

Photon Strength Function of ^{58}Fe using the Oslo and Shape Methods

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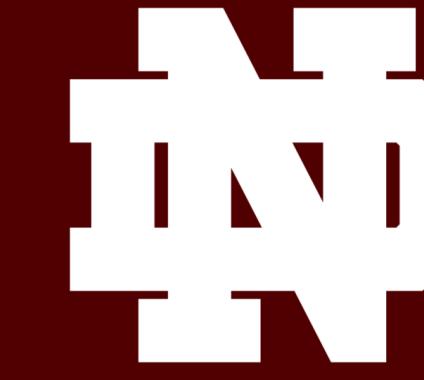
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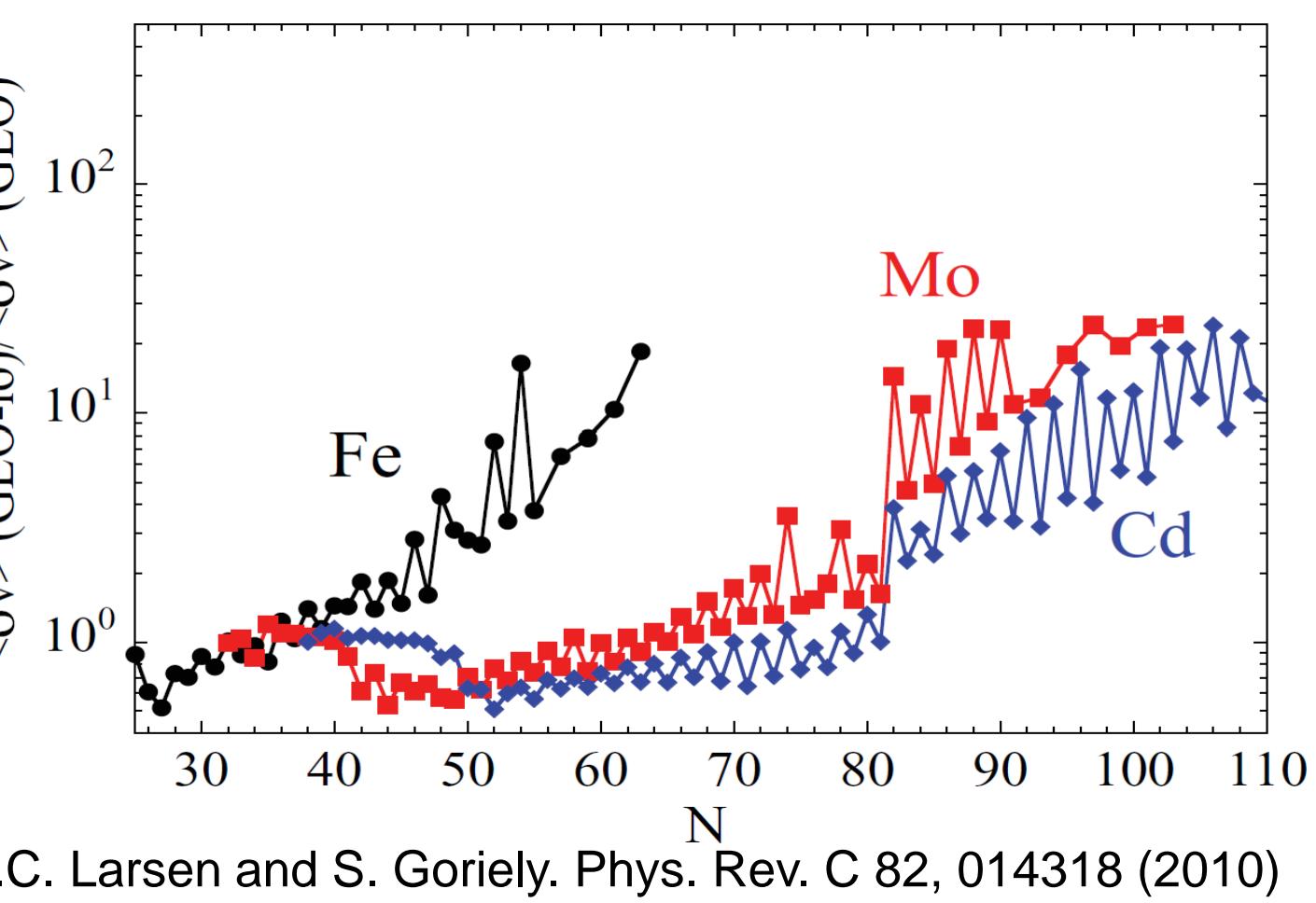
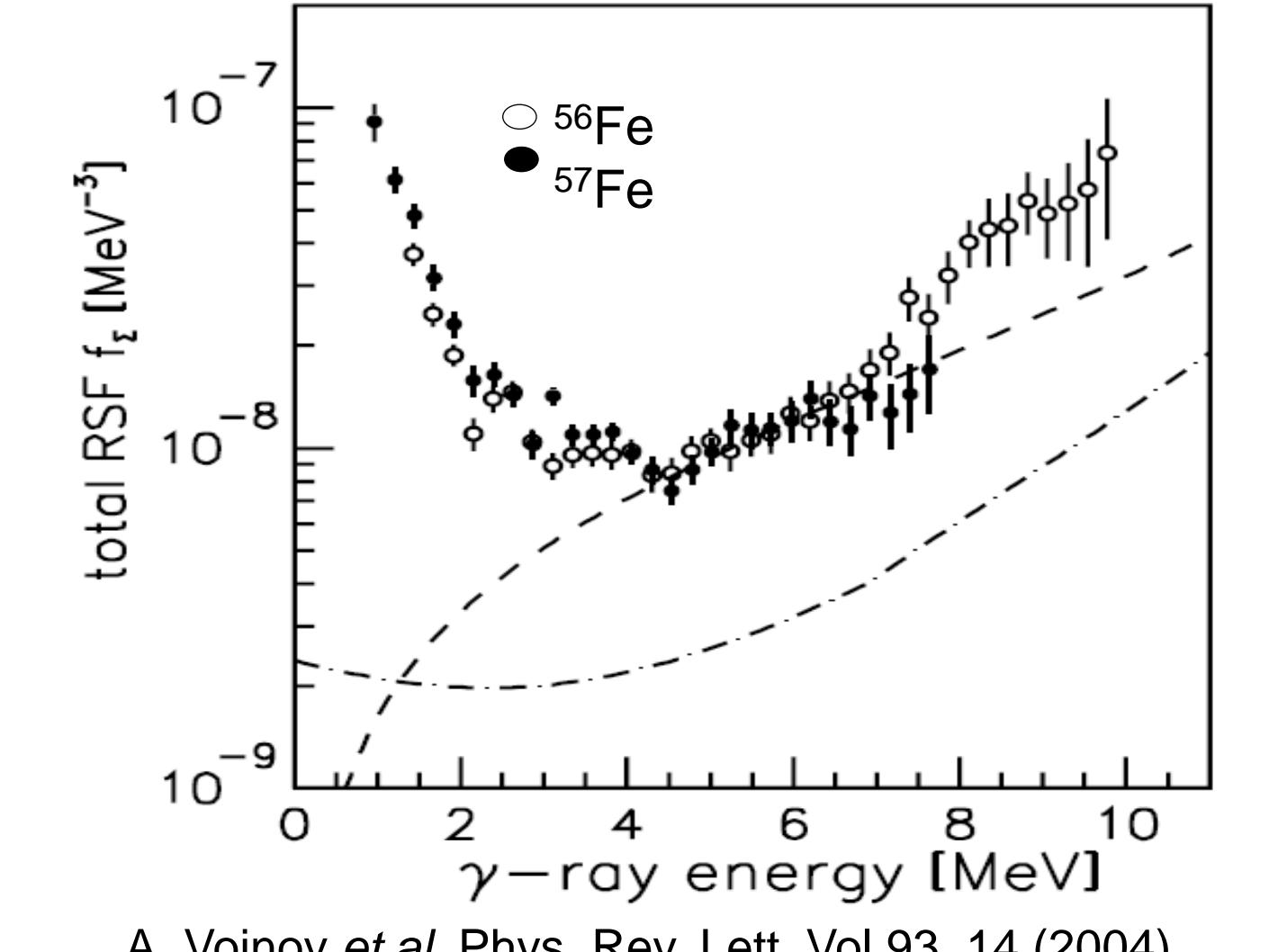
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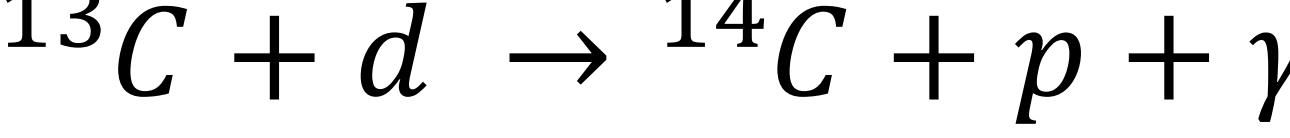
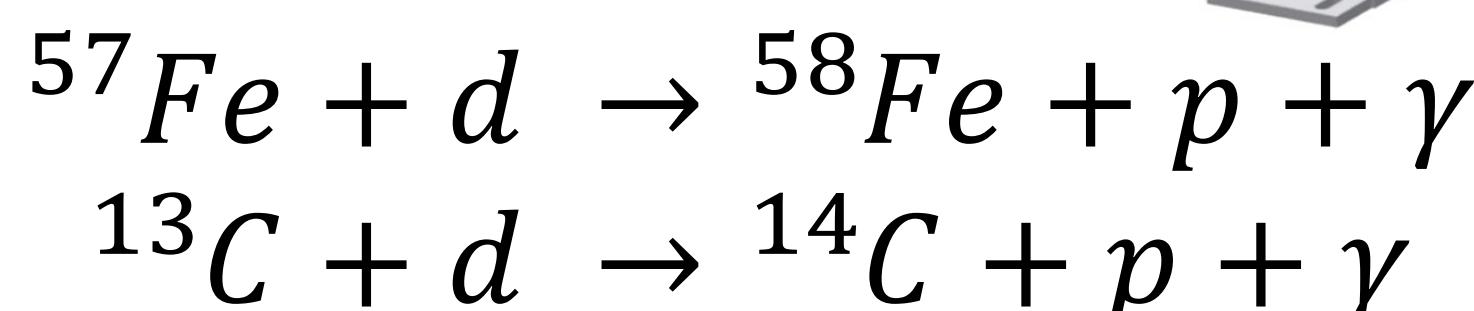
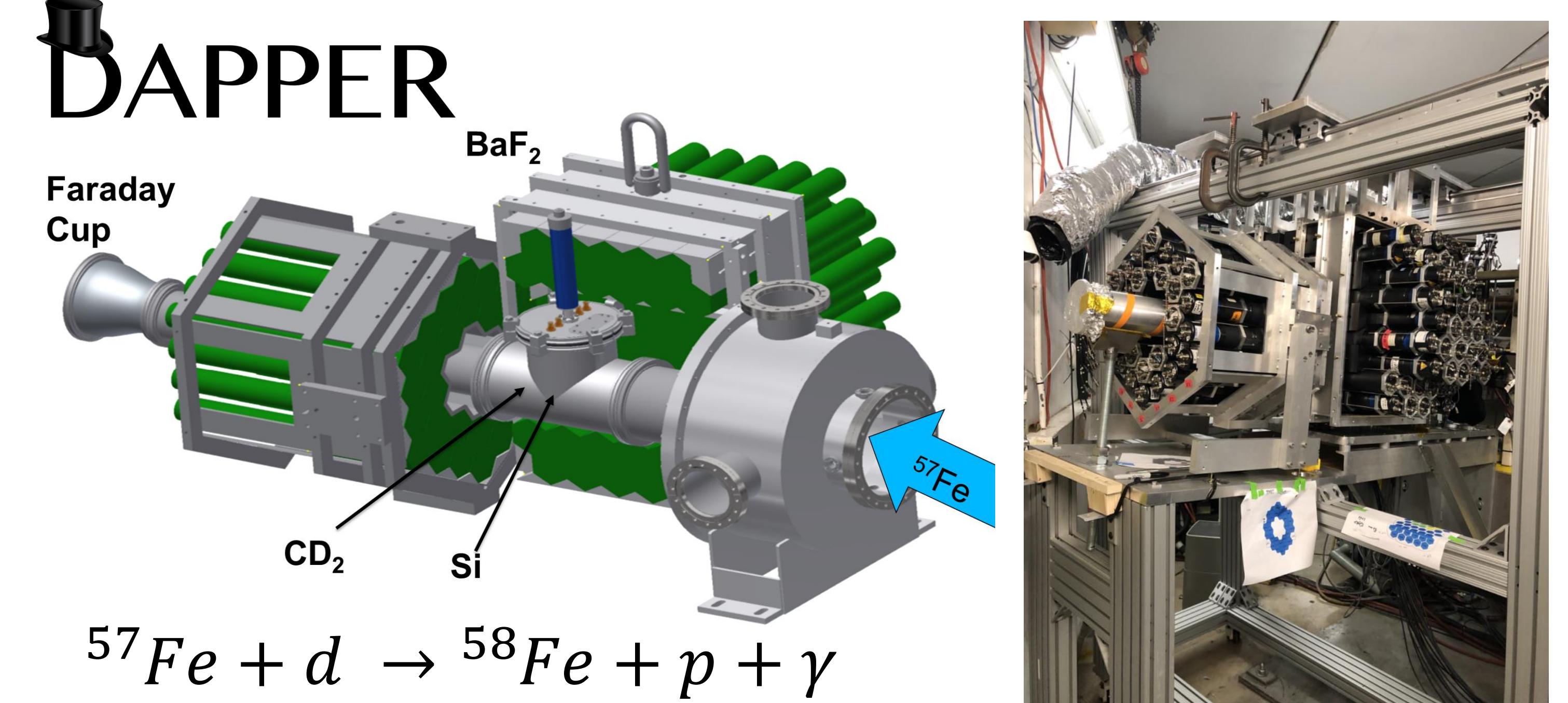
Introduction

- Neutron capture processes dominate production of $Z > 26$
- Models of stellar evolution require accurate cross-sections
- Photon strength functions often have the dominant uncertainty
- Low energy enhancement (LEE) first observed
- Reaction rates in r-process nucleosynthesis
- ^{60}Fe measurements on earth and space



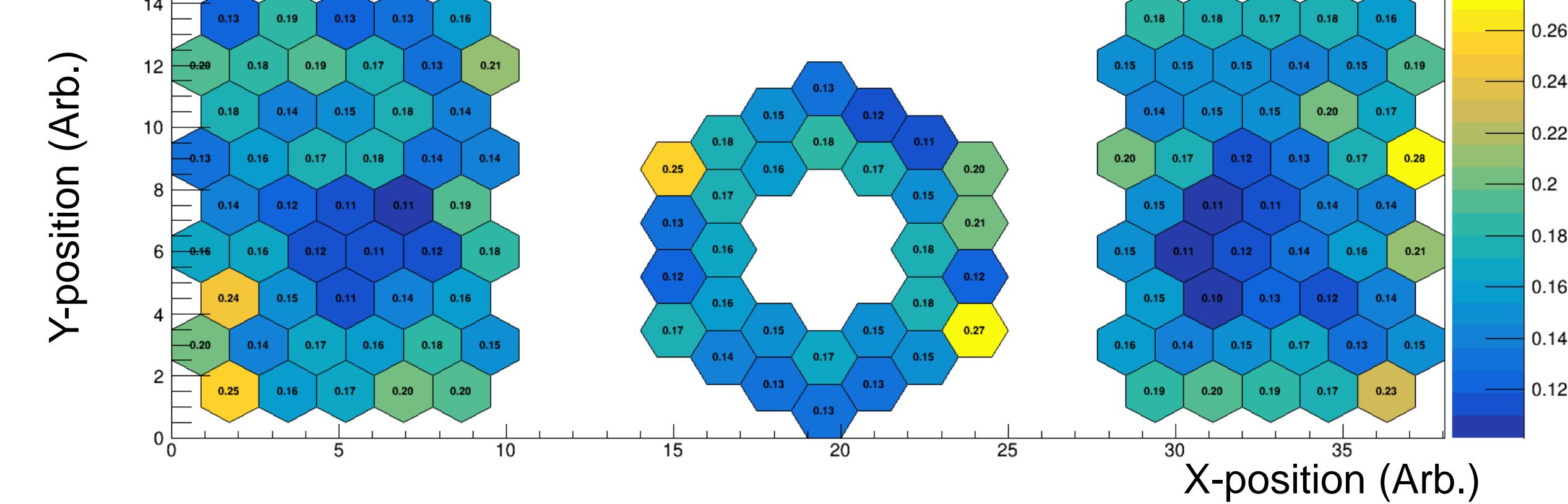
A. Voinov et al. Phys. Rev. Lett. Vol.93, 14 (2004)

A.C. Larsen and S. Goriely. Phys. Rev. C 82, 014318 (2010)



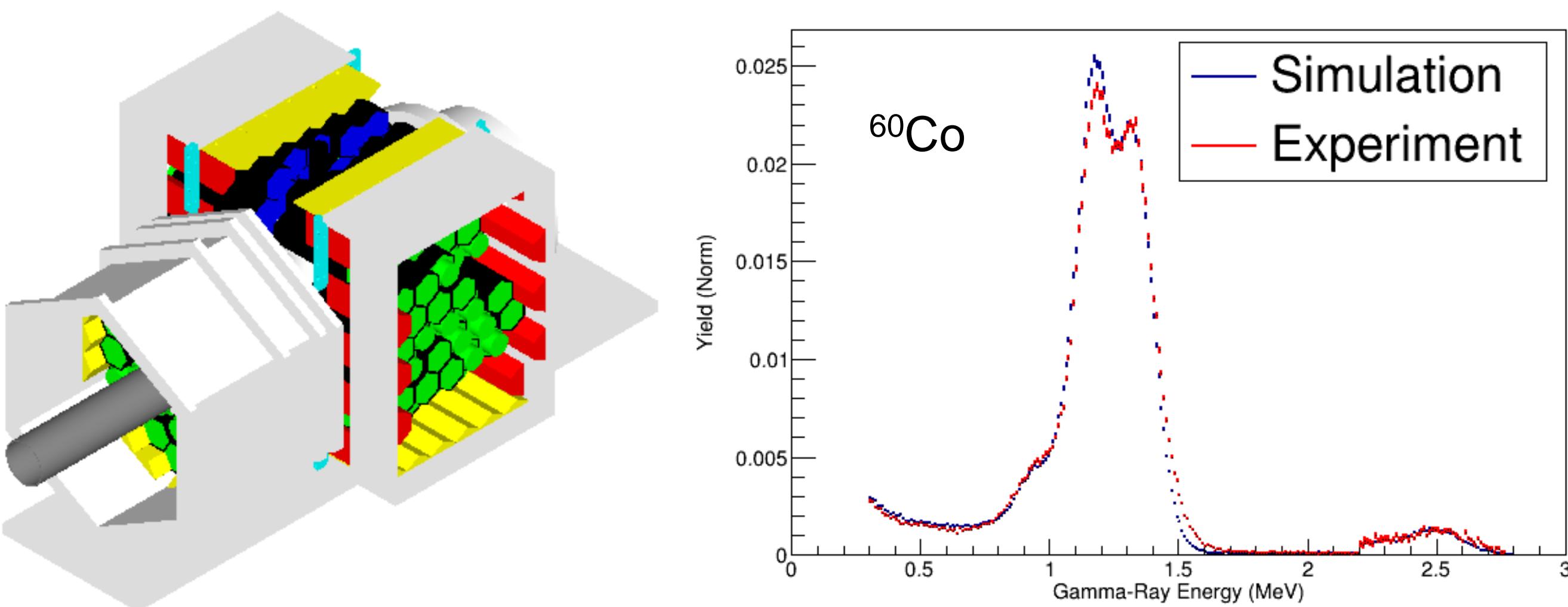
- Obtain PSF experimentally by measuring gamma-ray energies emitted from particular initial excited states
- Inverse kinematics opens possibilities of studying rare isotopes

Relative Energy Res (FWHM) at 1 MeV



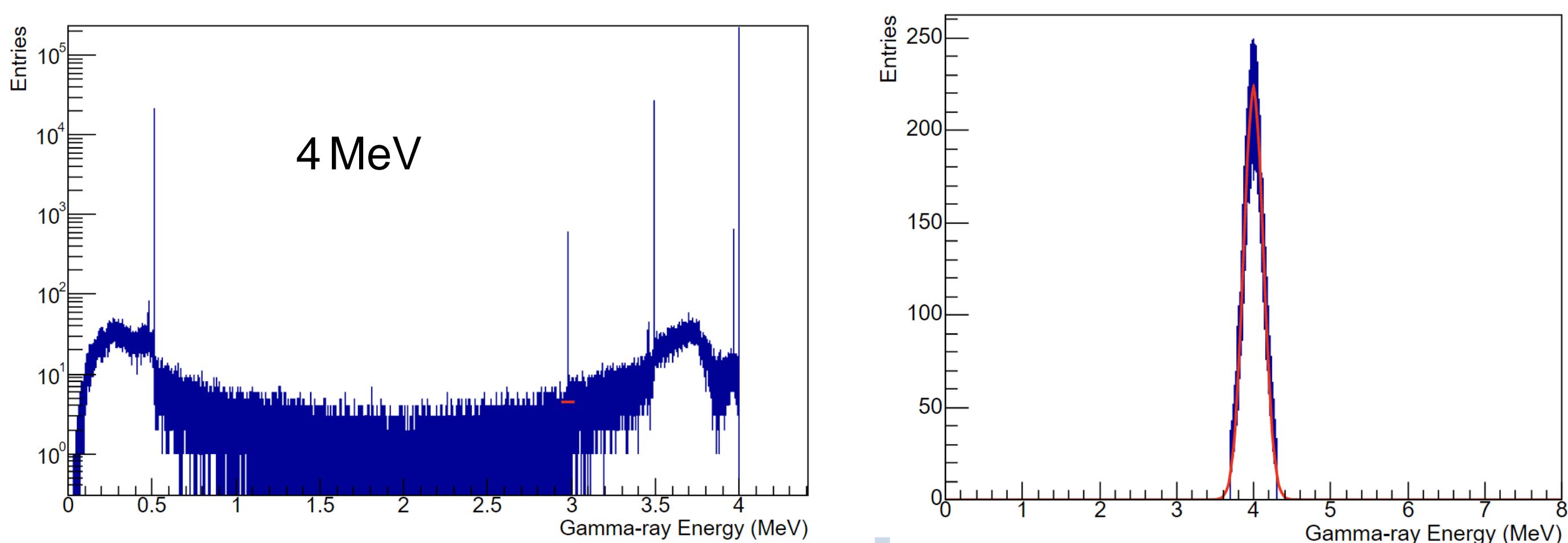
Simulation

- GEANT4 simulation of DAPPER
- Accuracy tested with source data

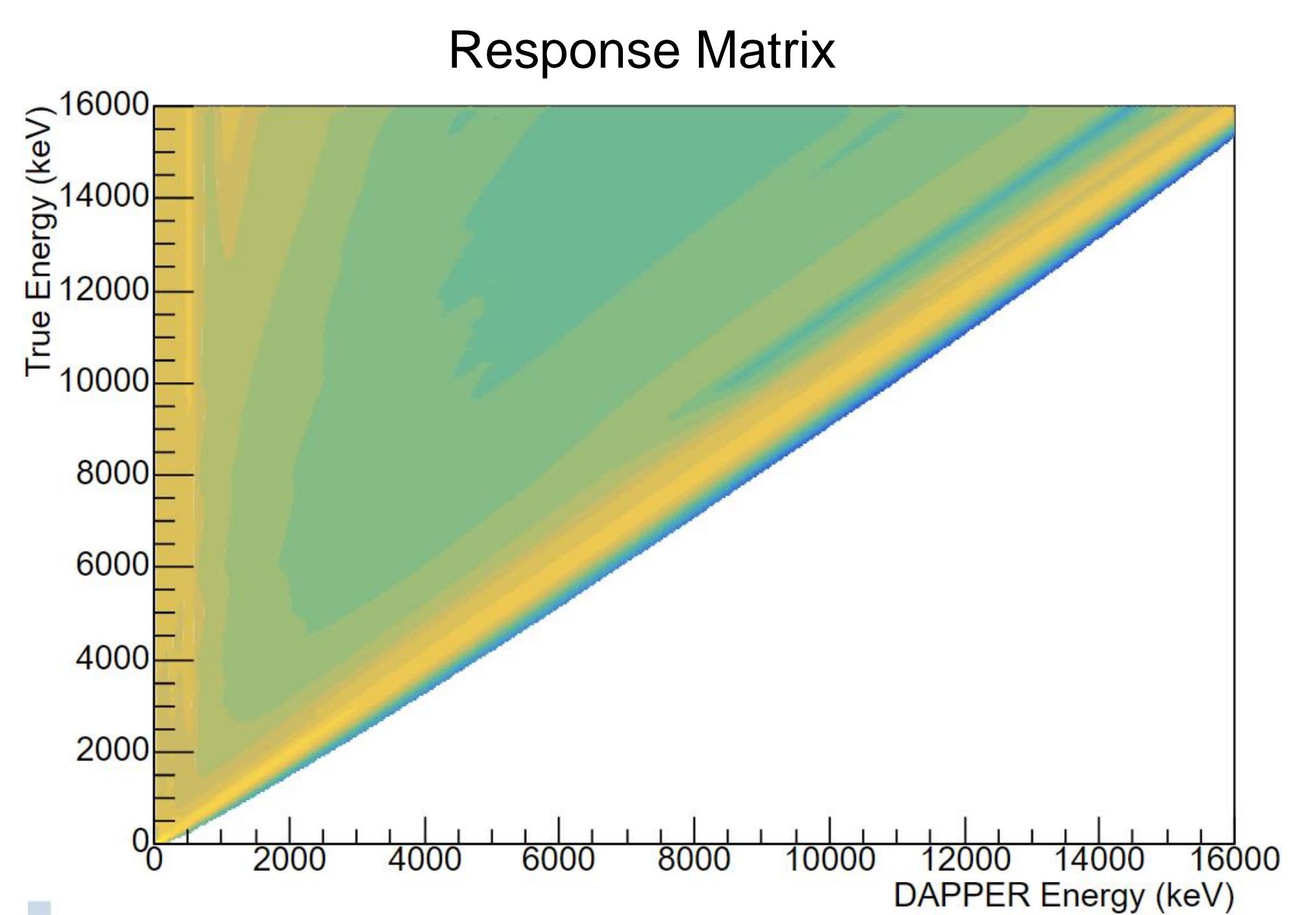


Response Function

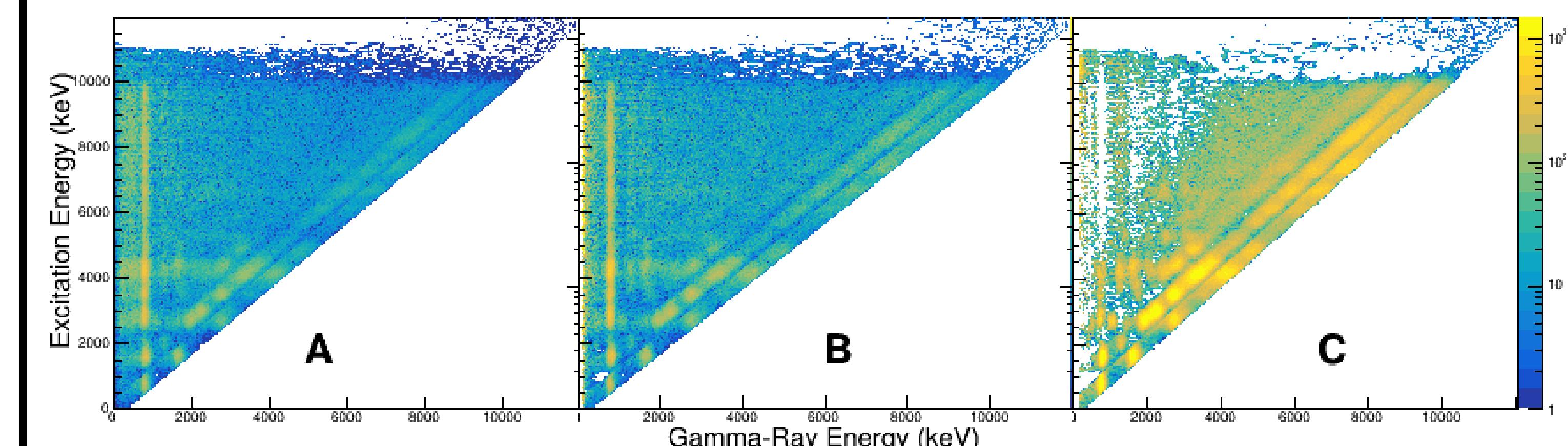
- Simulate gamma-rays of various energies



- Determine efficiencies, resolution and relative entries of various peaks
- Energy dependent shape of Compton background
- Interpolated for all energies



Oslo and Shape



- A: Raw experimental matrix
- B : Unfolded matrix – utilizes the response matrix
- C: Primary matrix - $P(E_\gamma, E_x) \propto \rho(E_x - E_\gamma) T(E_\gamma)$

- Absolute normalization: requires external data
 - NLD: D_0 , known levels
 - PSF: $\langle \Gamma_\gamma \rangle$ at S_n

