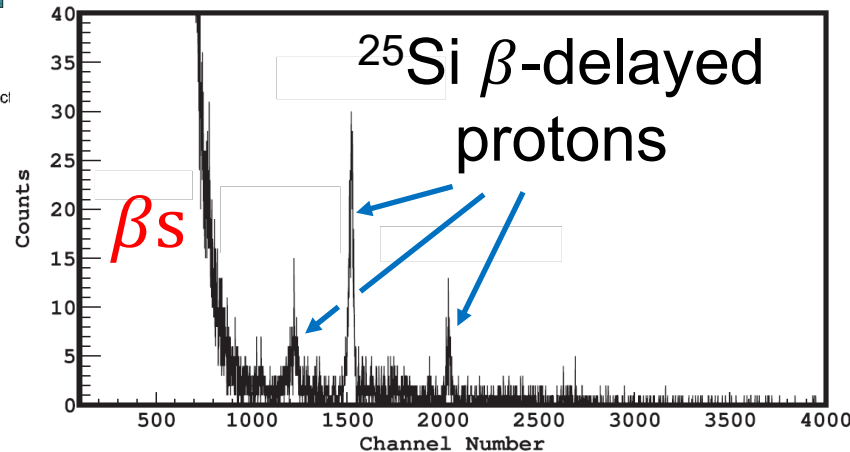


Tested prototype gas cell for ^3He LIG



50-keV ions to LSTAR

10 MeV/u ^3He on $^{24}(\text{nat})\text{Mg}$



- Efficiency $\sim 0.1\%$
- Improved design to be commissioned next year

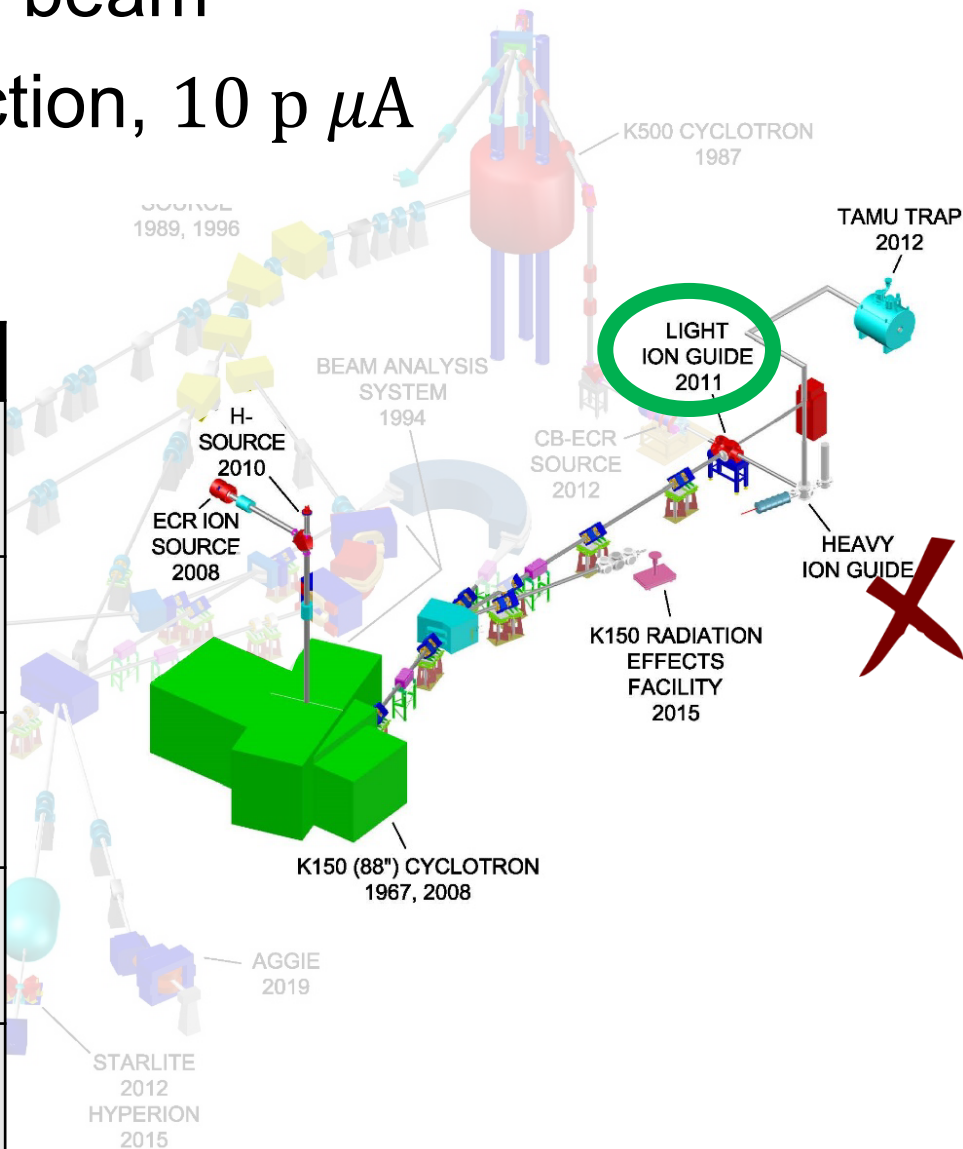
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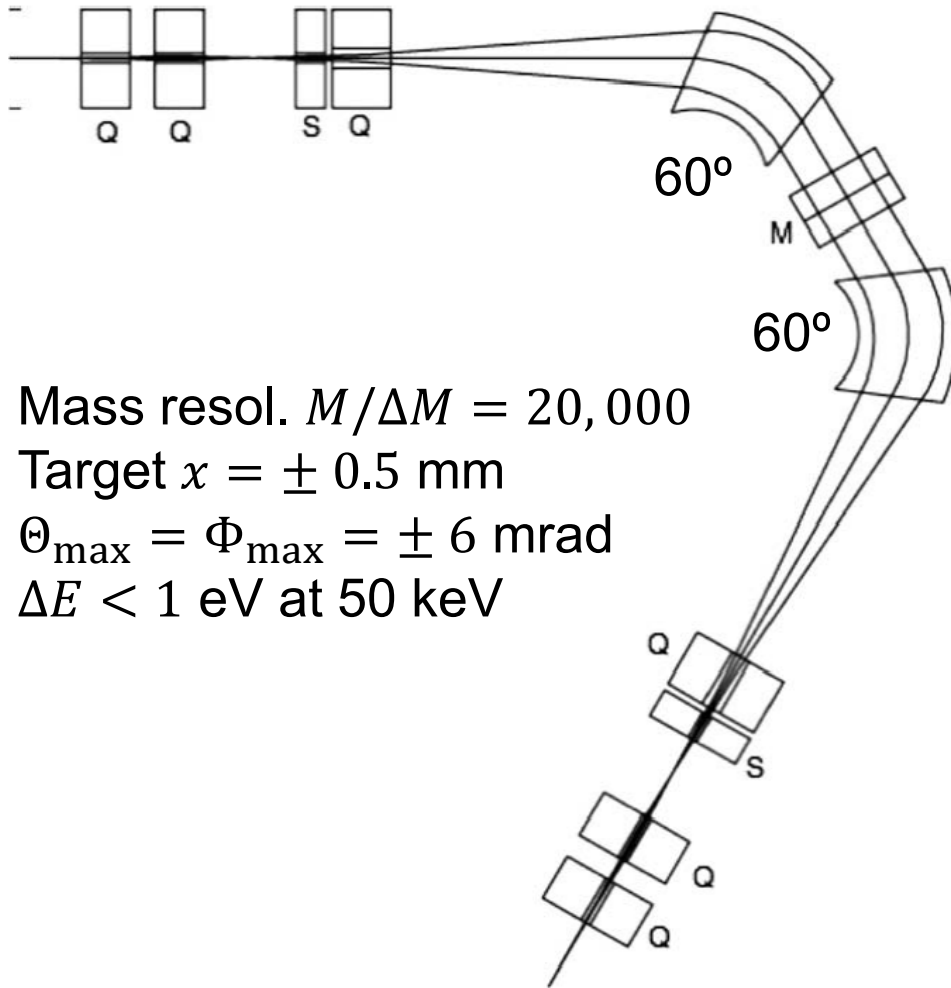
Fusion-evaporation reaction, $10 \text{ p } \mu\text{A}$

Target Product Production rate

Target	Product	Production rate
^{20}Ne	^{20}Mg	4×10^3
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^{28}Si	^{28}S	3×10^3
	^{29}S	8×10^4
^{32}S	^{32}Ar	0.9×10^3
	^{33}Ar	0.9×10^5
^{36}Ar	^{36}Ca	0.2×10^3
	^{37}Ca	0.2×10^5



Concept for mass separator: CARIBU (ANL)



- Mass resol. $M/\Delta M = 20,000$
- Target $x = \pm 0.5$ mm
- $\Theta_{\max} = \Phi_{\max} = \pm 6$ mrad
- $\Delta E < 1$ eV at 50 keV

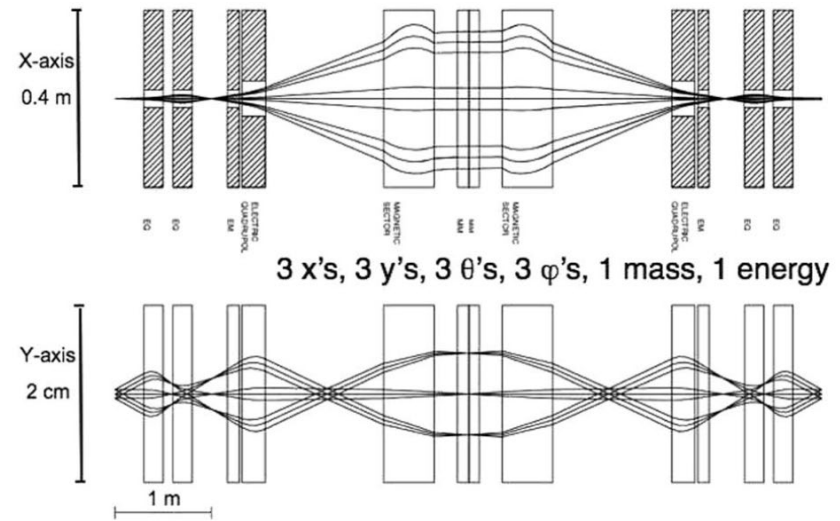


Fig. 1. Layout of the isobar separator in the x - (dispersive) and y - (non-dispersive) planes. Note the different scales

Dauids and Peterson,
NIMB 266, 4449 (2008)

Fig. 1. Layout of the isobar separator, showing the two 60° bending magnets with $\rho = 0.5$ m, two quadrupole (Q) doublets, two quadrupole singlets, two sextupole (S) singlets and an electrostatic multipole (M).

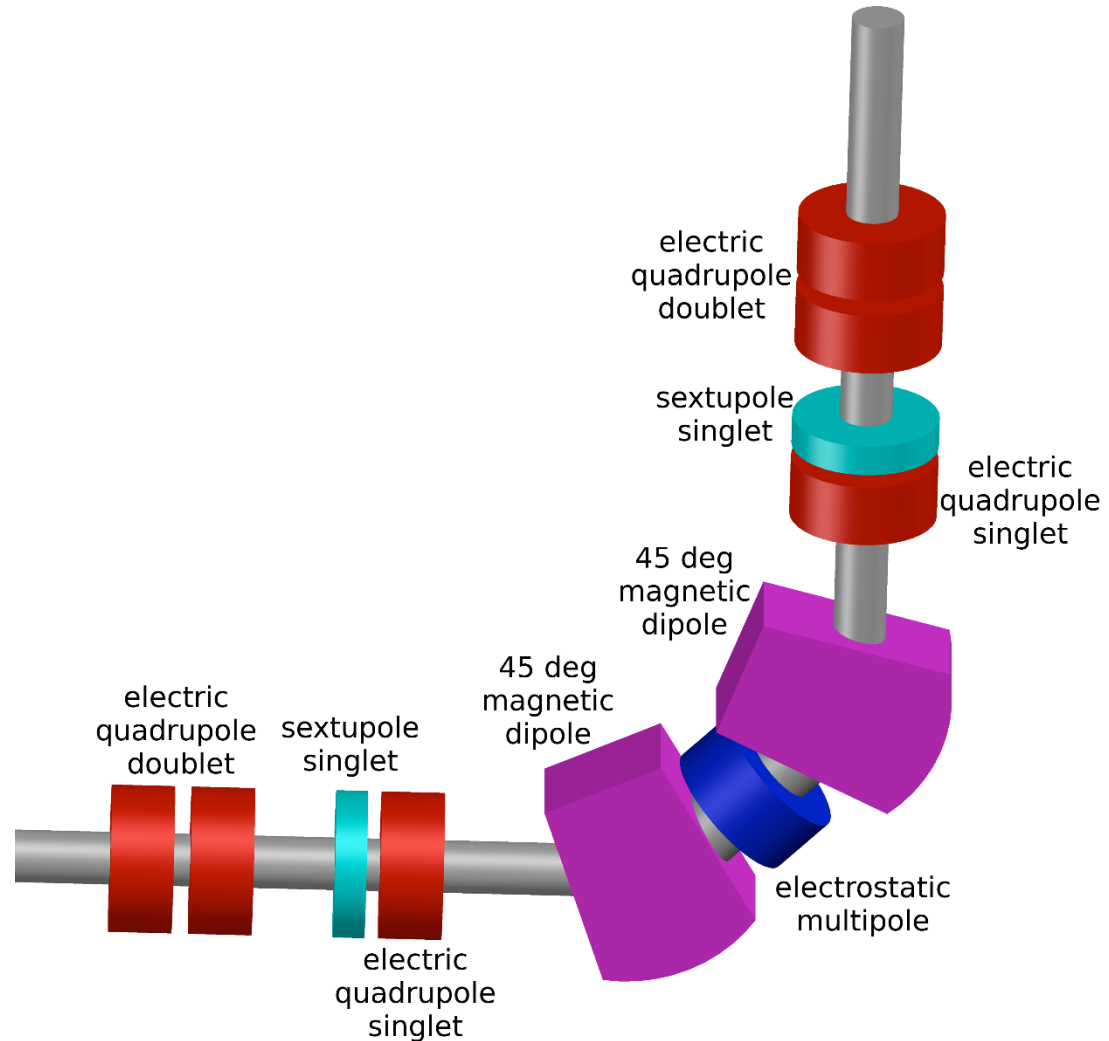
Our design



Georg Berg has come up with two designs:



☀️ $2 \times 45^\circ$ in a vertical direction



Our design

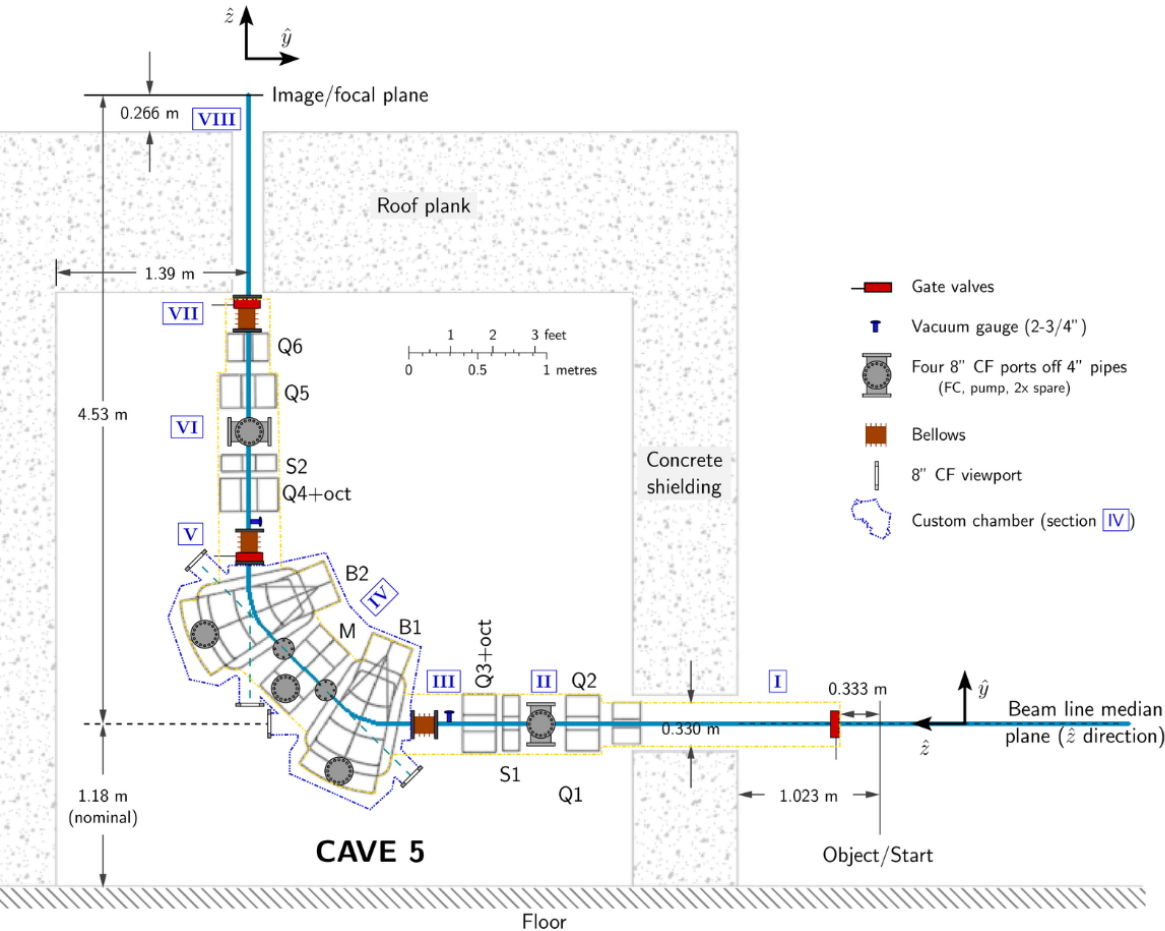


Georg Berg has come up with two designs:

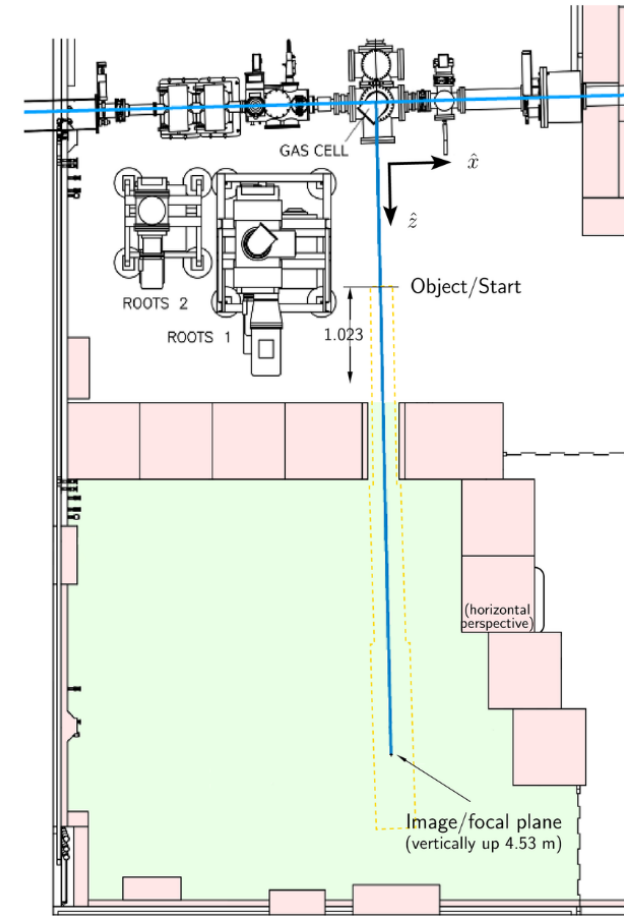


☀️ $2 \times 45^\circ$ in a vertical direction

Horizontal View



Top View



Our design

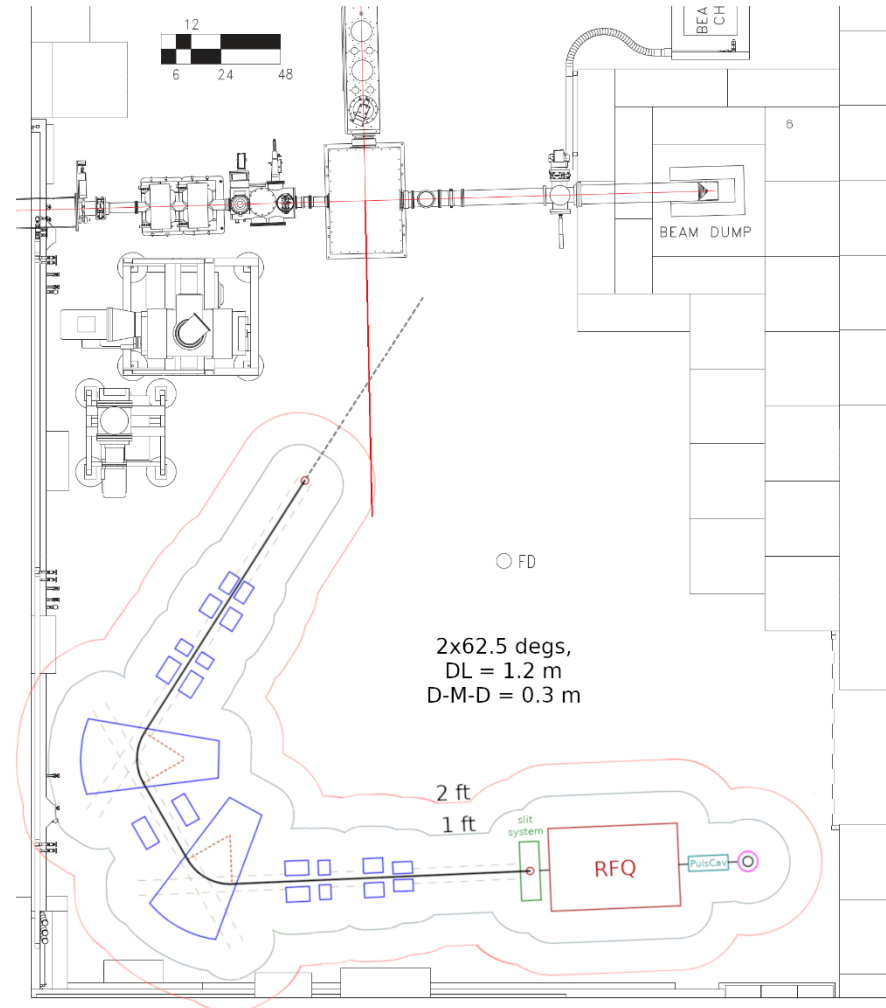
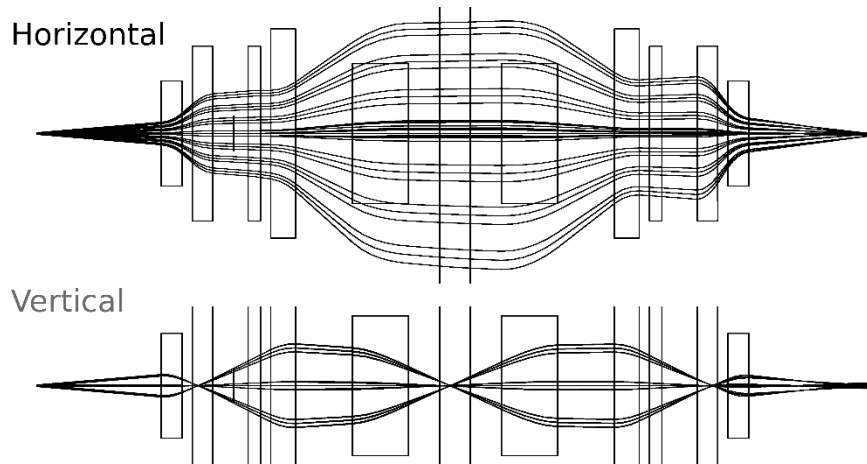


Georg Berg has come up with two designs:



☀ $2 \times 45^\circ$ in a vertical direction

☀ $2 \times 62.5^\circ$ in a horizontal plane



Our design

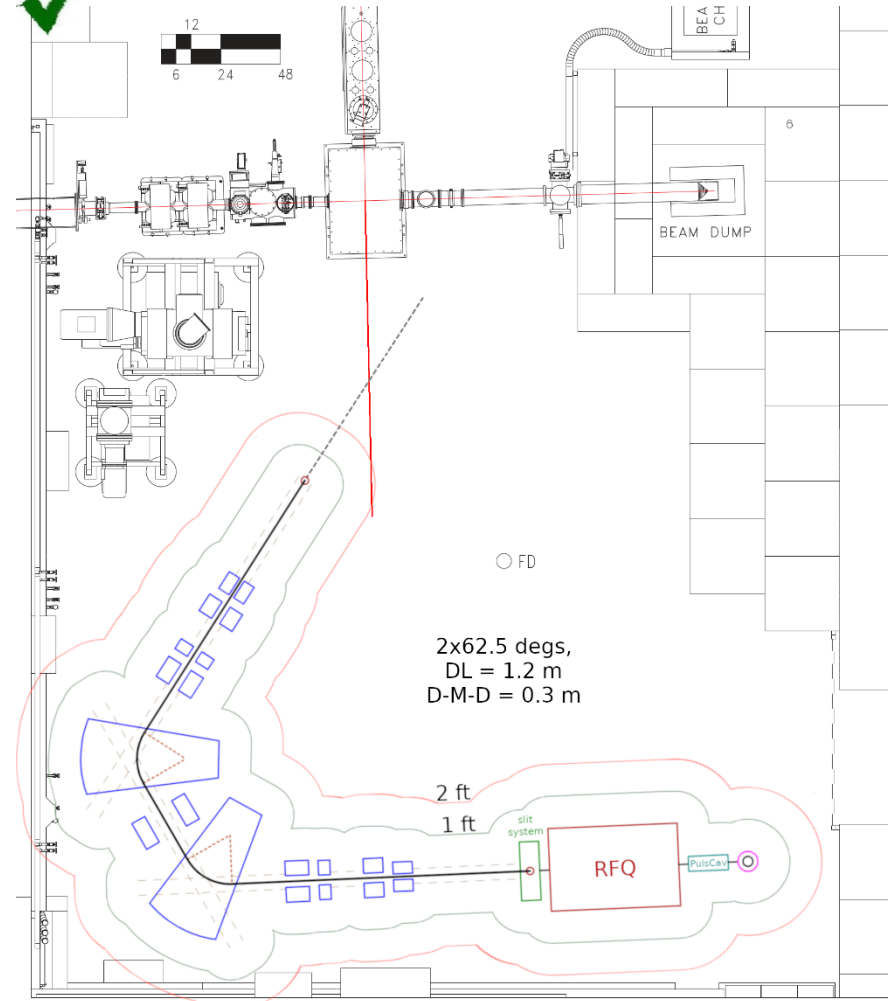
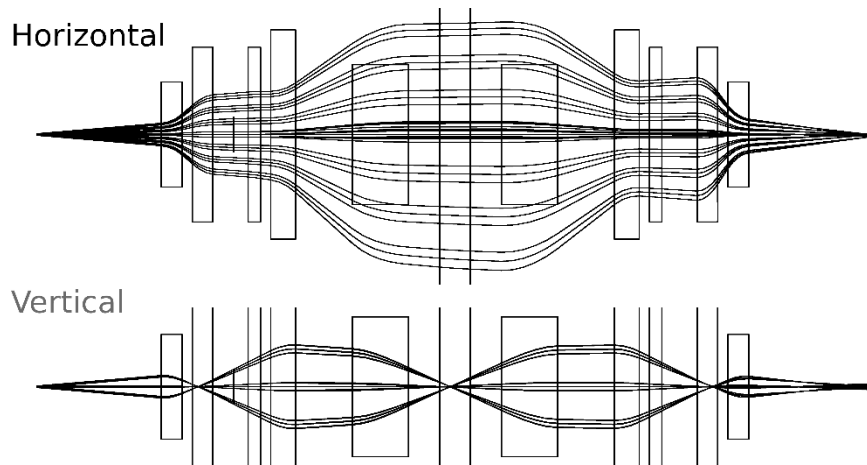


Georg Berg has come up with two designs:



✳️ $2 \times 45^\circ$ in a vertical direction ✖

✳️ $2 \times 62.5^\circ$ in a horizontal plane ✓

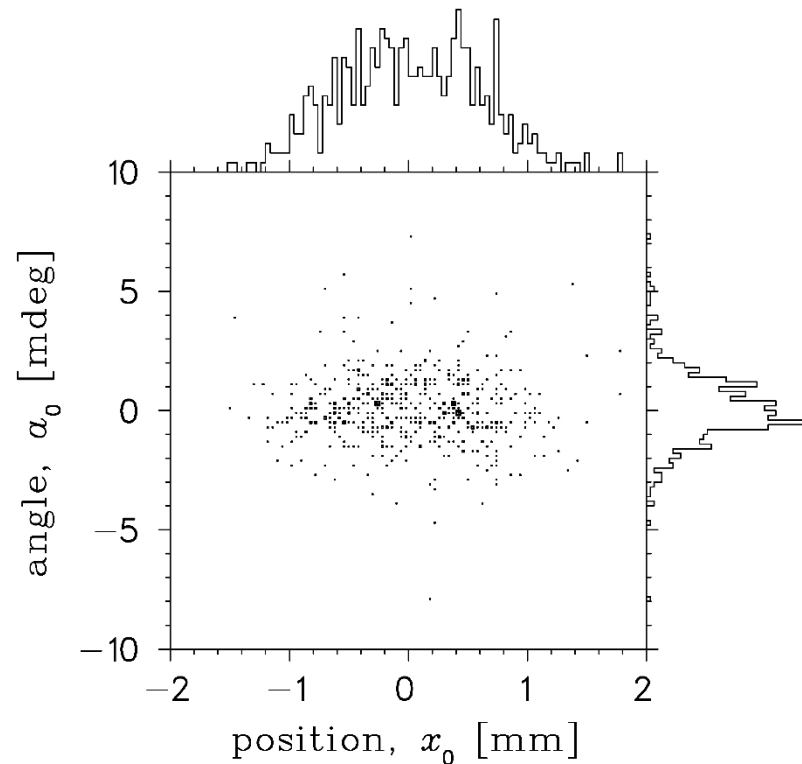


Calculated mass resolution:

- ✳️ For a 1-mm object beam,
 - 1st order: $M/\Delta M = 13\ 000$
 - 3rd order: $5\ 850$
 - 7th order: $5\ 480$

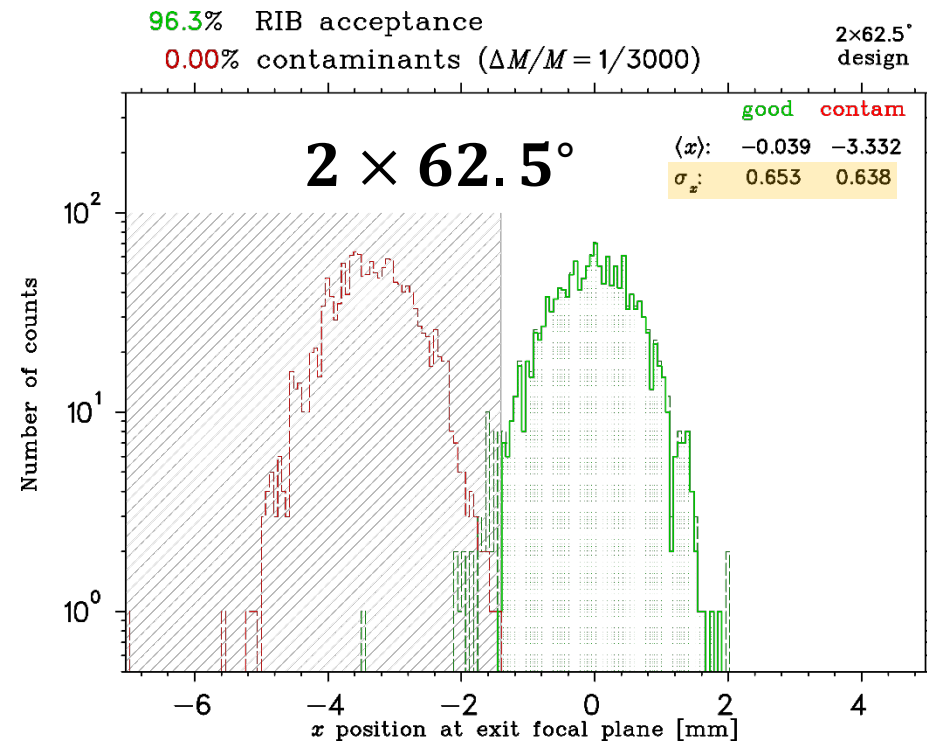
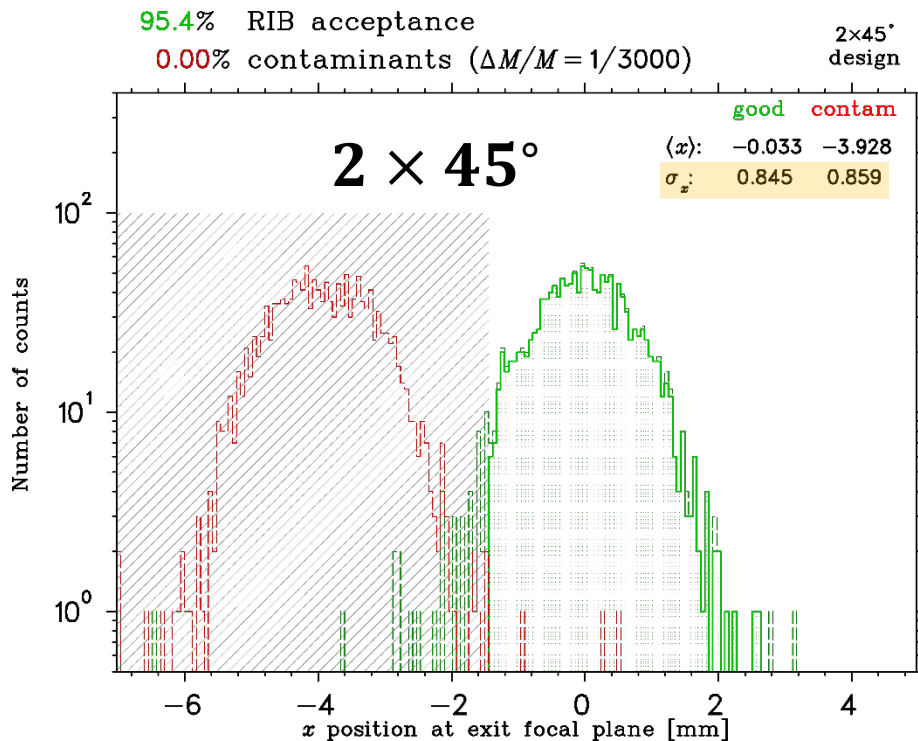
Since we know what we want...

- Simulate ion position and velocities from the He-LIG to the entrance of LSTAR
- With Einzel/steerers, can minimize either the spatial or the angular spreads
- Combined 0.65π mm mrad overall emittance is constant
- Use this as input for COSY ray tracing for a realistic estimation of the performance



First, let's compare $2 \times 45^\circ$ with $2 \times 62.5^\circ$

- Green curve: ion of interest (^{36}Ca , 65 keV, $\Delta E = 3.3$ eV)
- Red curve: nearest contaminant (^{36}K , $\Delta M/M = 3.057$)
- Solid/dashed: accepted/vetoed by separator
- New design has 30% smaller widths



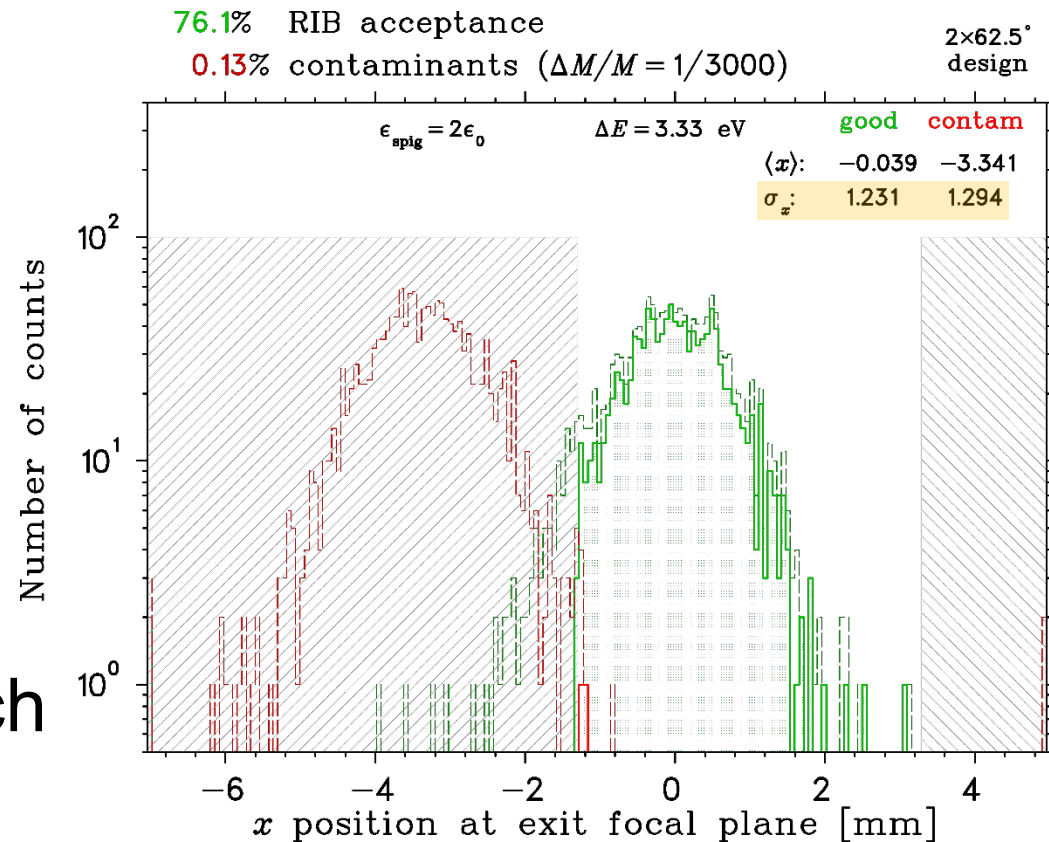
Let's get real...

How well can we trust SIMION?

Image of ion positions downstream of TAMUTRAP's RFQ was almost $2 \times$ larger than predicted

If we increase emittance by $2 \times$, widths increase by the same factor

Though we'd take a hit in efficiency, we can still remove nearby contaminants (molecules produced in the He-LIG might be much closer in mass...)

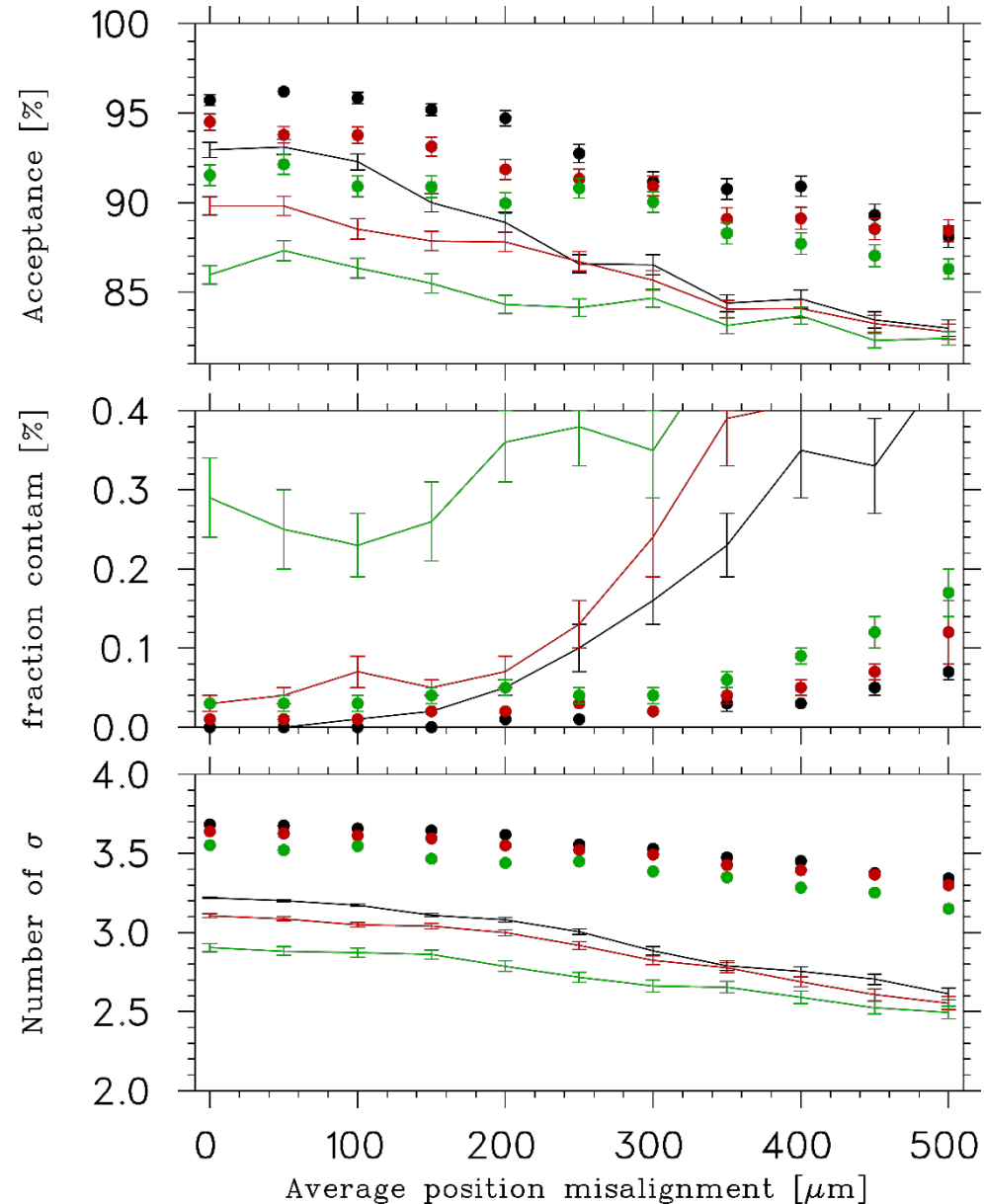


Position/angular misalignments?

Randomly move elements according to a Gaussian distr'n of width 0, 50, 100, ..., 500 μm , and tilts of 0 (black), 10 (red) and 20 mdeg (green)

$2 \times 45^\circ$ design (lines) is *much* more susceptible to misalignments than $2 \times 62.5^\circ$ (esp rotations)

Should have **>90% acceptance** with **± 0.25 mm** and **± 15 mdeg** tolerances



Looking forward

- Working in concert with the p -LIG group to utilize the K150 cyclotron for RIB production
- ^3He -LIG system should be commissioned next year
- It will probably take ~ 3 years to commission LSTAR
 - ✦ Supply-chain issues? Cost increase?
- Once commissioned for TAMUTRAP, the Cyclotron Institute will also have a **new beamline** providing **cooled and bunched proton-rich RIBs at low energy**
 - ✦ Decay spectroscopy
 - ✦ p decays for astrophysics? In TAMUTRAP?
- Looking only a little farther, it wouldn't be hard to charge-breed the ions and inject them in the K500 for **reaccelerated RIBs** (complementing the p -LIG effort)

Final thoughts, collaborators and thanks

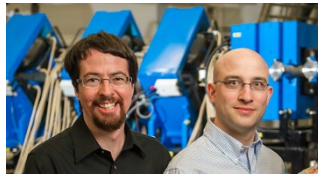
- TAMUTRAP: commissioned with offline sources, just need radioactive ions...
- ^3He -LIG: Initial tests were promising; commission new system early next year, measure emittance
- LSTAR: Design is finalized, specification document being written; bid request will be out before Christmas



G. Chubarian
V. Kolhinen
D. Melconian
P.D. Shidling
Gabriel Tabacaru
Operations group



G. Berg
M. Brodeur
M. Couder



Funded by



U.S. DEPARTMENT OF
ENERGY

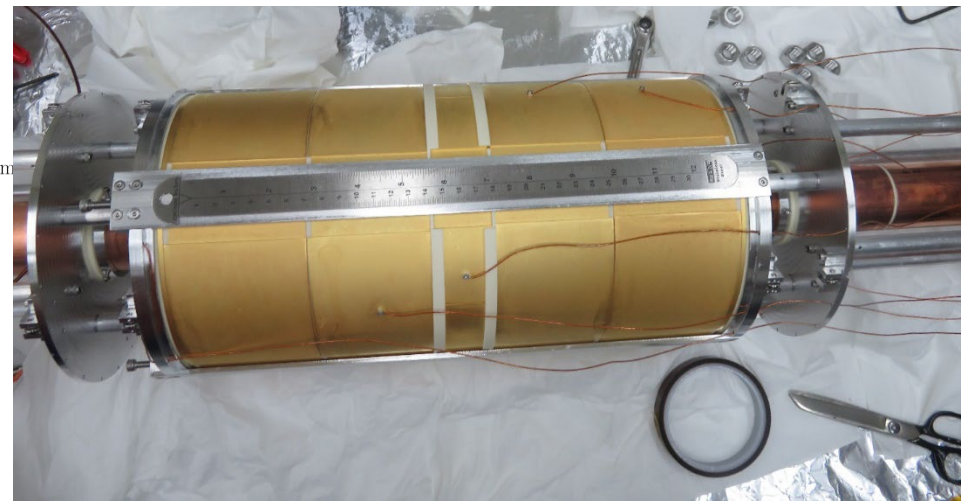
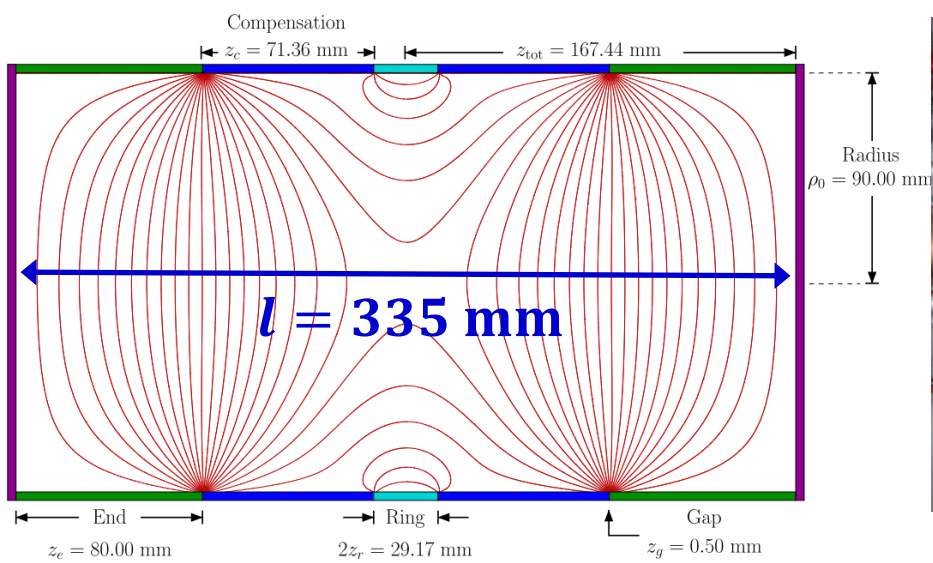
Office of
Science



The State of Texas

World's largest Penning trap commissioned

- Most cylindrical Penning traps have a length-to-radius ratio of $l/r = 11.75$
- To confine the protons from $T = 2$ decays, need $r = 90$ mm
- Needed a new design to make it fit in the 7T magnet



$$l/r = 3.72$$

M.Mehlman *et al.*, NIMA 712, 9 (2013)
P.Shidling *et al.*, Hyperfine Interact 240, 40 (2019)

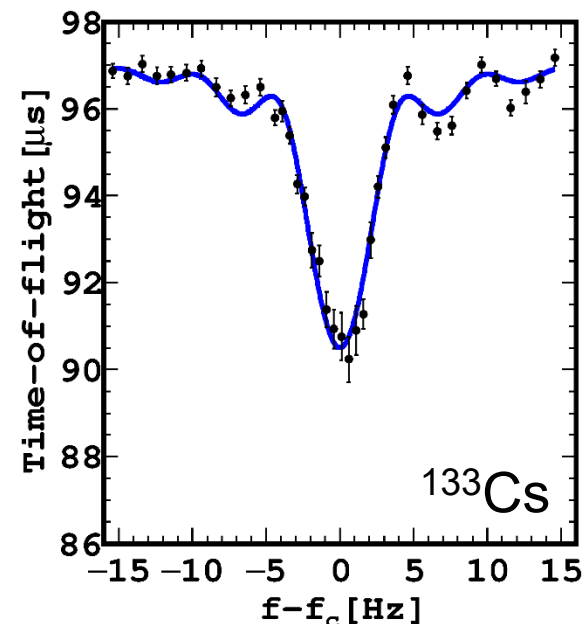
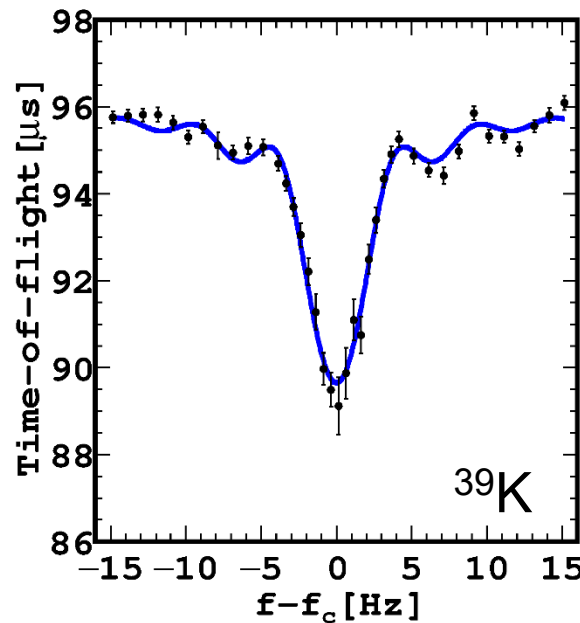
Mass measurement of stable ions

- Find resonant frequencies for ^{23}Na , $^{85,87}\text{Rb}$, ^{133}Cs & ^{39}K
- Use AME value for ^{39}K , and calculate other masses
- Good agreement with AME values (within uncertainties)

Precision

- ^{23}Na : 240 ppb
- ^{85}Rb : 5 ppb
- ^{87}Rb : 6 ppb
- ^{133}Cs : 7 ppb

200 ms excitation time

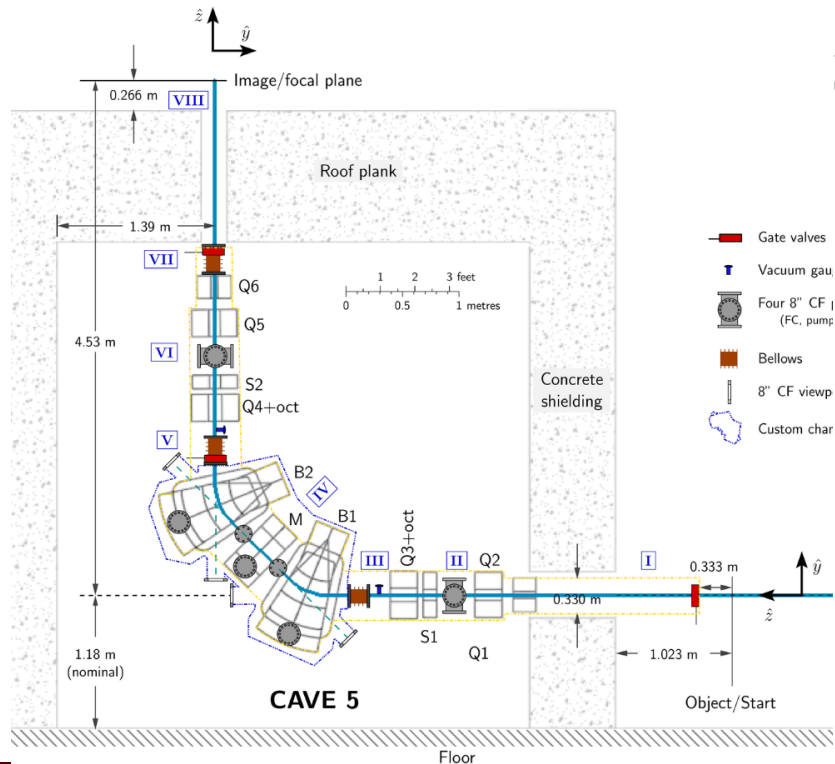


P.Shidling *et al.*, Int J Mass Spectr 468 (2021) 116636

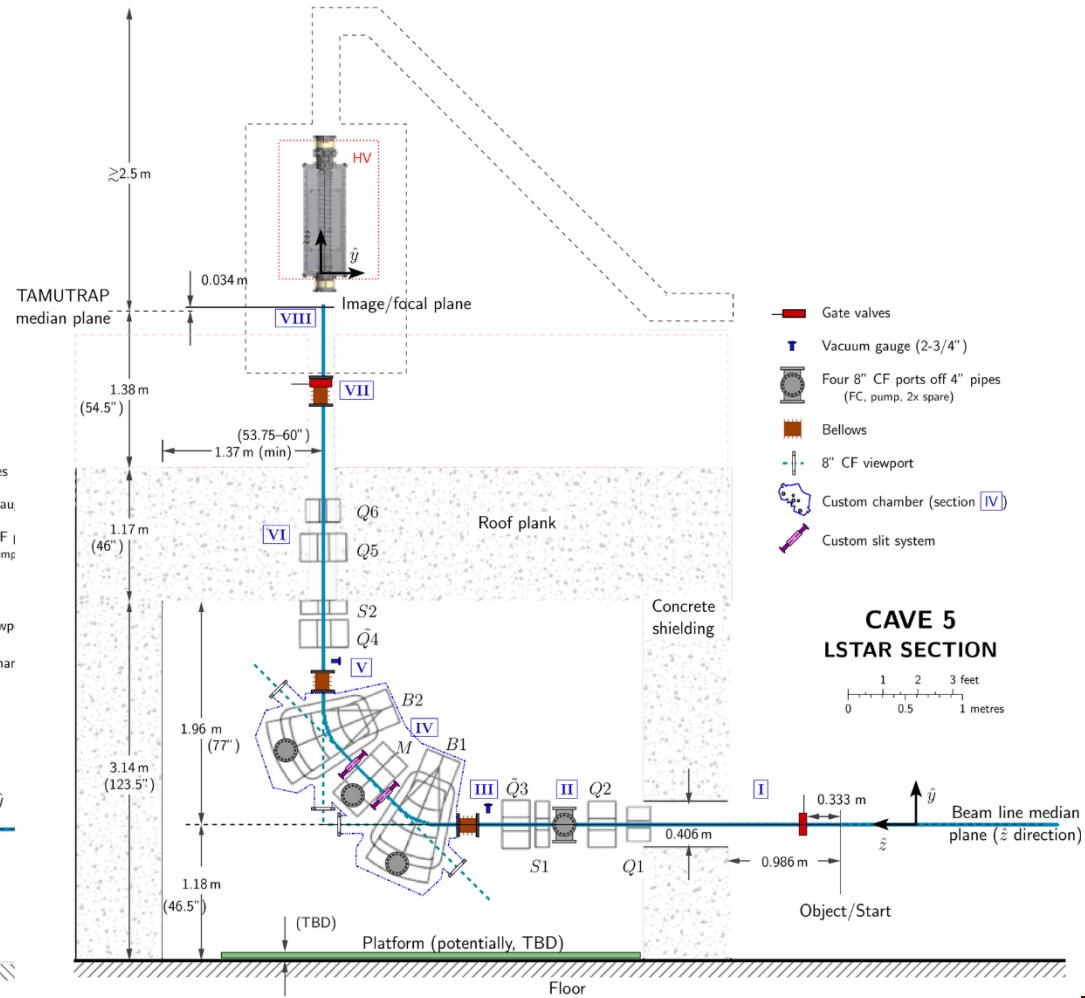
Why did we change our design?

As we were generating the specification document, we found a very significant discrepancy b/n AutoCAD and reality....

Horizontal View



Horizontal View



CAVE 5 LSTAR SECTION

