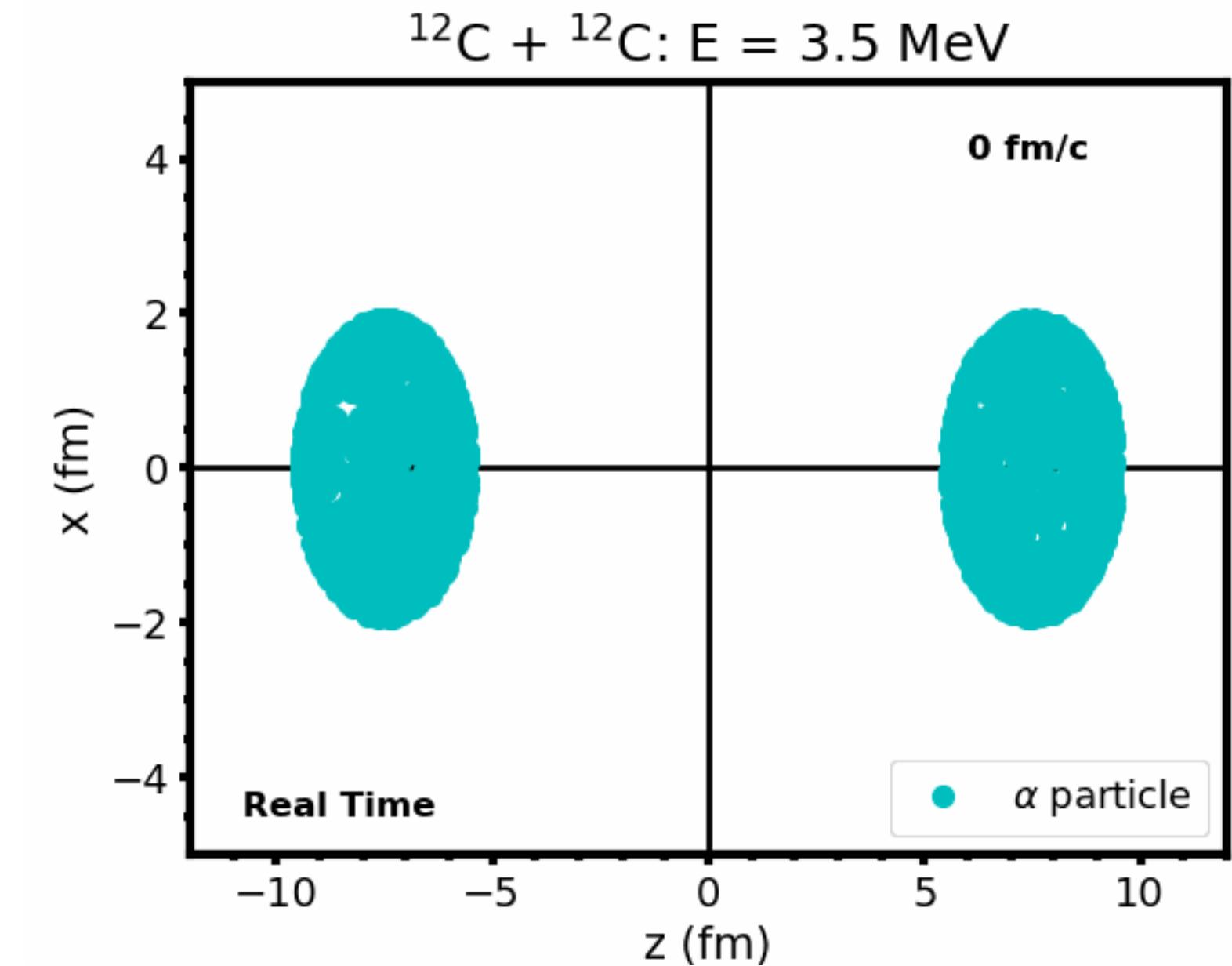


# Exploring sub-barrier fusion reactions

Aldo Bonasera

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LNS-INFN, Catania-Italy

## Introduction



Microscopic calculations: Vlasov dynamics in imaginary times.

Future directions-Constrained Molecular Dynamics (CoMD)

Hybrid Alpha Cluster model (T.Depastas).

Macroscopic calculations: Neck model in imaginary times (S. Sun).

Experimental challenges: pair production in sub barrier fusion reactions (T. Settlemyre)

## Conclusions

# Exploring sub-barrier fusion reactions

A.B. and J.B.Natowitz

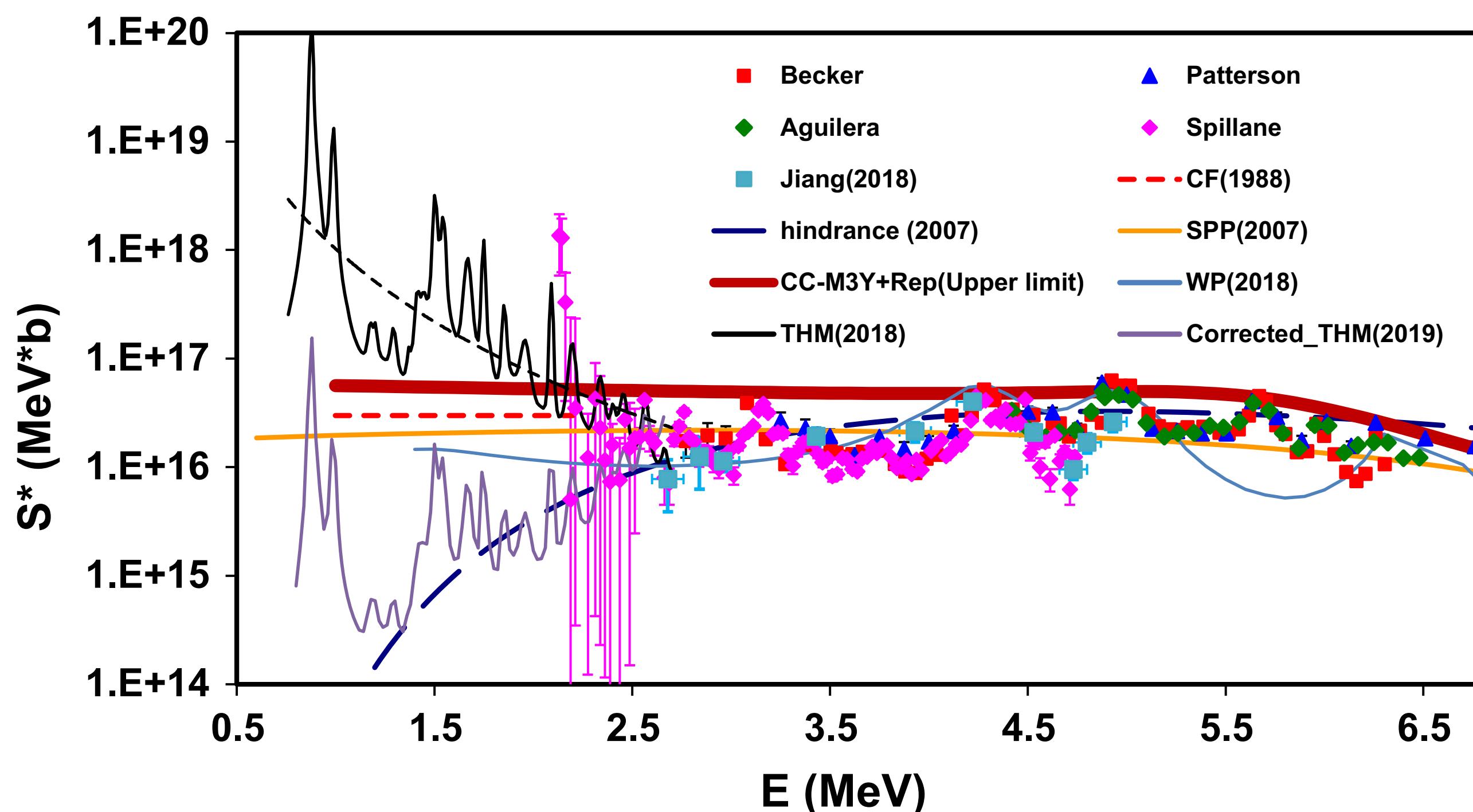
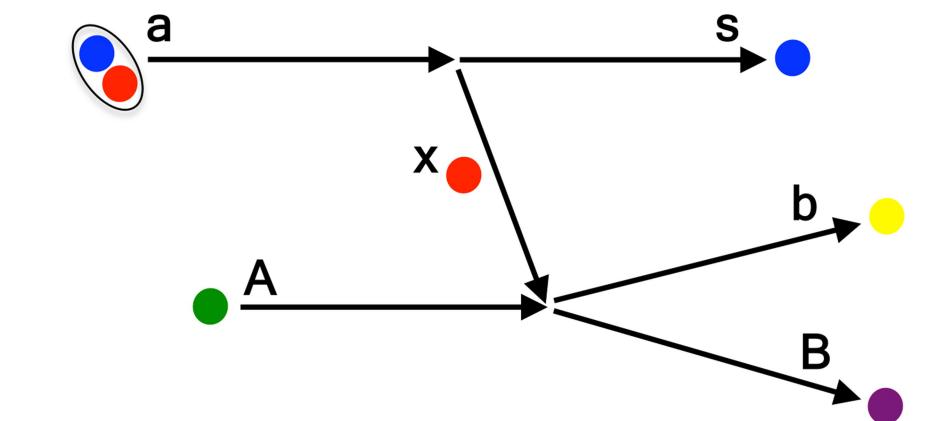
PHYSICAL REVIEW C 102, 061602(R) (2020)

## INTRODUCTION&MOTIVATION

Letter 688 | Nature | VOL 557 | 31 MAY 2018

### An increase in the $^{12}\text{C} + ^{12}\text{C}$ fusion rate from resonances at astrophysical energies

A. Tumino<sup>1,2\*</sup>, C. Spitaleri<sup>2,3</sup>, M. La Cognata<sup>2</sup>, S. Cherubini<sup>2,3</sup>, G. L. Guardo<sup>2,4</sup>, M. Gulino<sup>1,2</sup>, S. Hayakawa<sup>2,5</sup>, I. Indelicato<sup>2</sup>, L. Lamia<sup>2,3</sup>, H. Petrascu<sup>4</sup>, r. G. Pizzone<sup>2</sup>, S. M. r. Puglia<sup>2</sup>, G. G. rapisarda<sup>2</sup>, S. romano<sup>2,3</sup>, M. L. Sergi<sup>2</sup>, r. Spartá<sup>2</sup> & L. Trache<sup>4</sup>



Status on  $^{12}\text{C} + ^{12}\text{C}$  fusion at deep subbarrier energies:  
impact of resonances on astrophysical  $S^*$  factors

C. Beck<sup>1,a</sup>, A. M. Mukhamedzhanov<sup>2,b</sup>, X. Tang<sup>3,4,c</sup>

Eur. Phys. J. A (2020) 56:87

$$\begin{aligned} S^*(E_{\text{c.m.}}) &= E_{\text{c.m.}} \sigma(E_{\text{c.m.}}) \exp(87.12E_{\text{c.m.}}^{-1/2} + 0.46E_{\text{c.m.}}) \\ &= S(E_{\text{c.m.}}) \exp(0.46E_{\text{c.m.}}) \end{aligned} \quad (1)$$

# Feynman path integration in phase space

Physics Letters B 339 (1994) 207–210

Aldo Bonasera, Vladimir N. Kondratyev<sup>1</sup>

Phys.Rev.Lett.78(1997)187

Laboratorio Nazionale del Sud, INFN, v. S. Sofia 44, 95125 Catania, Italy

Solve the Vlasov equation in imaginary time. Define collective variables R&P

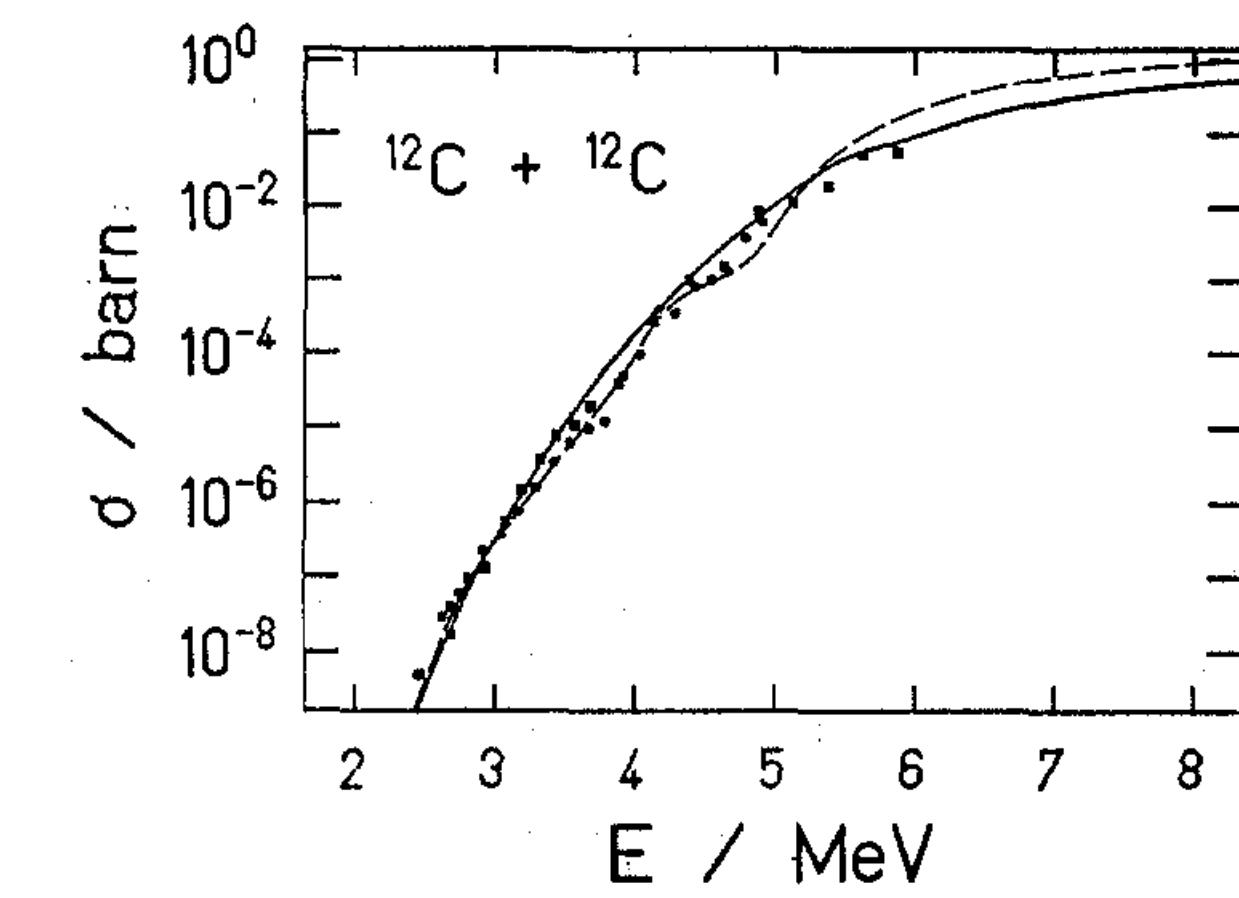
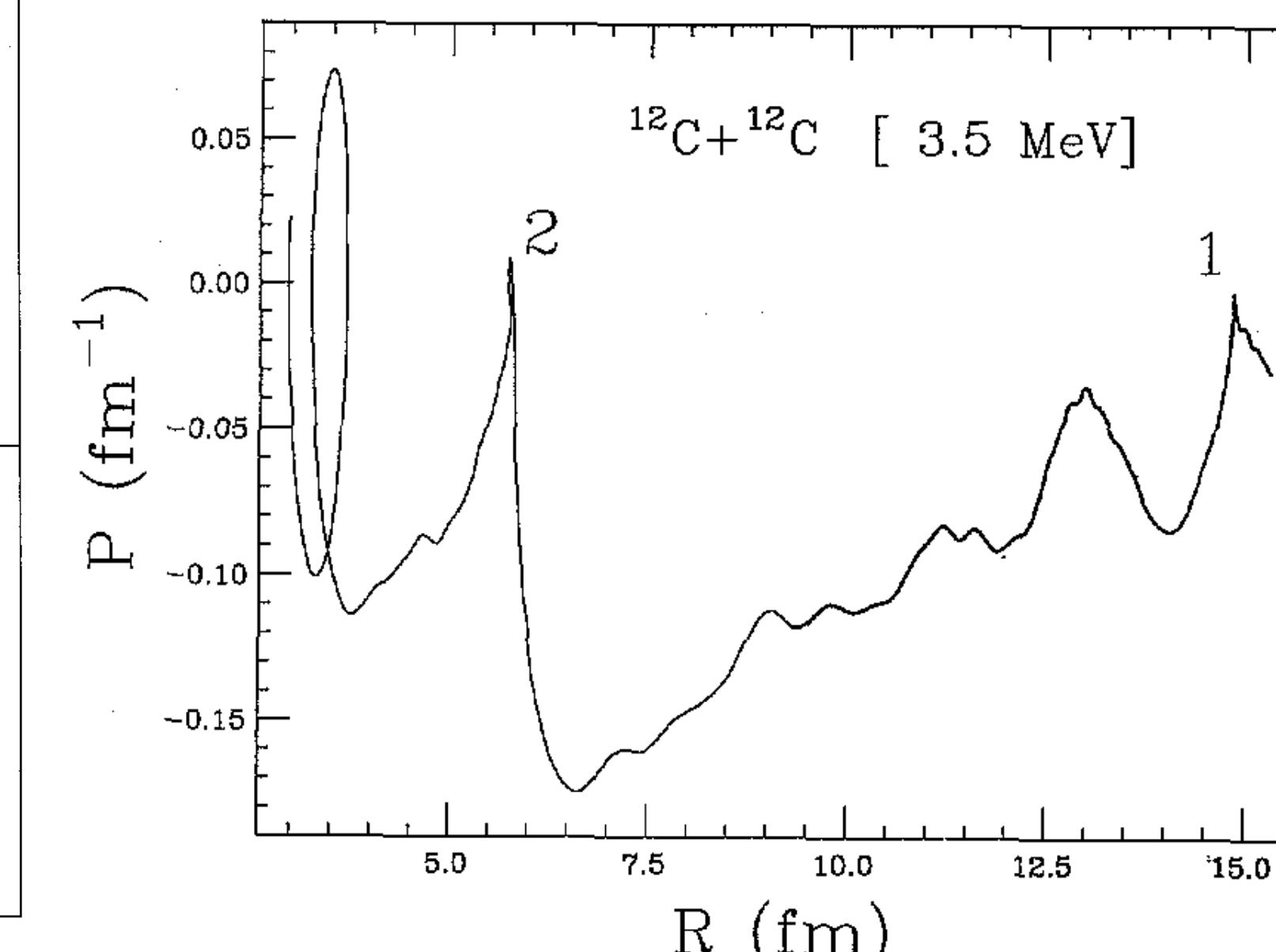
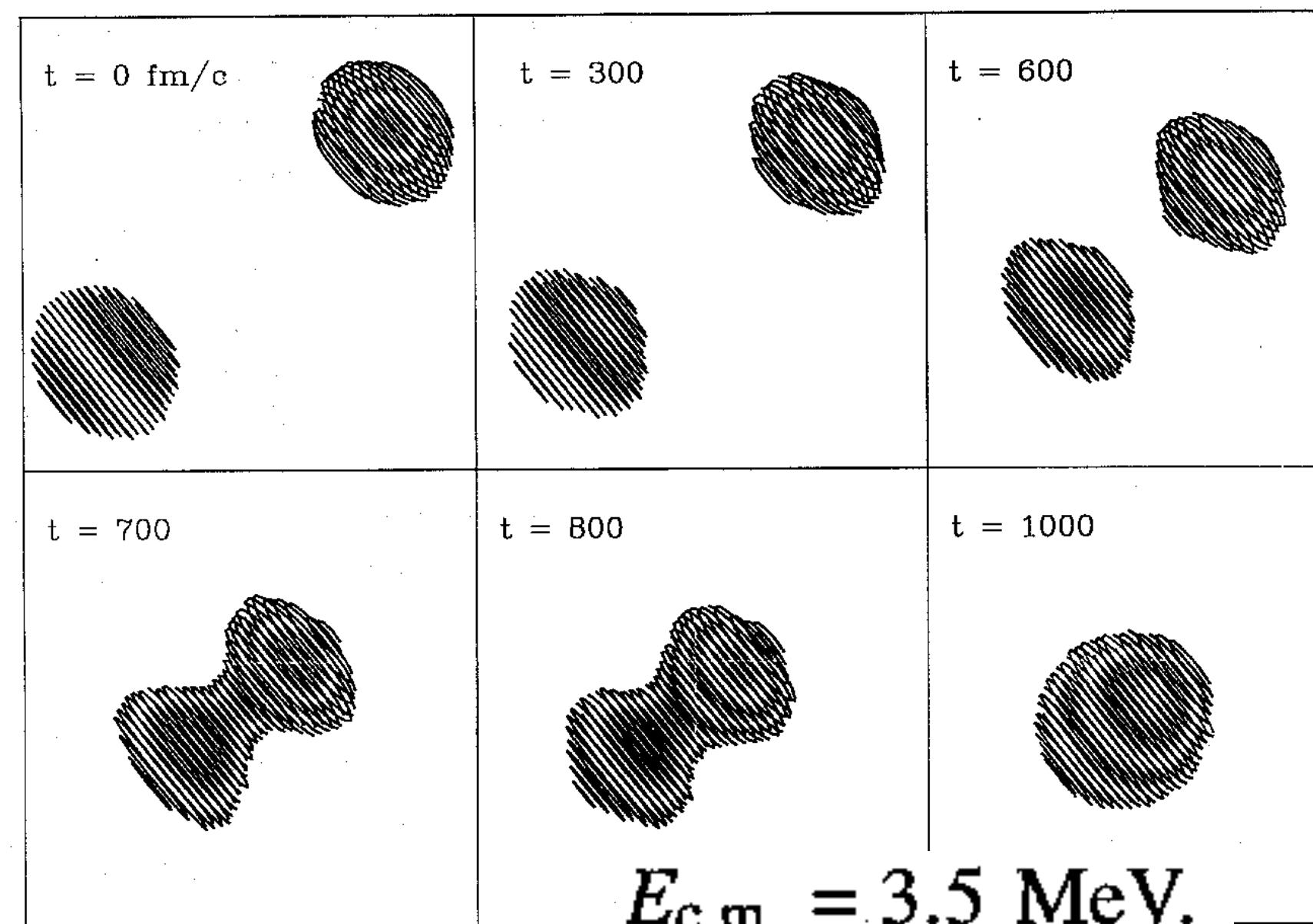
$$\left\{ \begin{array}{l} \mathbf{R} \\ \mathbf{P} \end{array} \right\}_A = \int d\mathbf{r} d\mathbf{p} \left\{ \begin{array}{l} \mathbf{r} \\ \mathbf{p} \end{array} \right\} f(\mathbf{r}, \mathbf{p}; t)$$

$$\frac{d\mathbf{R}_{A(B)}}{dt} = \frac{\mathbf{P}_{A(B)}}{m} ; \quad \frac{d\mathbf{P}_{A(B)}}{dt} = \mathbf{F}_{A(B)}$$

$$- \int_B d\mathbf{r} d\mathbf{p} \left\{ \begin{array}{l} \mathbf{r} \\ \mathbf{p} \end{array} \right\} f(\mathbf{r}, \mathbf{p}; t)$$

in imaginary time  $t \rightarrow it$

$$\frac{d\mathbf{R}_{A(B)}^i}{dt} = \frac{\mathbf{P}_{A(B)}^i}{m} ; \quad \frac{d\mathbf{P}_{A(B)}^i}{dt} = -\mathbf{F}_{A(B)}$$



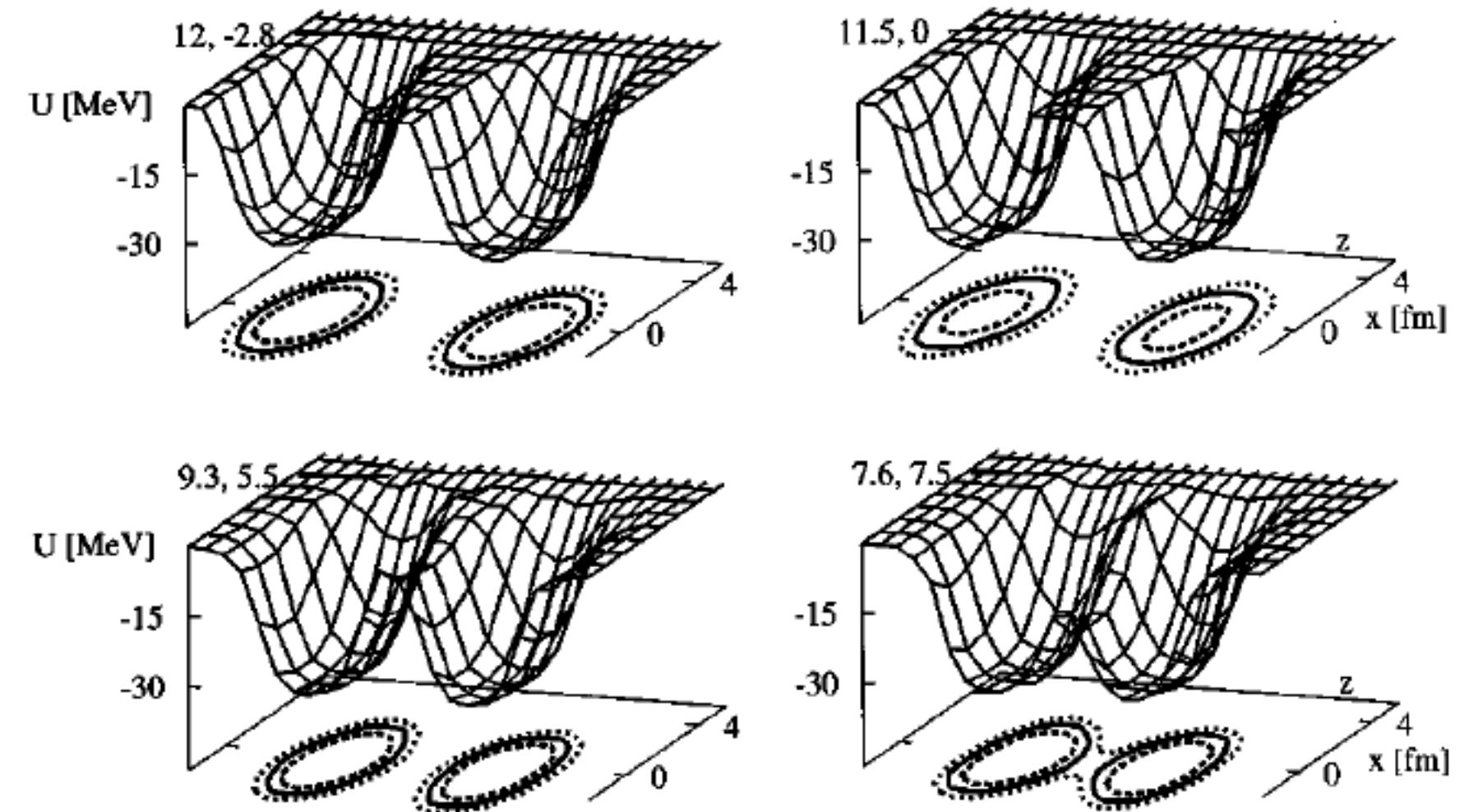


FIG. 1. Snapshots from the mean-field simulation of the fusion reaction for a head-on  $^{16}\text{O} + ^{16}\text{O}$  collision at the energy 8 MeV. The

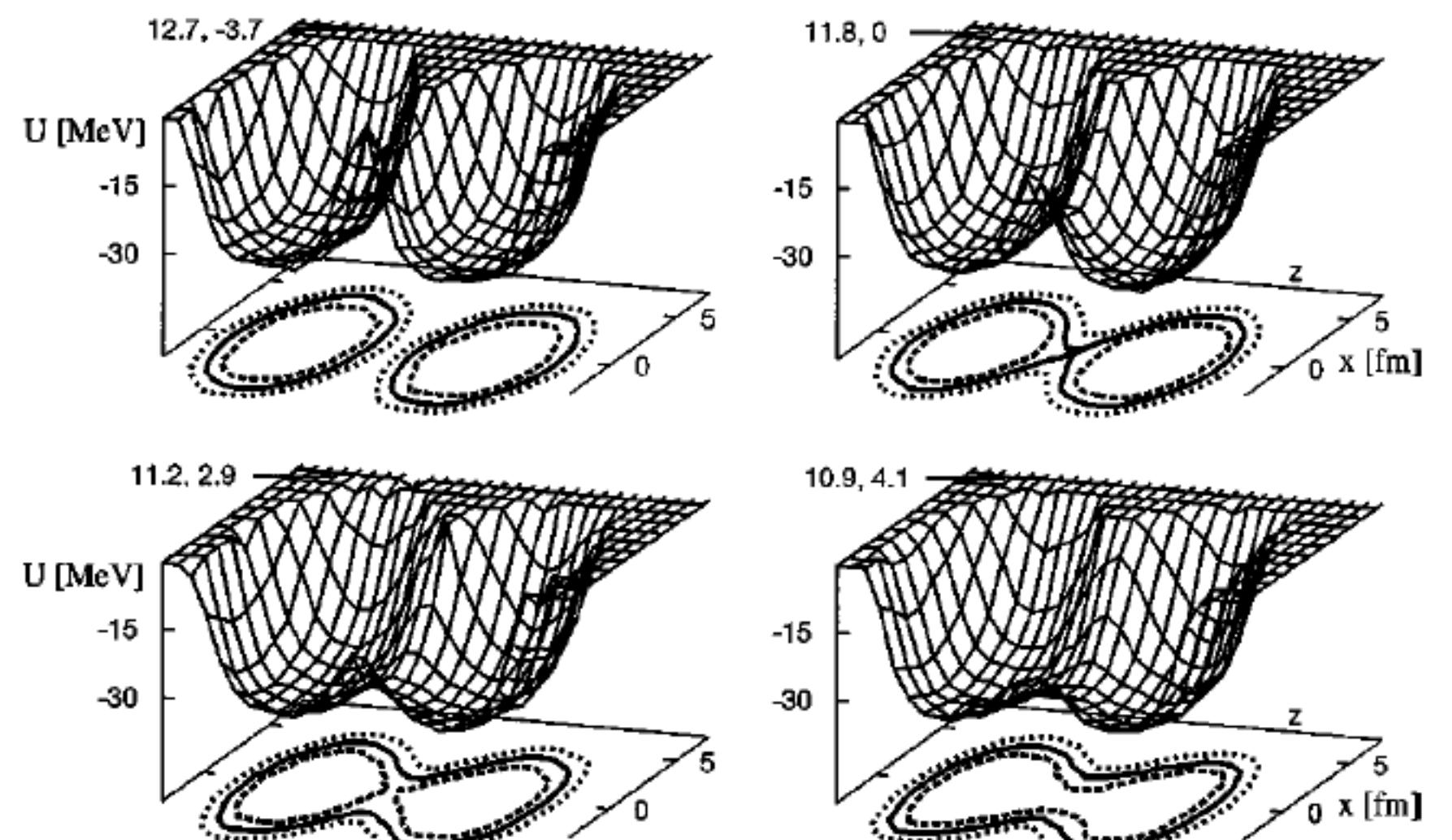


FIG. 2. The same as Fig. 1 for a  $^{58}\text{Ni} + ^{58}\text{Ni}$  collision at the energy 93 MeV.

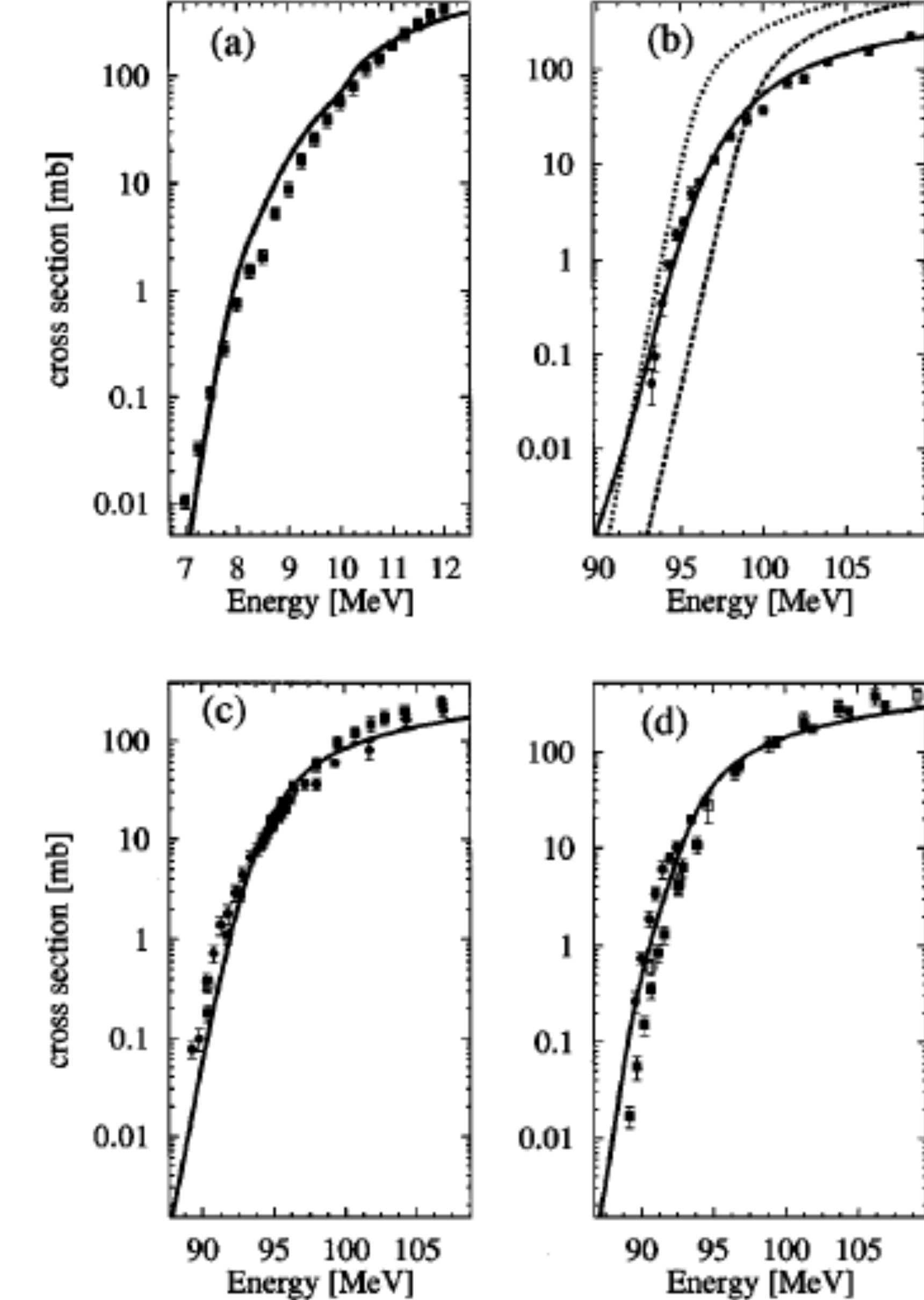
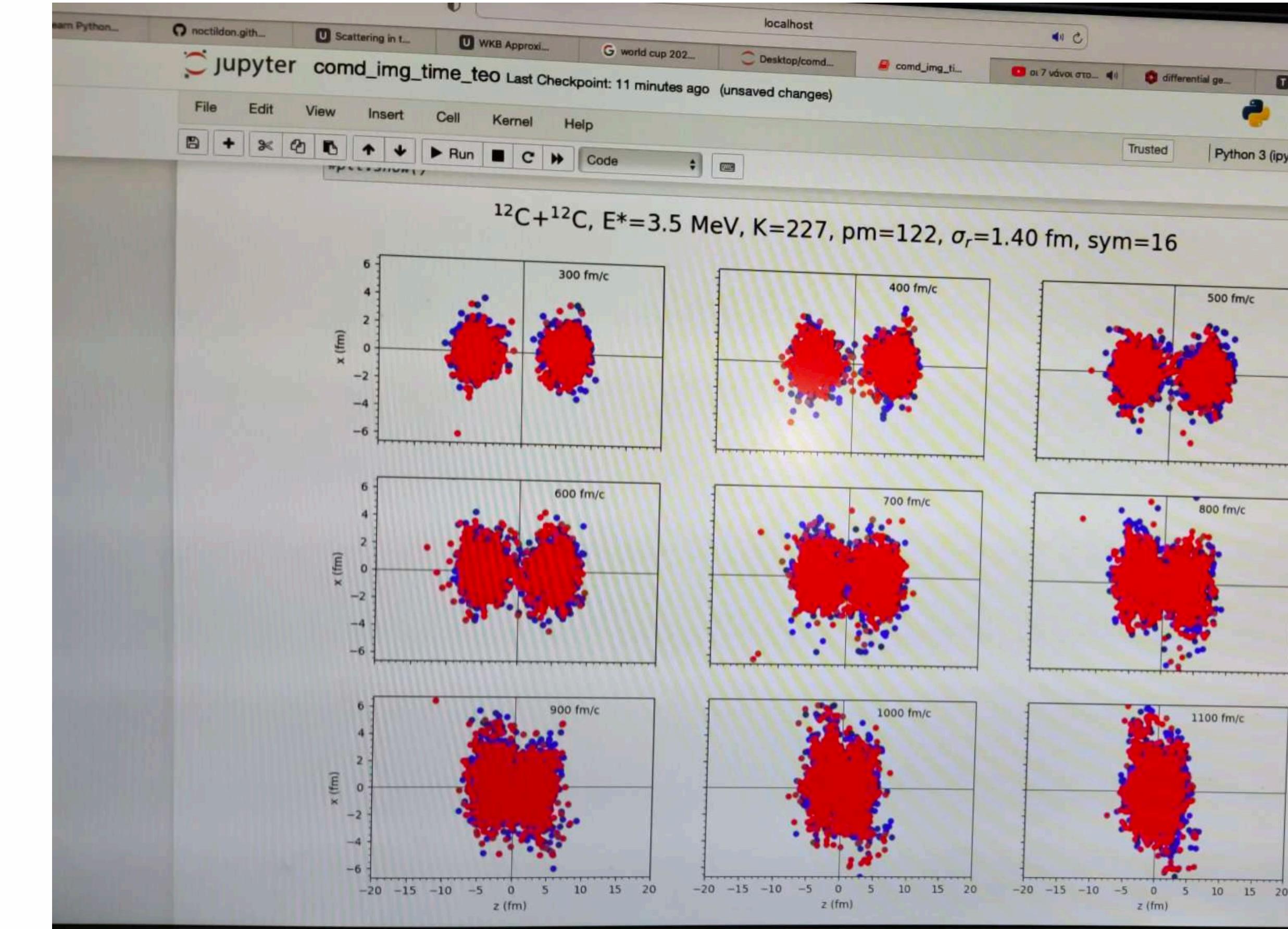
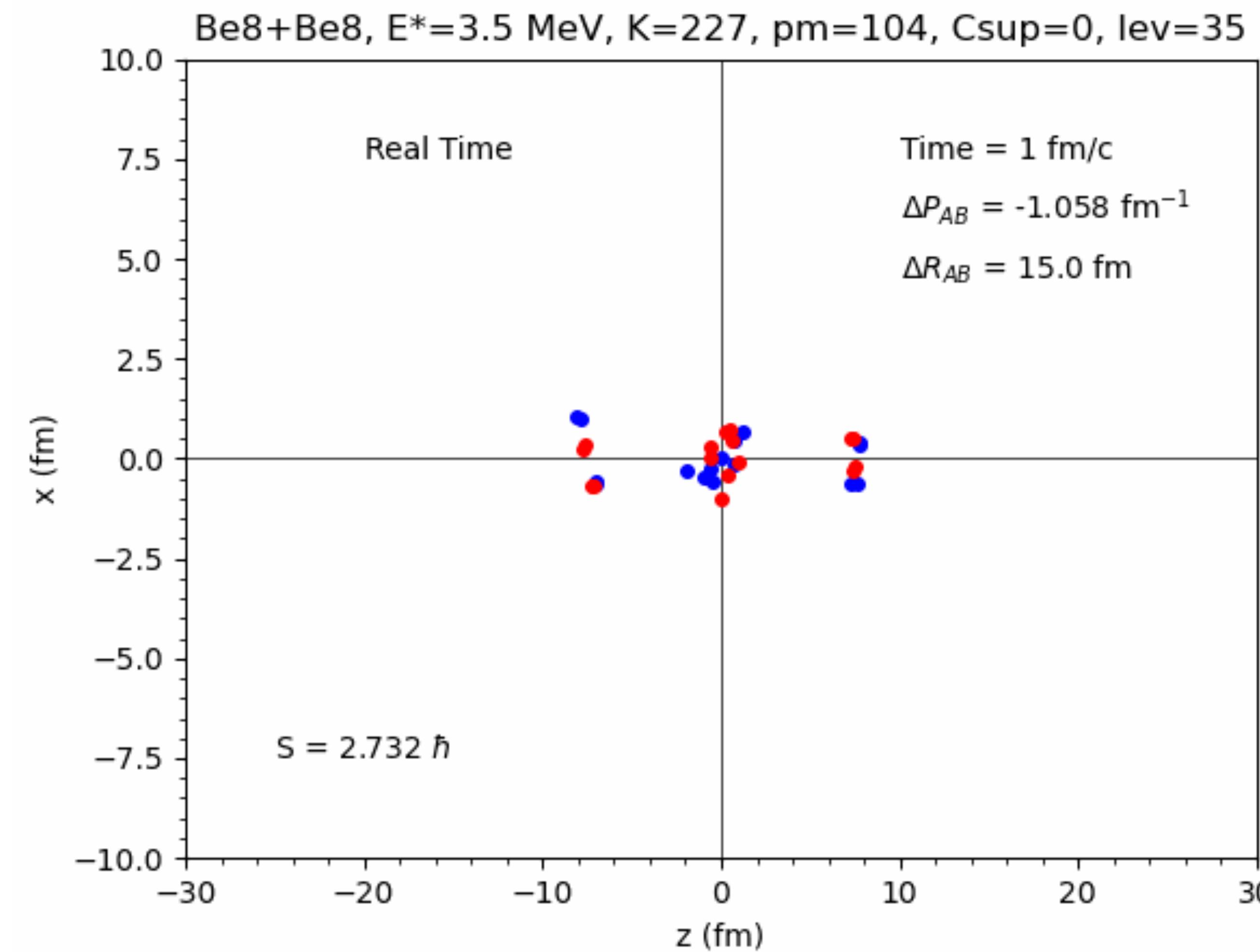


FIG. 5. The fusion excitation functions. Part (a) represents the  $^{16}\text{O} + ^{16}\text{O}$  system with the experimental data from Ref. [47]. Parts (b)–(d) display the results for the isotope pairs: (b)  $^{58}\text{Ni} + ^{58}\text{Ni}$ ; (c)  $^{58}\text{Ni} + ^{64}\text{Ni}$ ; and (d)  $^{64}\text{Ni} + ^{64}\text{Ni}$ . The circles show the experimental

# Microscopic calculations in progress- T. Depastas

Preliminary attempts using CoMD

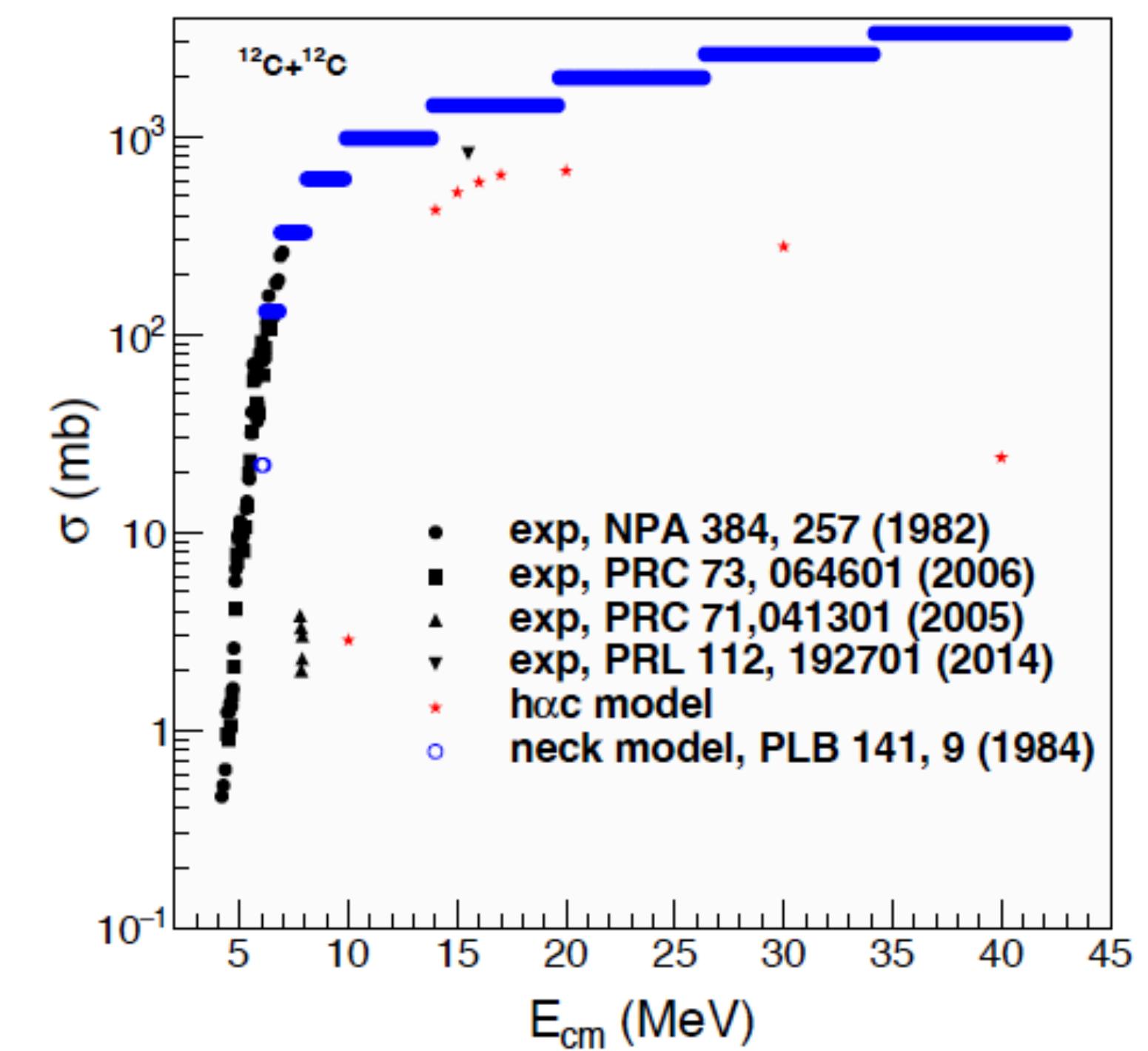
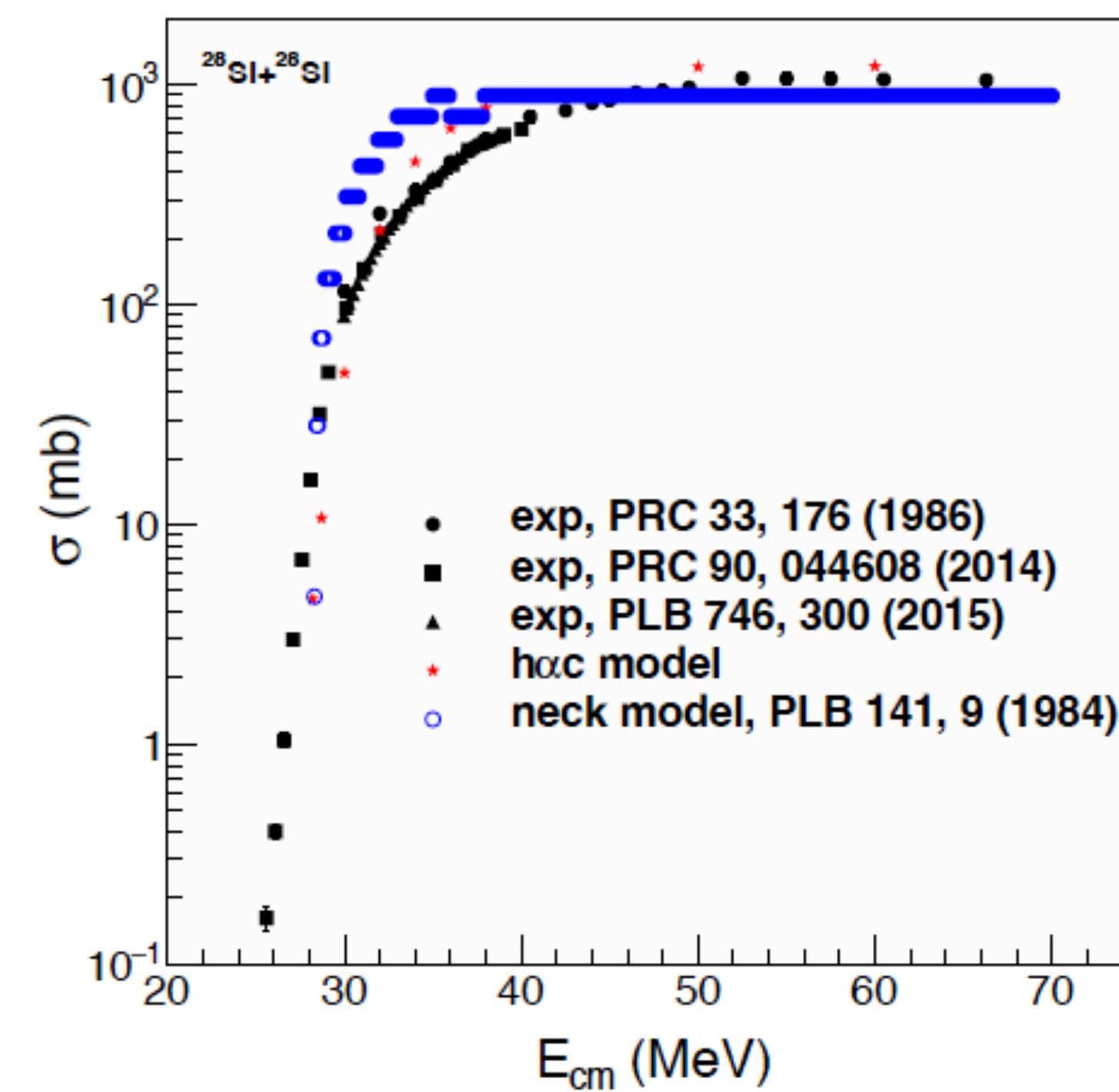
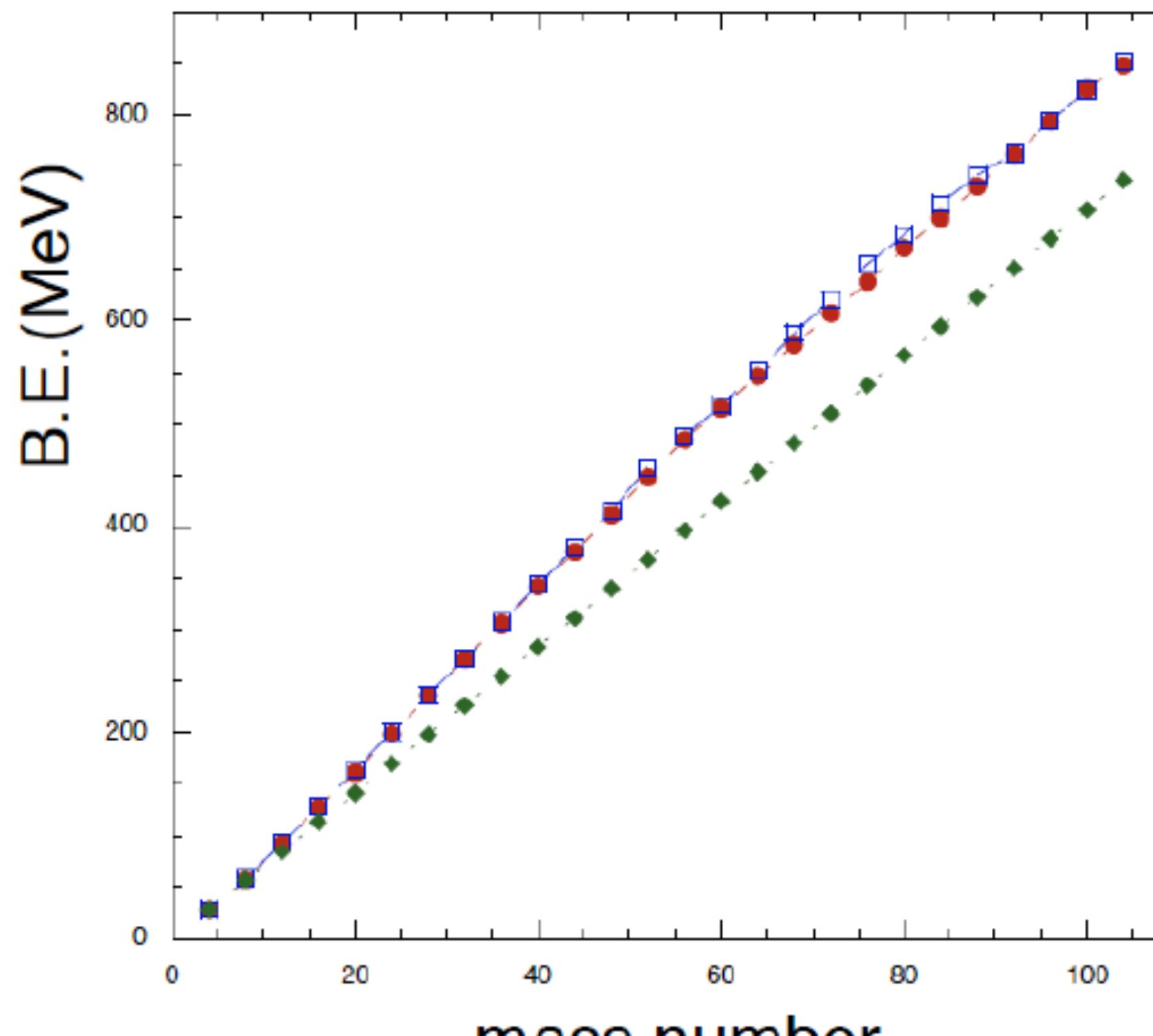


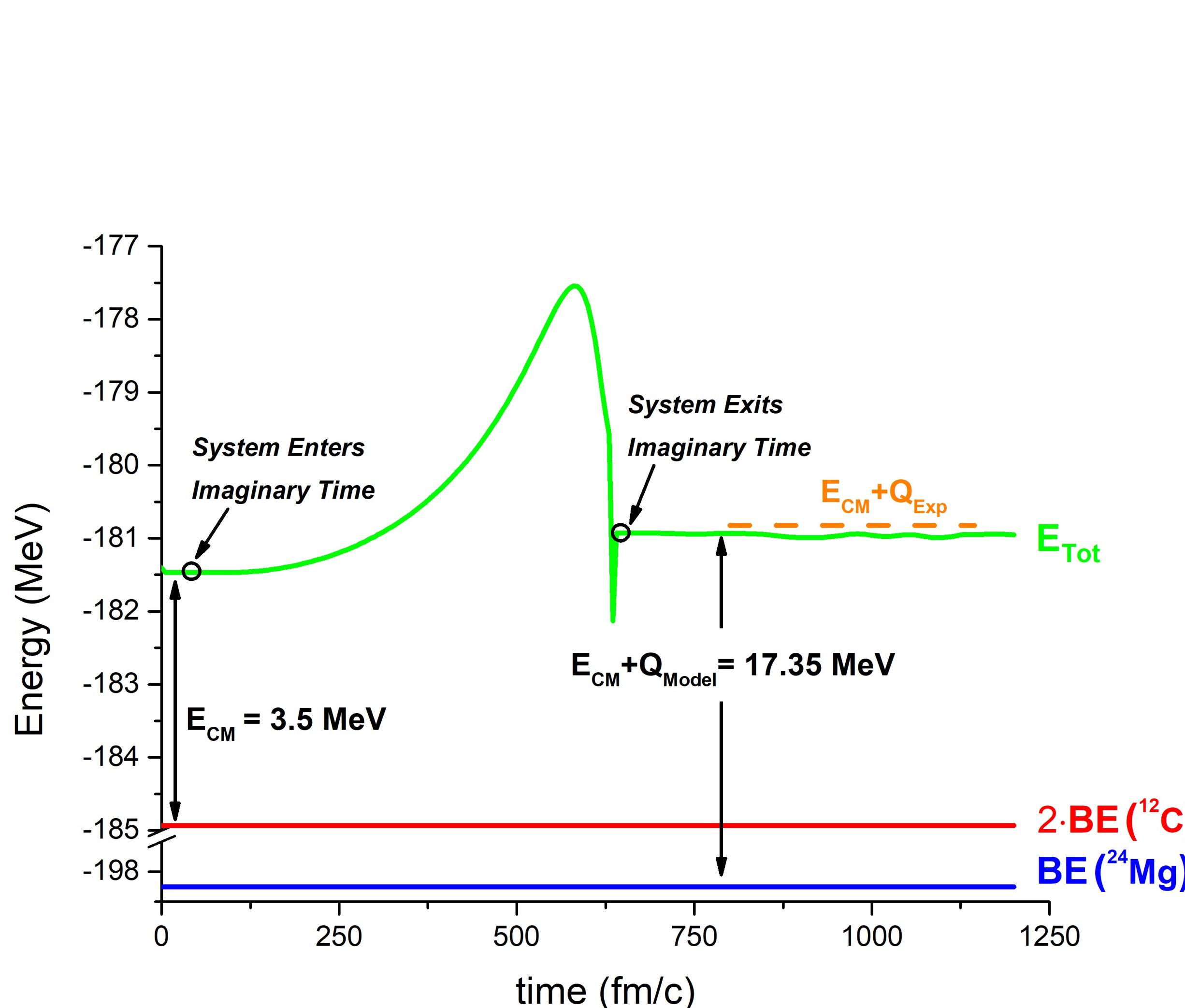
# Hybrid Alpha Cluster model (HAC)

Zheng, H.; Bonasera, A. **Spin Quantization in Heavy Ion Collision.** Symmetry 2021, 13, 1777. <https://doi.org/10.3390/sym13101777>

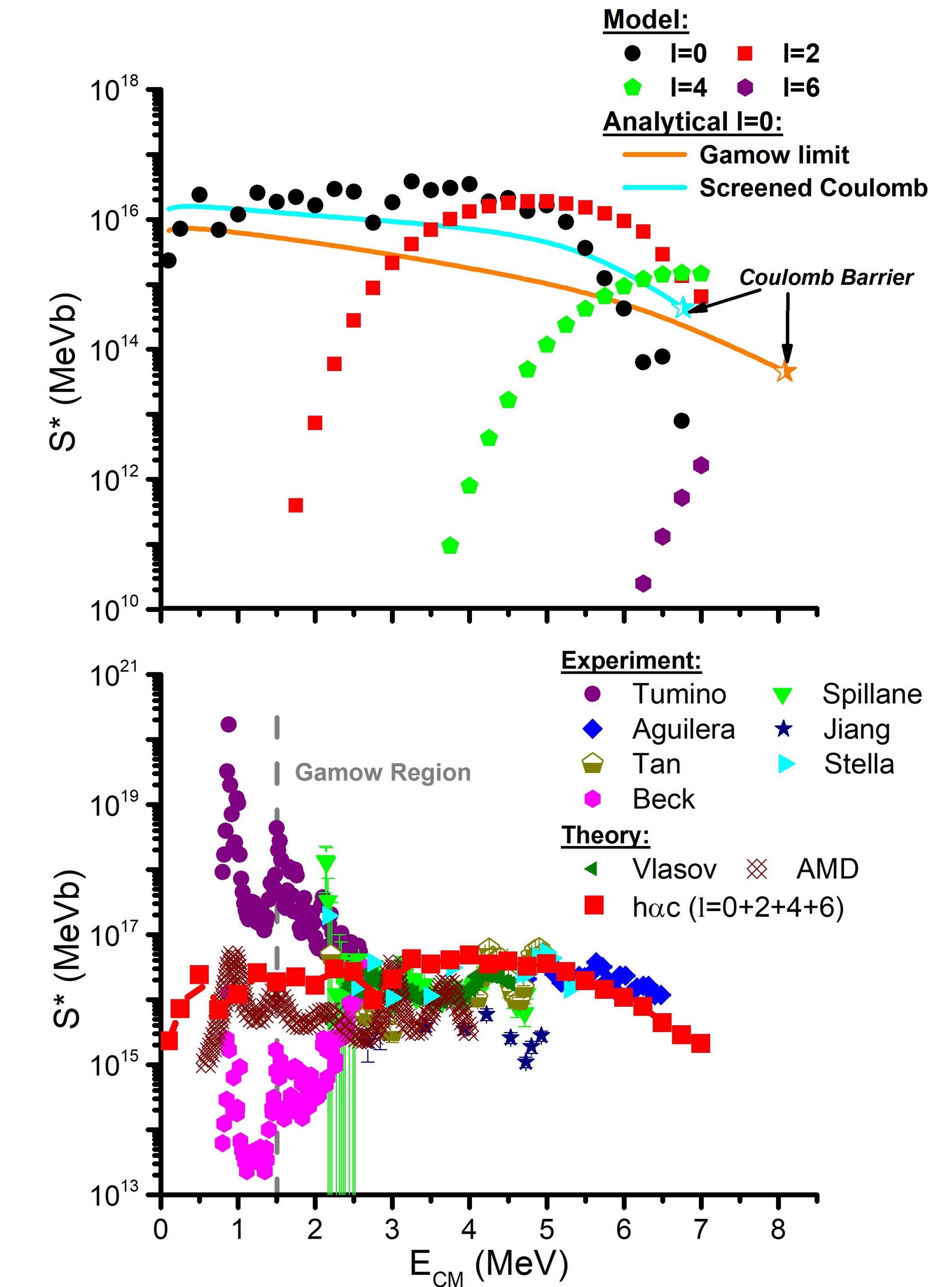
In our model a-degrees of freedom are treated explicitly while nucleon (protons and neutrons) degrees of freedom are treated implicitly hence the hac acronym. The interaction between the a-particles is given by the Coulomb repulsion (in the monopole–monopole approximation for simplicity) and the nuclear attraction. The latter is approximated as  $V_{aa} = V_{Bass}(A = 4)$ , i.e., the Bass potential for mass  $A = 4$  nuclei [30]. Coulomb repulsion is not sufficient to prevent a strong overlap among a-particles. Overlapping nuclei increase the repulsion due to the combined action of the Pauli principle and Heisenberg uncertainty principle, in particular the Fermi energy (per a-particle) is given by:

$$\frac{E_F}{N_\alpha} = 4 \cdot x_F \cdot \bar{\varepsilon}_F \cdot \bar{\rho}^{2/3}.$$



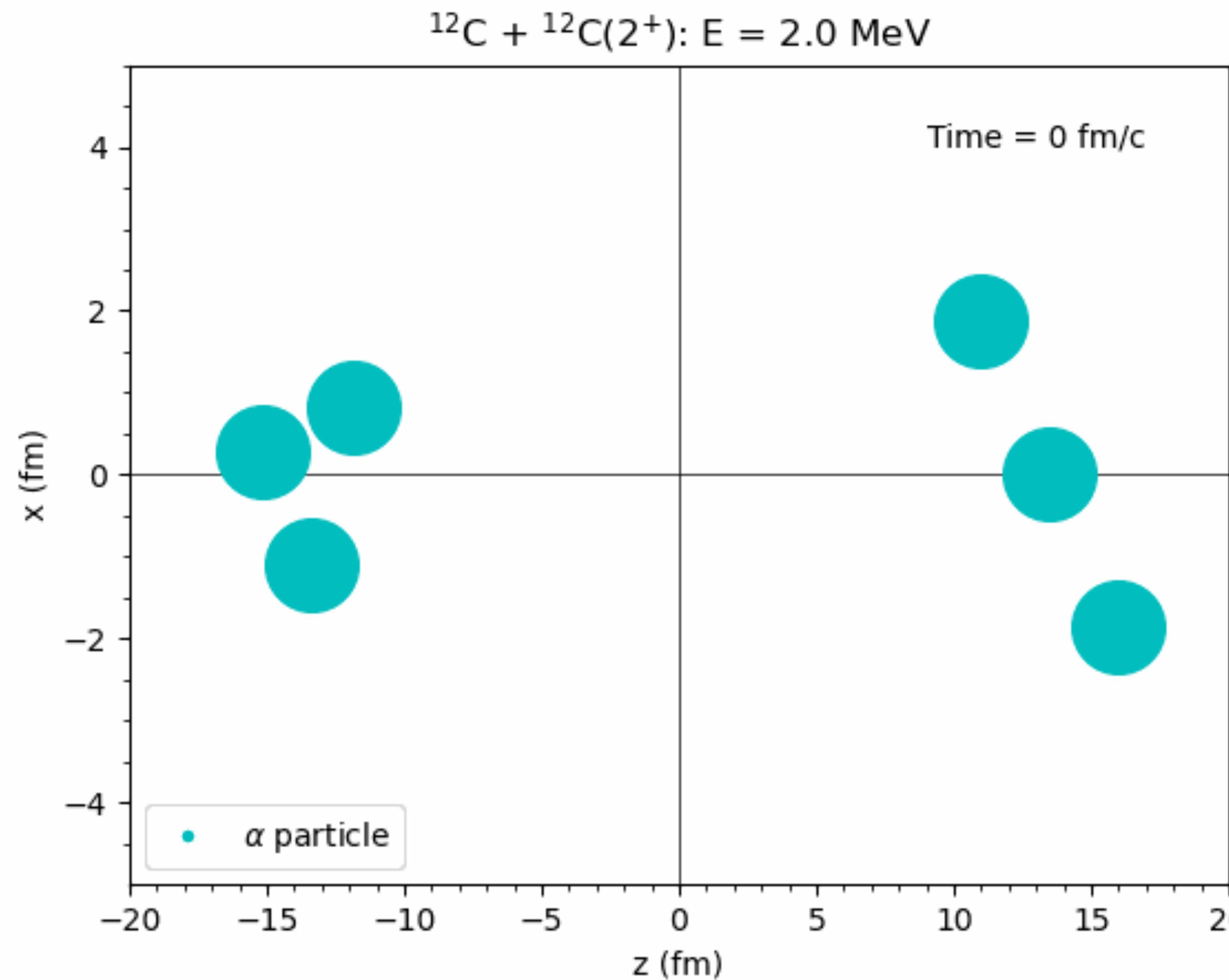


(T.Depastas)



Quare id faciam fortasse requiris!

$^{12}\text{C}$   $E^*=4.44\text{MeV}$   $I=2$

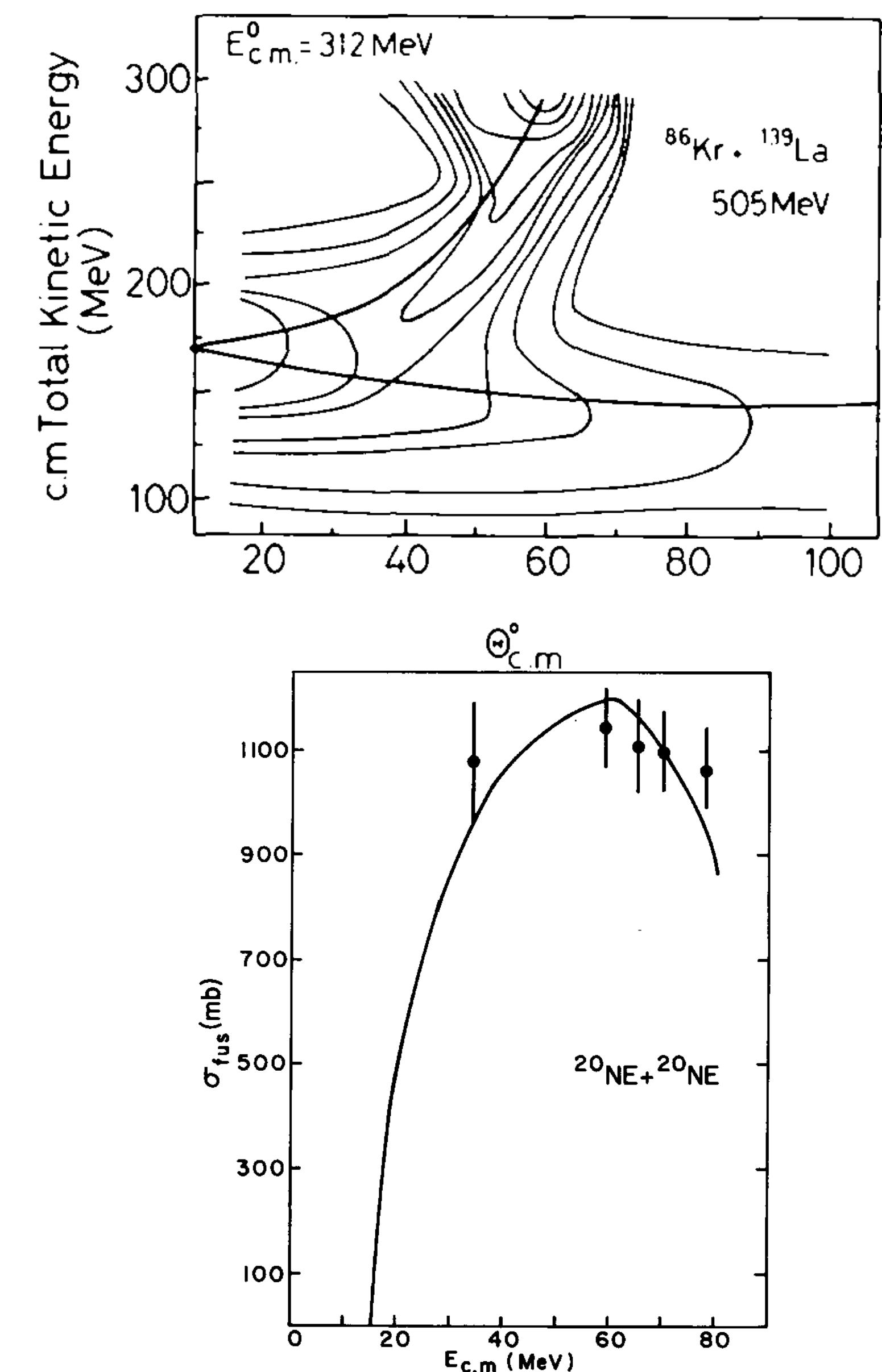
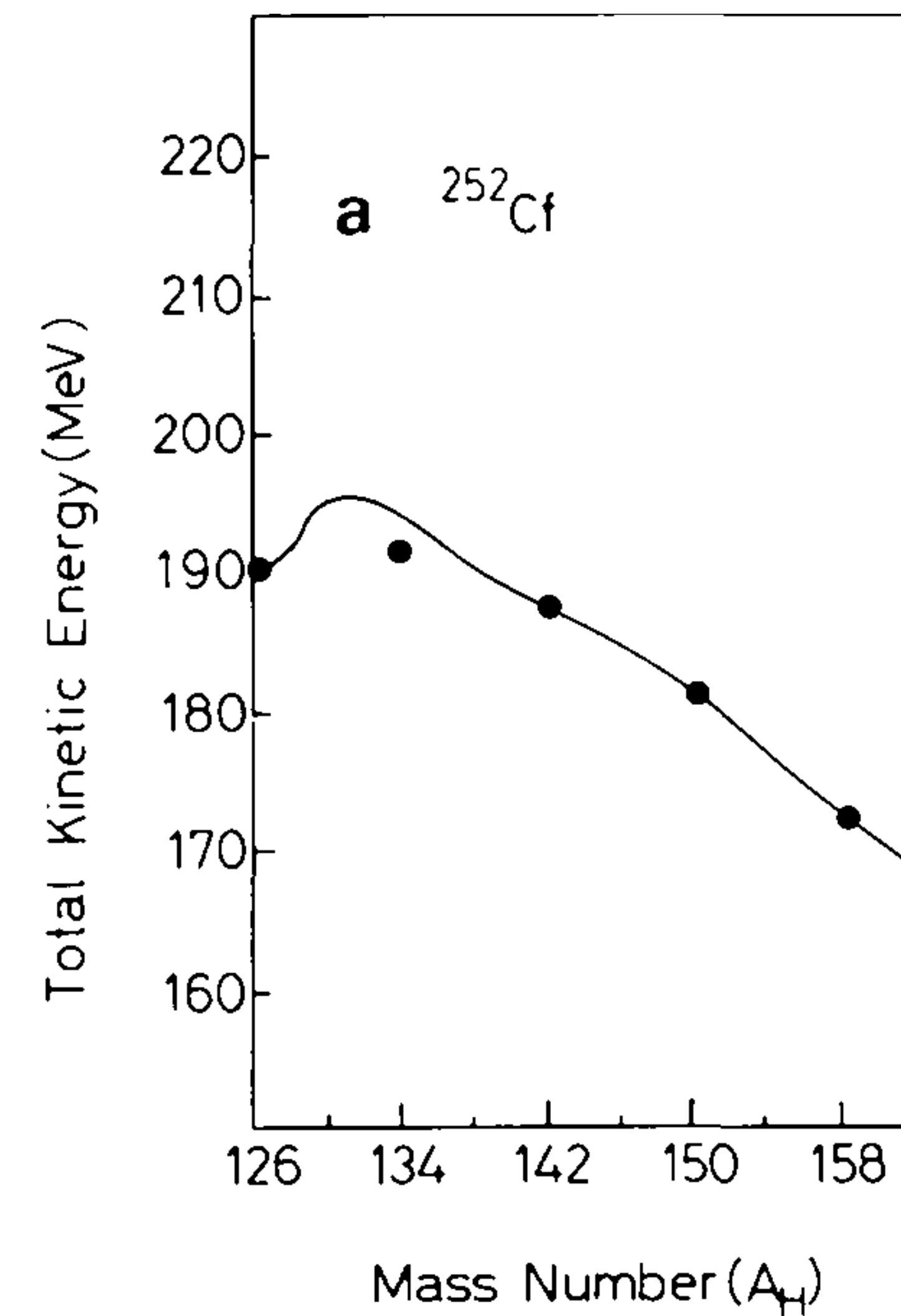
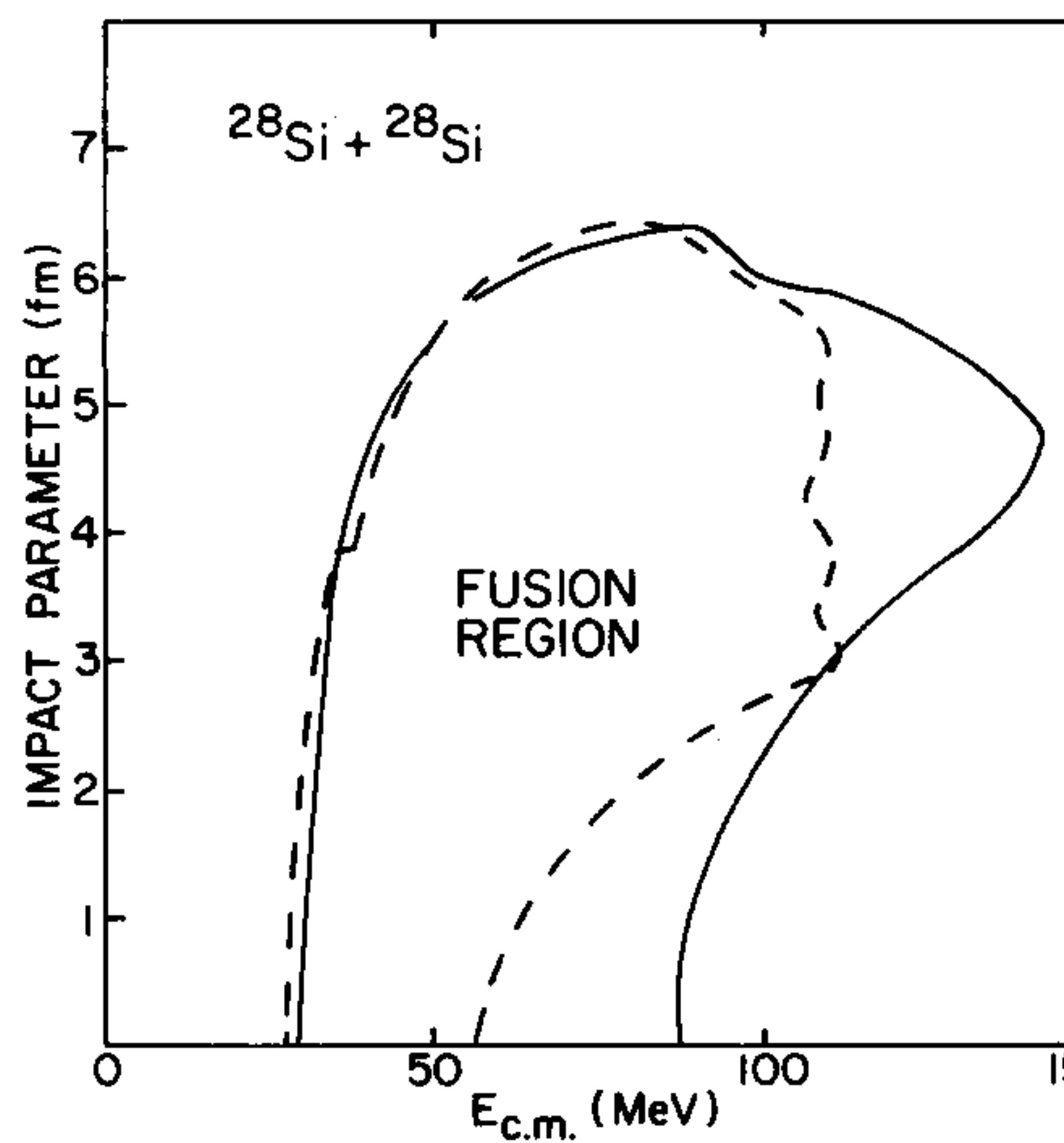


# NEWTONIAN DYNAMICS OF TIME-DEPENDENT MEAN FIELD THEORY

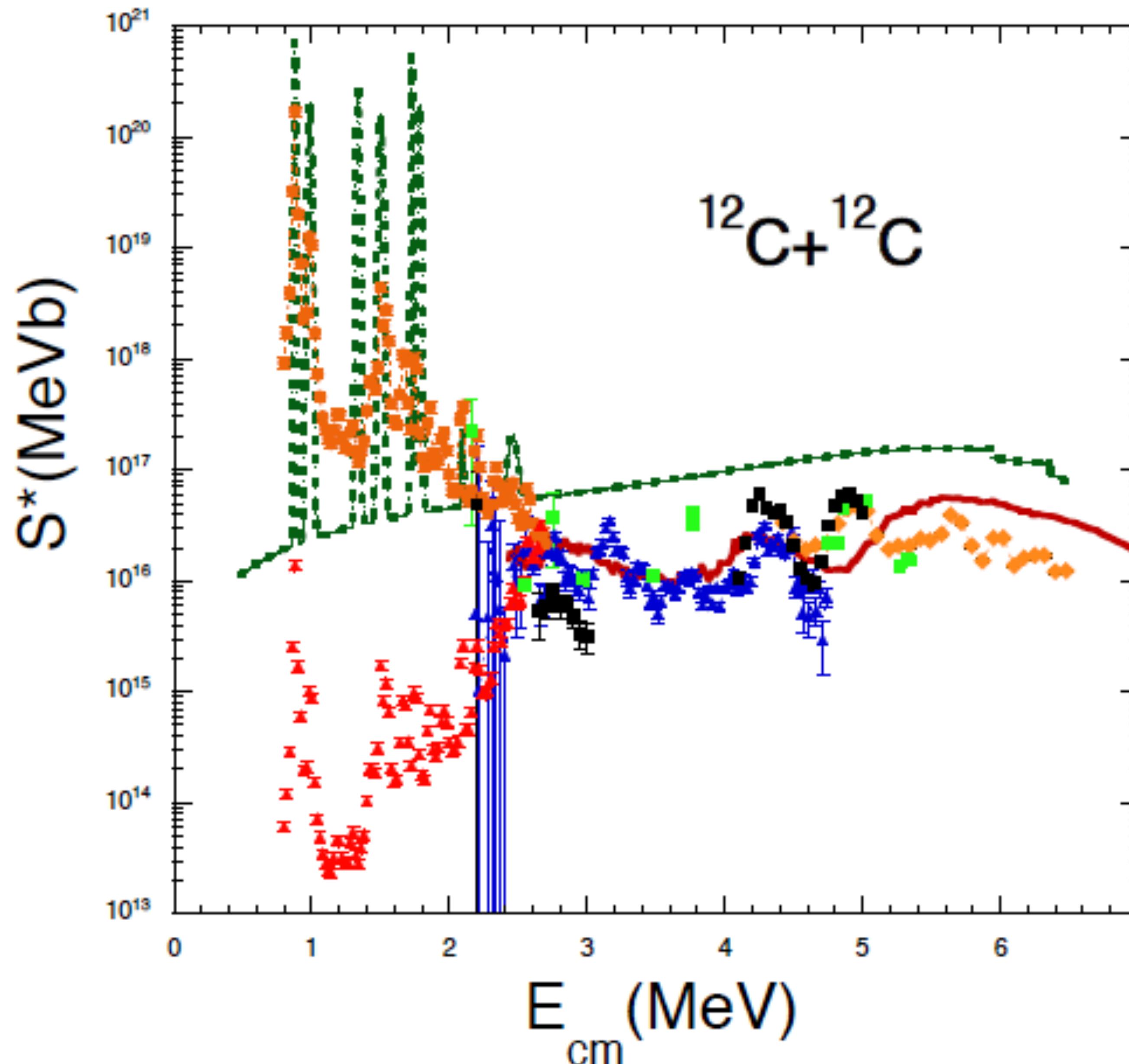
Phys.Lett.B141(1984)9; 168B(1986)35.

A. BONASERA, G.F. BERTSCH and E.N. EL-SAYED

Cyclotron Laboratory, Michigan State University, East Lansing, MI 48824, USA



The probability of fusion for the  $l$  th-partial wave is given by  $Tl = 1/(1 + \exp\{2A\})$ ,  $A = \int_1^2 P \ dR$ .



To take into account resonances modify the Bass potential as:

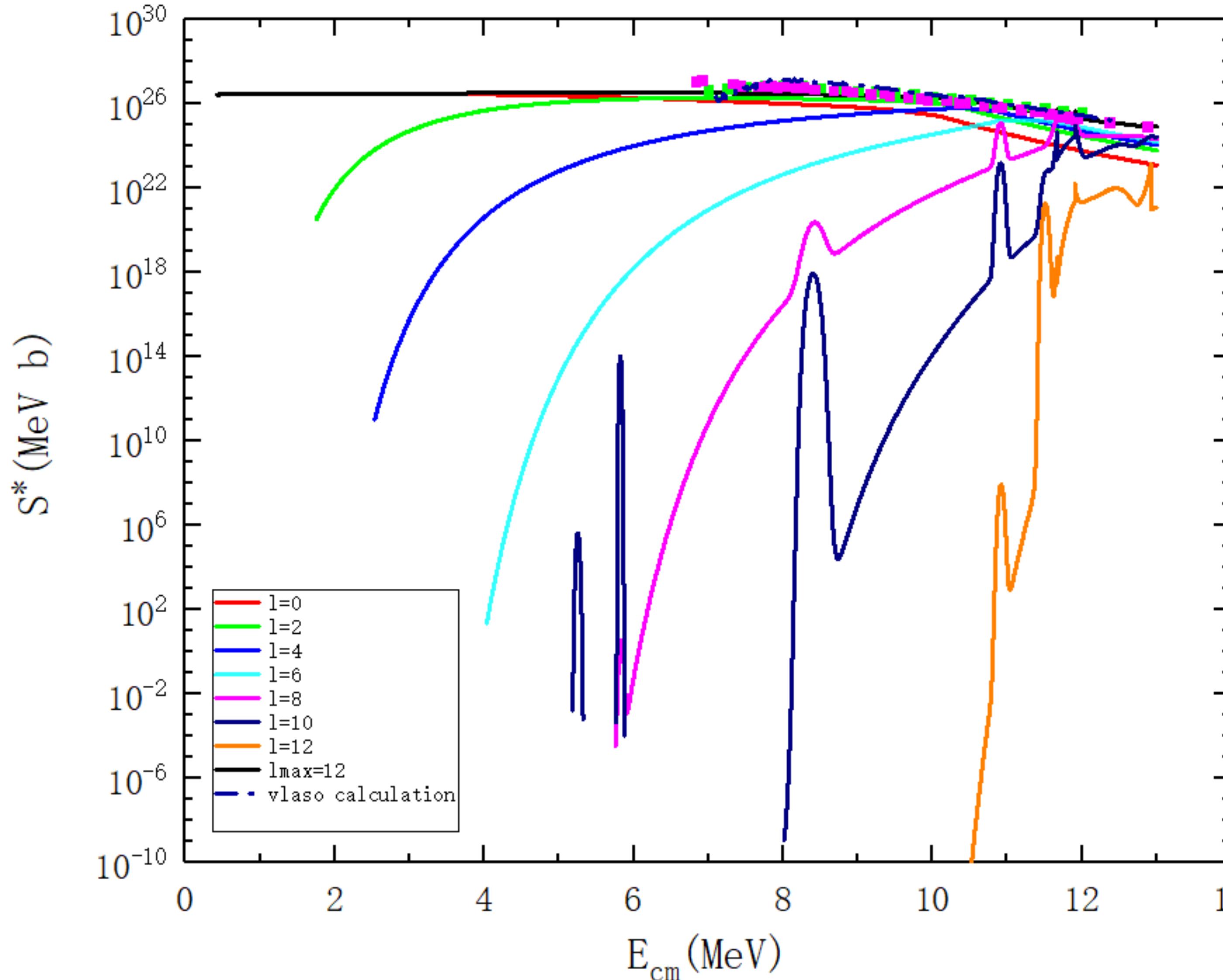
$$V_B \rightarrow V_B[1 + g(x, \gamma, \sigma)],$$

# Calculation of the O16+O16 sub\_barrier fusion cross section

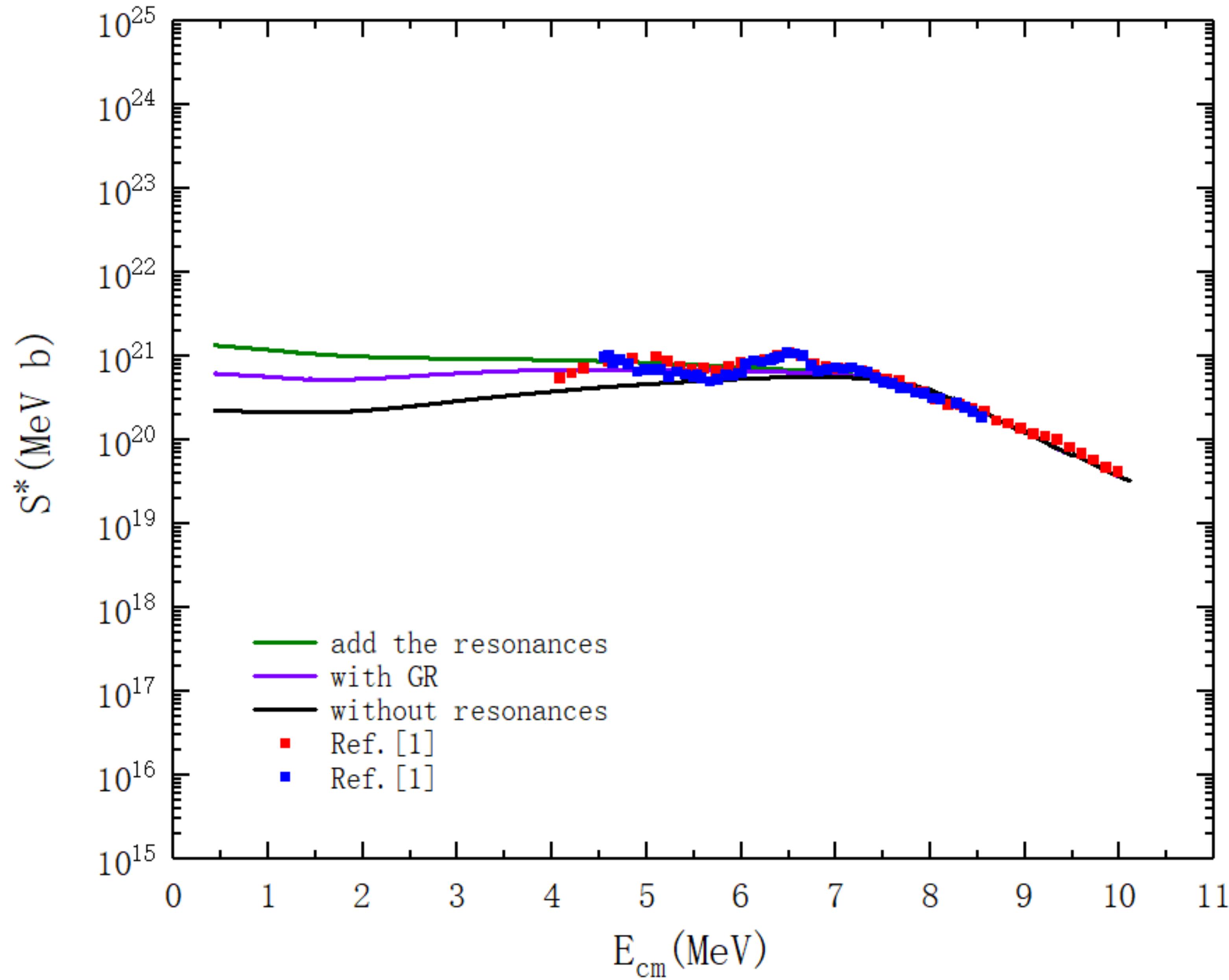
# Shuting Sun

supervisors: H. Zheng and A. Bonasera

School of Physics & Information Technology , Shaanxi Normal University

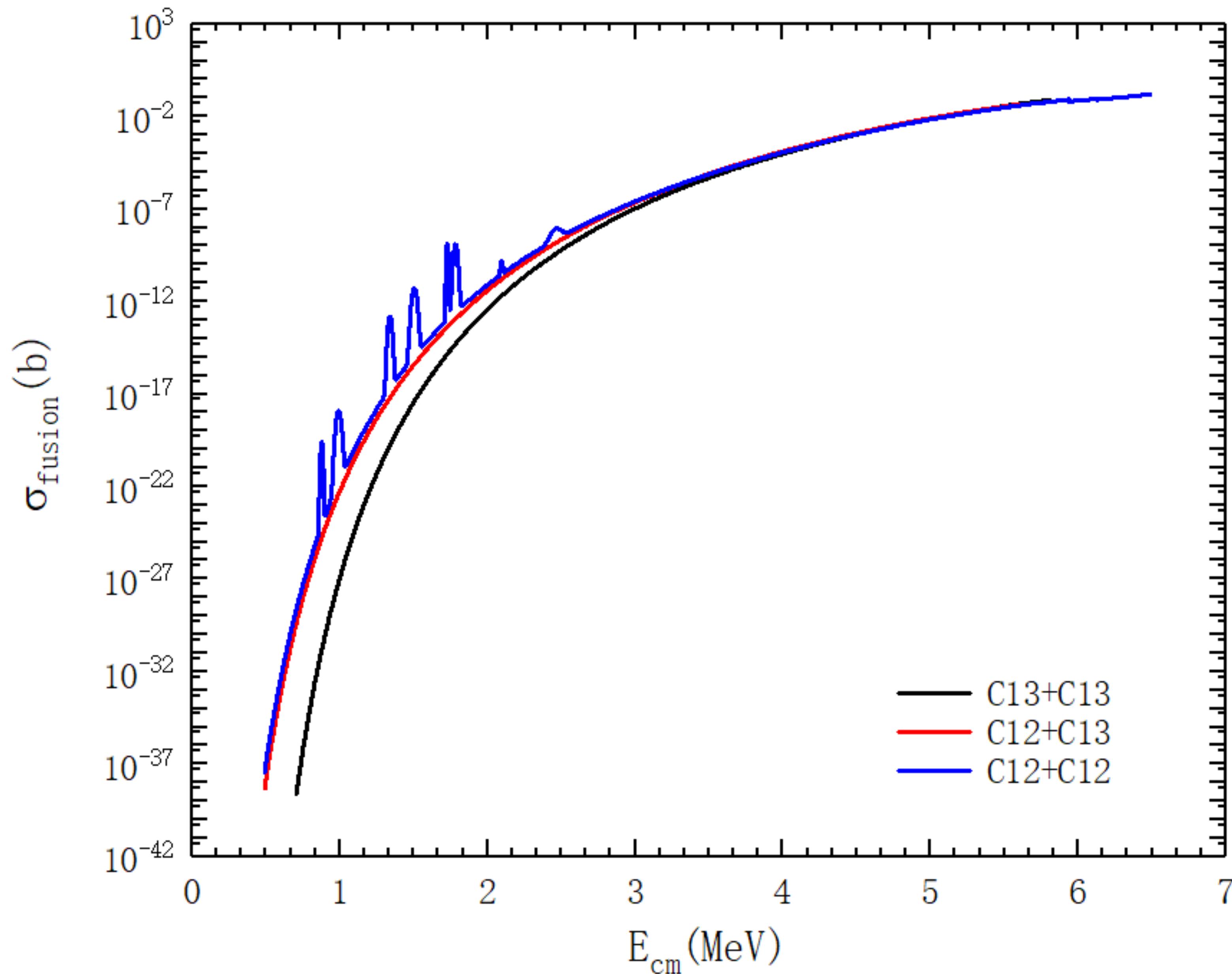


**12C+16O**



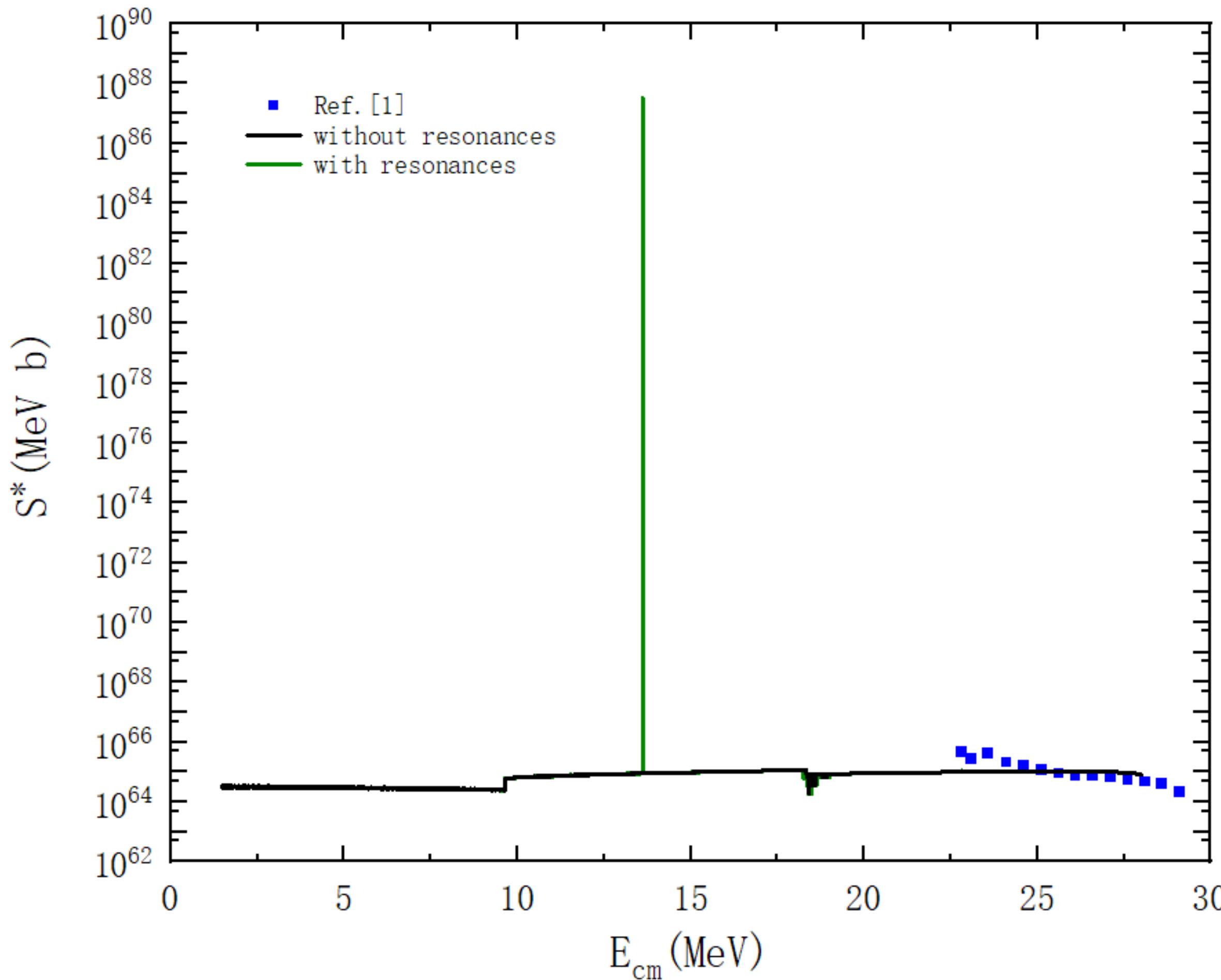
**Shutting Sun**

Because available resonance parameters are not found in nndc, the calculation of C12+C13 and C13+C13 does not include resonance.



**Shuting Sun**

# Calculation of the Si28+ Si28 sub\_barrierfusion cross section



Shuting Sun



National Nuclear Security Administration



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# Progress in Research, Fall 2022

## Dynamical pair production at sub-barrier energies for light nuclei

*Thomas Settemire*, Hua Zheng, and Aldo Bonasera, [arXiv:2207.06900 \[nucl-th\]](https://arxiv.org/abs/2207.06900)



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# Schwinger mechanism

- Dirac particles approximately satisfy the Klein-Gordon Equation

$$(p^2 + m^2) \psi = \left( i\hbar \frac{\partial}{\partial t} - V(x) \right)^2 \psi$$

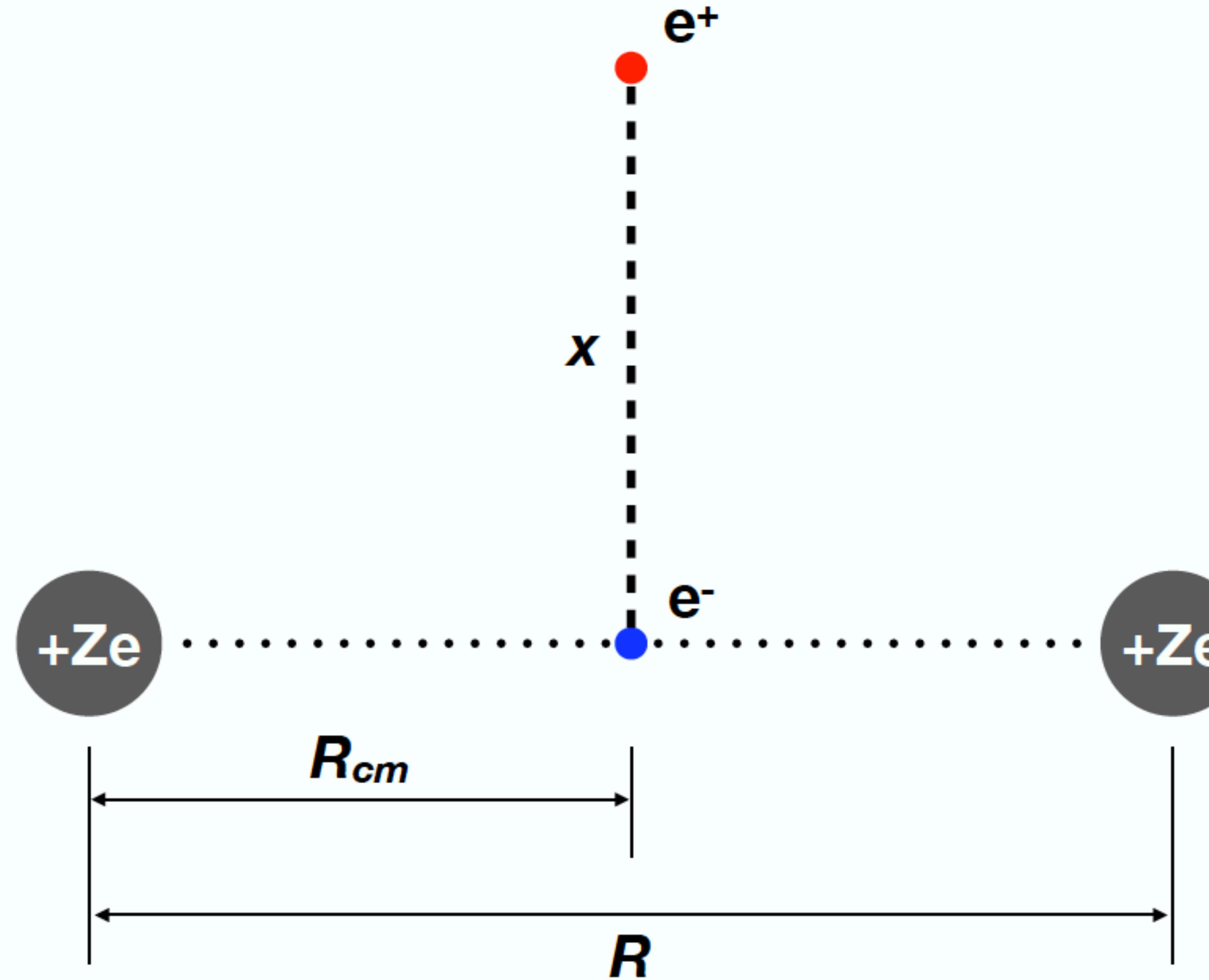
- Manipulate into a Schrödinger equation.

$$\left( \frac{p_x^2}{2m_T} + V_{eff}(x) \right) \psi = 0$$

$$V_{eff}(x) = \frac{m_T}{2} - \frac{(E - V(x))^2}{2m_T} \qquad m_T = \sqrt{m^2 + p_y^2 + p_z^2}$$

- Negative energy particles in Dirac sea can tunnel through the effective potential and become real.

- The two nuclei come together with impact parameter zero.
- Suppose the  $e^-$  is created at the center of mass of the two nuclei and the  $e^+$  is on an axis perpendicular to the beam axis.
- Symmetric and energetically favorable.
- The ions get accelerated by the  $e^-$  in the middle, encouraging fusion.



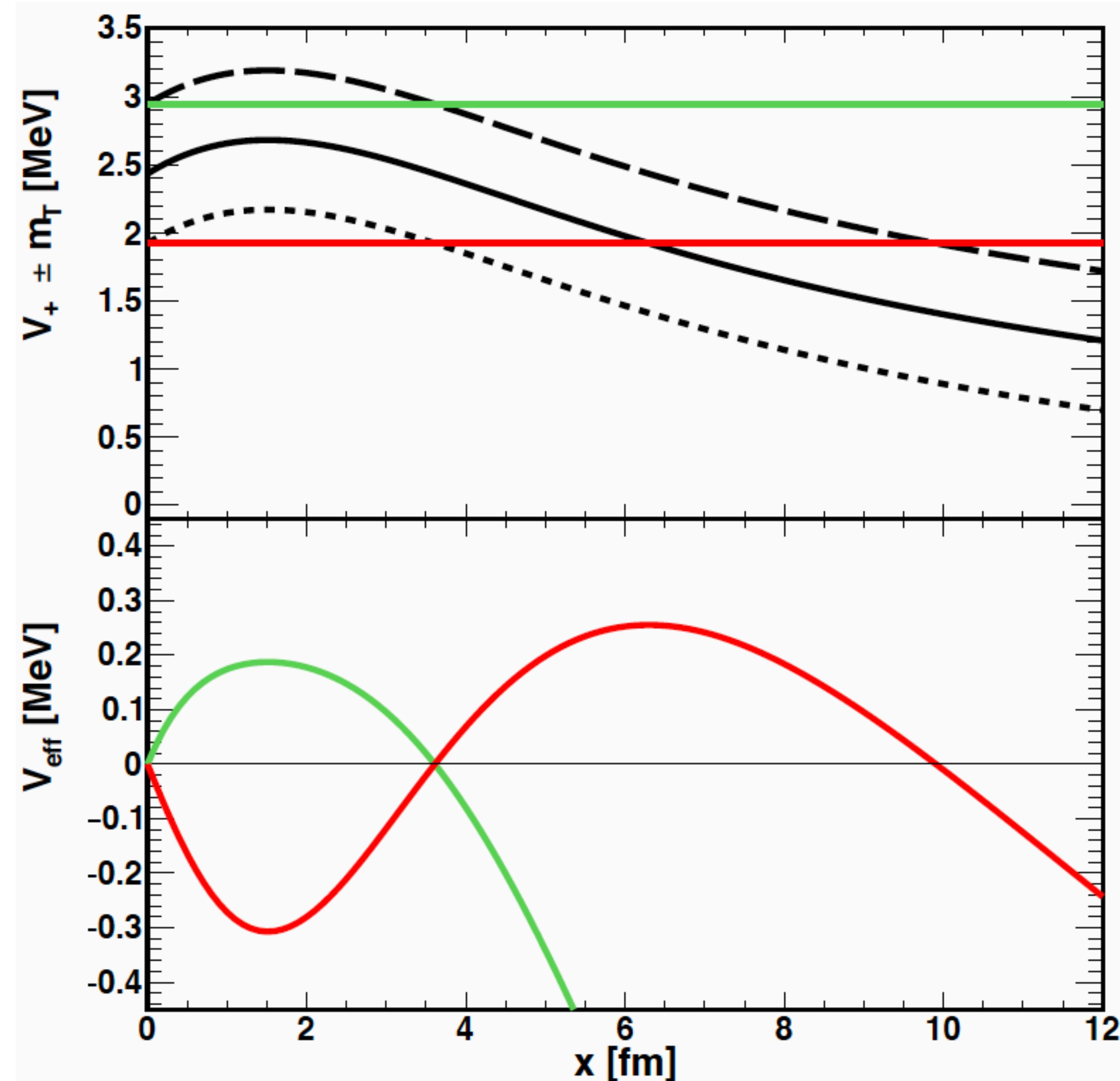
# Effective potential

- Potential energy of positron

$$V_+(R, x) = \frac{2Ze^2}{\sqrt{\left(\frac{R}{2}\right)^2 + x^2}} - S(x) \frac{e^2}{x},$$

- Effective potential

$$V_{eff}(x) = \frac{m_T}{2} - \frac{(E_+ - V_+(R, x))^2}{2m_T},$$

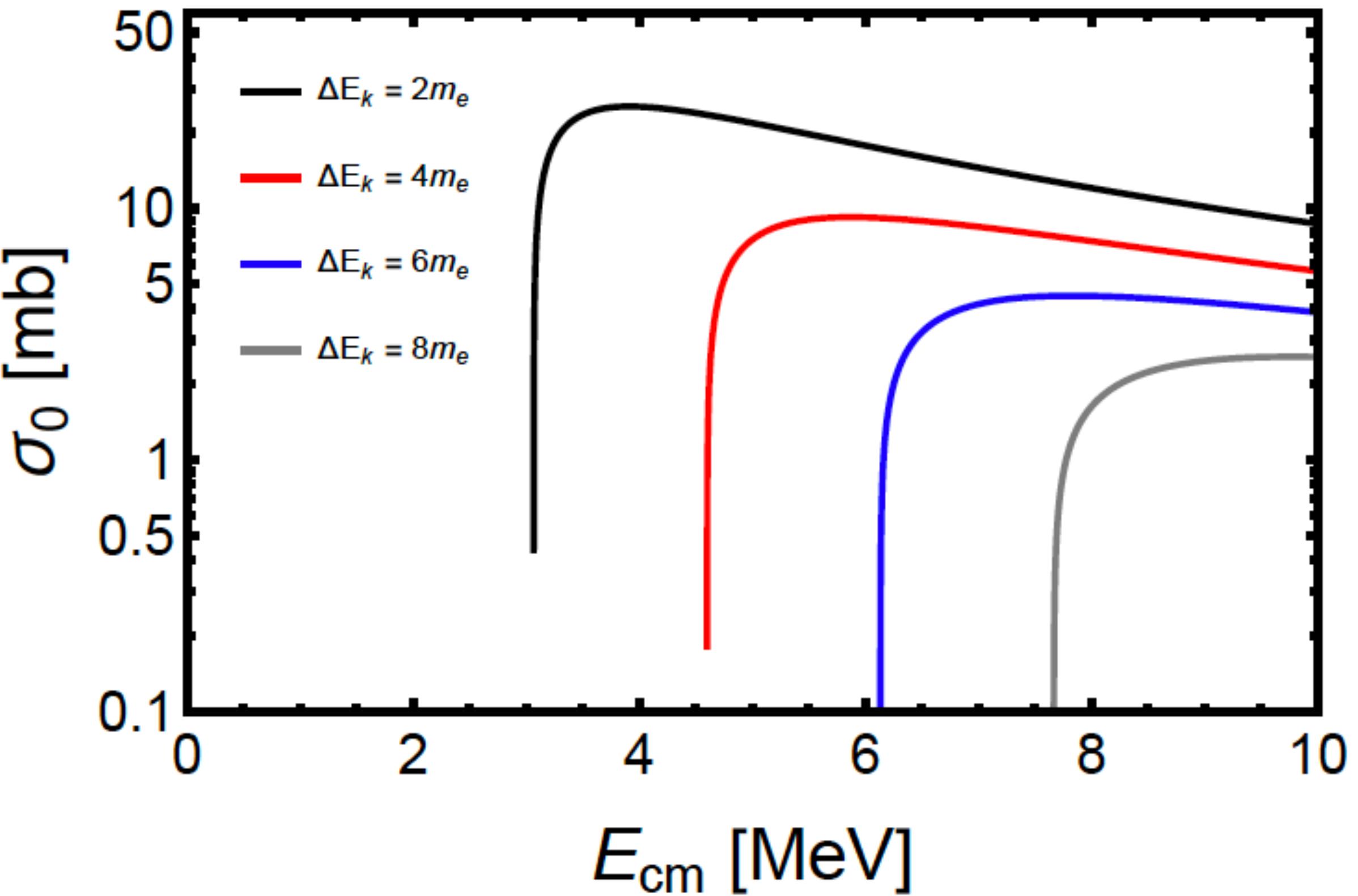


# Results

$$\sigma(E_{c.m.}) = \frac{\pi\hbar^2}{2\mu E_{c.m.}} \sum_{l=0}^n (2l+1) \Pi_l P_H.$$

$$P_H = \tau/\Delta\tau, \quad \Delta\tau = \frac{\hbar}{2m_T},$$

- Square root in  $\sigma$  gives minimum value for  $E_{cm}$



$$\begin{aligned} \sigma_0(E_{c.m.}) &= \frac{\pi\hbar^2}{2\mu E_{c.m.}} 0.5 \frac{\tau}{\Delta\tau} \\ &= \frac{1}{N_{\pm}^{max}} \frac{6\pi\hbar Z e^2 m_e}{\sqrt{2\mu} E_{c.m.}^2 (\Delta E_k + 2m_e)} \\ &\times \sqrt{E_{c.m.} - \frac{Z}{4}(\Delta E_k + 2m_e)}. \end{aligned}$$

# Pair production as a probe for the dynamics of nuclear fission and $\alpha$ decay

T. Settemyre,<sup>1</sup> H. Zheng,<sup>2</sup> and A. Bonasera<sup>1,3</sup>

<sup>1</sup>*Cyclotron Institute, Texas A&M University, College Station, Texas 77843, USA*

<sup>2</sup>*School of Physics and Information Technology, Shaanxi Normal University, Xi'an 710119, China*

<sup>3</sup>*Laboratori Nazionali del Sud, INFN, via Santa Sofia, 62, 95123 Catania, Italy*

Electron-positron pairs can be produced via the Schwinger mechanism in the presence of strong electric fields. In particular, the fields involved in  $\alpha$  decay and nuclear fission are strong enough to produce them. The energy of the  $e^+e^-$  pair is related to the relative distance and velocity of the daughter nuclei. Thus, the energy distribution of the produced pairs can give information about the dynamics of the fission and  $\alpha$  decay processes. A neck model of nuclear fission is used to illustrate how the pairs can be used as a probe of the dynamics.

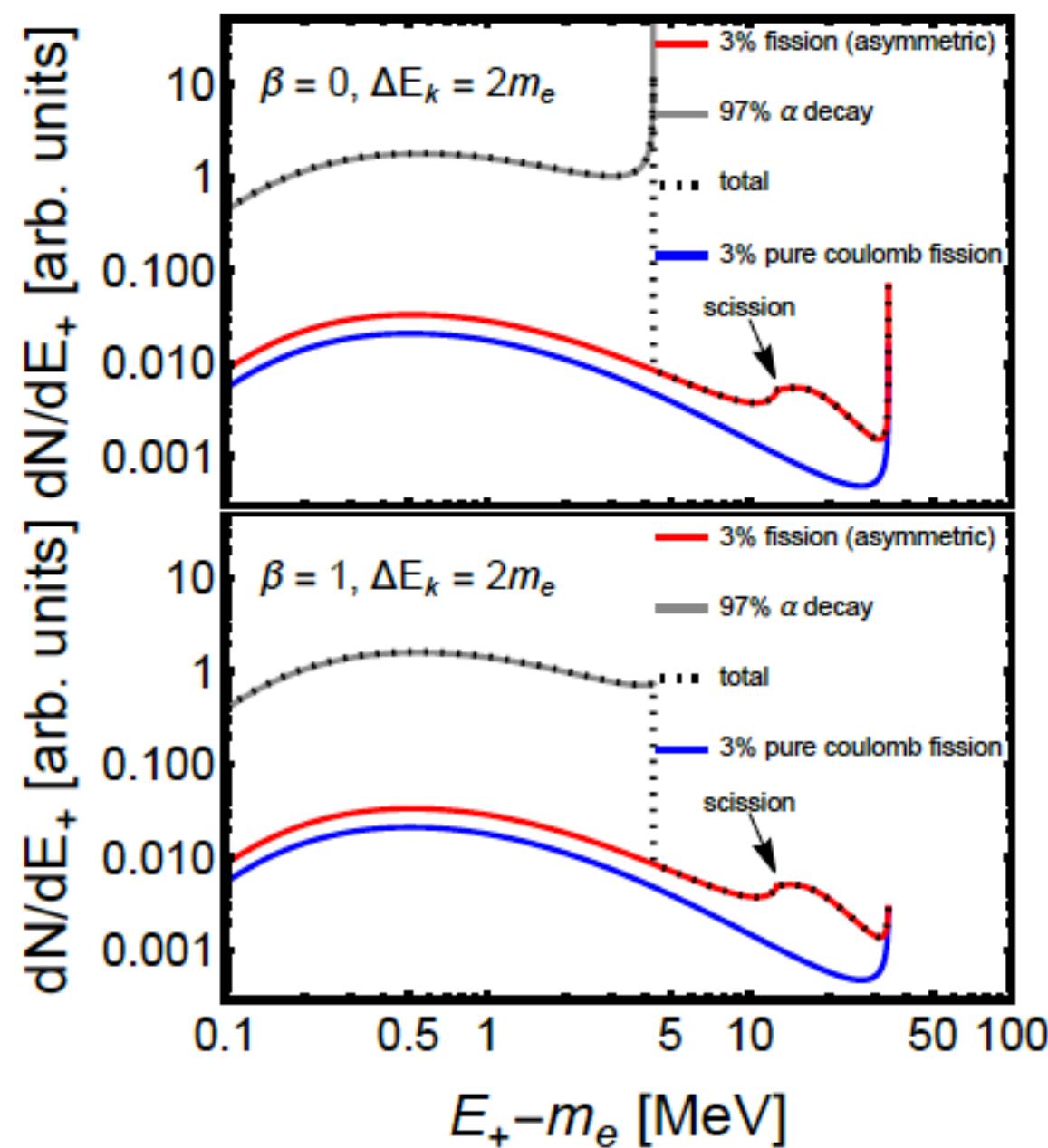


FIG. 3:  $dN/dE_+$  for fission and  $\alpha$  decay of  $^{252}\text{Cf}$ , weighted by the corresponding branching ratios. Here,  $m_T = m_e$  and  $\Delta E_k = 2m_e$ . In the top panel we assume that the extra acceleration ( $\Delta E_k$ ) is given to the ions after the pair creation resulting in the divergence at time zero,  $E_+ \approx 30 \text{ MeV}$ ,  $\beta = 0$ . The Coulomb divergence disappears if  $\beta = 1$ , i.e. the ions are accelerated at the beginning of the pair production process (bottom panel). Similarly for  $\alpha$  decay.

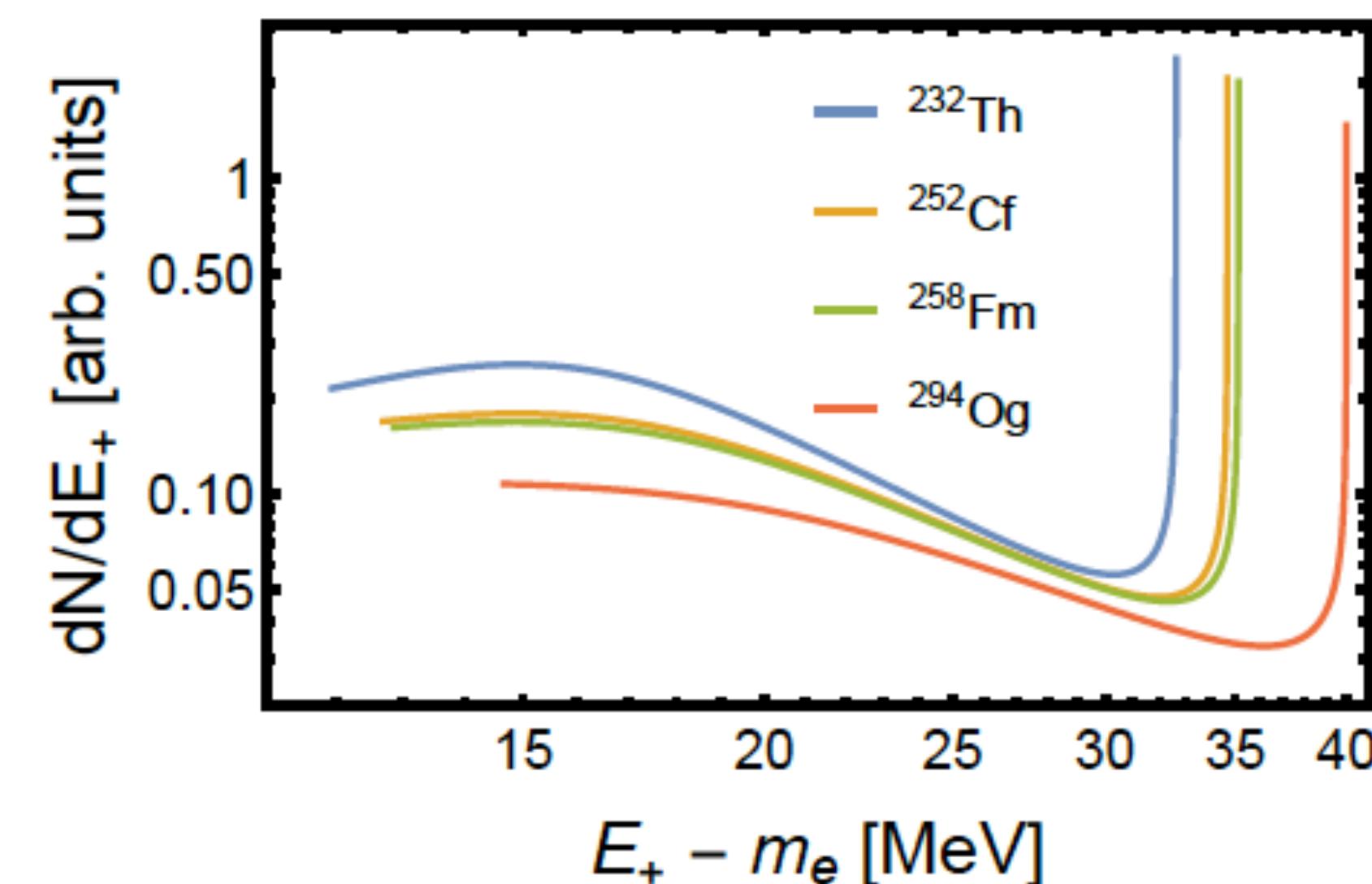


FIG. 4:  $dN/dE_+$  for fission of a selection of nuclei. The fission dynamics are based on the neck model. Here,  $m_T = m_e$ ,  $\beta = 0$  and  $\Delta E_k = 2m_e$ .

# Conclusions

$^{12}\text{C}+^{12}\text{C}$ : The Neck model and the Vlasov approach in imaginary time give  $S^*>e16\text{MeVb}$  for  $E_{\text{cm}}>0.5\text{ MeV}$

Adding resonances is in some agreement with the THM  
 $I=0$  channel is dominant up to  $E_{\text{cm}}=3\text{MeV}$

For heavier systems resonances not so important because of the large Q-value

Microscopic models in progress and some interesting results have been obtained for fusions in C stars.

Pair production in heavy and superheavy nuclei plus other systems

Applications to superheavy formation

THANKS



DOE grant: DE-FG03-93ER40773