## Physics 305 – Sample Final Exam

There are four problems on this exam. Each problem is worth 25 points. Start each problem on a new sheet of paper, and use only one side of each sheet. GOOD LUCK !!!

- (1) A particle with charge q is moving with constant speed  $\beta c$  along the z axis. Use the tools that we developed in Chapter 12 to determine the electric and magnetic fields at an arbitrary point in space at the time *t*=0 when the particle passes through the origin. Express your answers in cylindrical coordinates.
- (2) Consider two thin, concentric spherical shells with radii *a* and *b* (with a < b). The region in between the two shells is filled with a material that has relative dielectric  $\varepsilon_r = 1$ , relative permeability  $\mu_r = 1$ , and conductivity  $\sigma$ . At t=0, there is a total charge  $+Q_0$  on the inner shell and  $-Q_0$  on the outer shell. Assume the conductivity is small enough so that retardation effects can be neglected.
- (a) Find the electric and magnetic fields throughout space for all times t>0. Be sure to justify any assumptions that you make.
- (b) Show that your fields from part (a) obey Maxwell's equations.
- (3) The Bohr planetary model of the atom was a natural outgrowth from Rutherford's discovery of the atomic nucleus. But there was a problem! Bohr's model had an electron with charge *-e* moving with constant speed *v* in a circle of radius *a* in the *x-y* plane, centered on the origin. Classically, such a system should radiate and lose energy. Calculate the intensity distribution of the emitted radiation,  $dP/d\Omega$ , and the average total power radiated. Note that, in the Bohr model,  $v \ll c$ , so the long-wavelength approximation is valid.

**Side comment**: The calculated total power isn't too far off for excited states in the hydrogen atom, but it fails miserably for the *ground* state.

(4) My former post-doc, now an Assistant Professor at James Madison University, would like to perform an experiment that requires a tunable, mono-energetic beam of 5-10 MeV gamma rays. Such a beam can be produced by shining light from a tunable dye laser operating in the IR or visible range (photon energies of <1 to a few eV) into a high energy electron beam (typical electron energies for this purpose are in the range of 0.5 to 5 GeV), as shown in the figure below. If Adriana places her target where it will see only those laser photons that backscatter off the electron beam in a narrow cone at 180<sup>o</sup>, it will see mono-energetic gamma rays. Calculate the energy of the gamma rays,  $E_{\gamma,f}$ , expressed in terms of the incident photon energy  $E_{\gamma,i}$ , the electron beam energy  $E_{e,i}$ , the mass *m* of the electron (note:  $mc^2 \approx 0.5$  MeV), and any physical constants that you need.

**Note**: Adriana needs to know the gamma ray energy to within  $\sim 0.1\%$  (*i.e.*, around 10 keV). So it will be okay to make approximations that change the answer less than that. If you find you must make an approximation that has a larger impact on your final answer, also estimate the error in your answer arising from your approximation.

 $\tilde{E}_{\gamma,i}$  $E_{e,i}$ 

Final Exam, Problem 1 We want the E+B fields in S, whereas we know  $\vec{E}^2 \frac{4}{4\pi\epsilon_0} \vec{F}_2$  and B=Oin S'. First, locate the common points in 4-space:  $Z' = Y(z - \beta x^{\circ}) = Y(z - \beta c \cdot 0) = Yz.$ x'\_= x → y'=y, , t is not needed because E+B are constant in S'. At (x', y', 2') = (x, y, 82):  $E_{z} = \frac{1}{4\pi\xi_{0}} \frac{1}{s^{2} + (\xi_{z})^{2}} \frac{1}{\sqrt{s^{2} + (\xi_{z})^{2}}}$  $E_{s}^{\prime} = \frac{1}{4\pi s_{0}} \frac{1}{s^{2} + (Y_{z})^{2}} \frac{s}{\sqrt{s^{2} + (x_{z})^{2}}}, E_{\varphi}^{\prime} = 0, B^{\prime} = 0$ Then  $\vec{E}_{11} = \vec{E}_{11} \Rightarrow \vec{E}_{2} = \frac{q}{4\pi\epsilon_{0}} \frac{82}{(s^{2}+(s^{2})^{2})^{3}/2}$  $\vec{E}_{i} = Y(\vec{E}_{i} + \vec{\sigma} \times \vec{B}) \Rightarrow \vec{E}_{i} = Y(\vec{E}_{i} - \vec{\sigma} \times \vec{B}) \Rightarrow$  $E_{s} = \frac{9}{4\pi \xi_{s}} \frac{\chi_{s}}{(s^{2} + (\chi_{z})^{2})^{3}/2}$  $= \frac{q}{E} = \frac{q}{4\pi\epsilon_{o}} \frac{\chi}{(s^{2} + (\chi_{z})^{2})^{3}/2} (s\hat{s} + z\hat{z}) = \frac{q}{4\pi\epsilon_{o}} \frac{\chi}{(s^{2} + (\chi_{z})^{2})^{3}/2}$  $\vec{B}'_{1} = \left\{ \vec{B}_{1} - \vec{\beta} \times \vec{E}_{1} \right\} \Rightarrow \vec{B}_{1} = \left\{ \vec{U} \times \vec{E}_{1} \right\} = \left\{ \vec{B}_{1} \times \vec{E}_{1} \right\}$  $= \frac{q_s \gamma_c^\beta}{4\pi \epsilon_o} \frac{s}{\left(s^2 + (\gamma_z)^2\right)^{3/2}} q$ 

Final Exam, Problem 2 (a) Let Q(t) be on the inner shell and -Q(t) on the outer. Then for  $a < r < b, \vec{E} = \frac{Q(t)}{4\pi\epsilon_0} \vec{f} \Rightarrow \vec{f} = \sigma \vec{E} = \frac{\sigma Q(t)}{4\pi\epsilon_0} \vec{f} \Rightarrow$  $-\frac{dQ}{dt} = \int \vec{J} \cdot d\vec{a} = \frac{\tau Q(t)}{y_{tf_{to}}} \frac{y_{tf_{to}}}{\tau^2} = \frac{\tau Q(t)}{z_0} \rightarrow$ dQ + T Q=O ⇒ (Q(+)=Qoe t 1 Thus, Ê(t) = Qo = zt p 4TEO e r2 for acr<b, and yor elsewhere. The averent is spherically to symmetric, so no angular direction differs from any other. Thus, we must have  $B=B_{p}\hat{r}$ . But  $\vec{z}.\vec{B}=0$ ⇒ \$B. da = O. Considering a solare of radius ~ then gives: B(t)=0 frall r (b) E is just the electrostatic solution for Q(t) at time t, so it obeys Gauss' haw ), and has (= x = 0 = - 2B), so Farabay's haw is bear. By construction, we have assured that (= B=O). Finally, Maxwell-Anypere's haw gives:  $\overrightarrow{T} \times \overrightarrow{B} = O = \mu_0 \overrightarrow{J} + \mu_0 \overleftarrow{\varepsilon}_0 \overrightarrow{E} = \mu_0 \overrightarrow{U} \frac{Q(t)}{4\pi \varepsilon_0} \frac{\widehat{\Gamma}}{\Gamma^2} + \mu_0 \overleftarrow{\varepsilon}_0 \frac{1}{4\pi \varepsilon_0} \frac{Q(t)}{2\tau} \frac{\widehat{\Gamma}}{\Gamma^2}$  $=\frac{h_0T}{4\pi\epsilon_0}Q(t)\frac{F}{r^2}+\frac{h_0}{4\pi}\left(\frac{-\sigma}{\epsilon_0}Q(t)\right)\frac{F}{r^2}=0$ and Maxwell-Ampere's how is also okay.

Final Exam, Problem 3 Assume t= O is chosen when the electron is on the x-axis. then: W(t)=aco(a) k + a sin (a) g. Thus, p(t) = Sr'p(t)dr = -ea(as(ot)(+sin(+a)y), and  $p(t) = e_{\alpha} \frac{\omega^2}{\alpha^2} \left[ \cos\left(\frac{\omega t}{\alpha}\right) \hat{x} + \sin\left(\frac{\omega t}{\alpha}\right) \hat{y} \right].$ 1 p = ev, independent of time, so the time - averaged total radiated power is given by: P= 10 p² = 10 e²ut 6TTC 6TTC2C For the angular distribution, <cos? = <sin? = 2, while < sin cos? = C Therefore, we can think of this as two in dependent, harmonic oscillators one in E and one in g, that all incoherently. Each radiates at sin & relative to the oscillation direction. The direction cosine relative to x is sind cos q >>  $\sin^2 \theta_{p} = 1 - \cos^2 \theta_{\chi} = 1 - \sin^2 \theta \cos^2 \varphi$ . Likewise,  $\sin^2 \theta_q = [-\cos^2 \theta_q = [-\sin^2 \theta \sin^2 \varphi, and]$   $\left\langle \frac{dP}{dR} \right\rangle = r^2 \left\langle \frac{s}{s} \right\rangle \cdot \hat{\boldsymbol{r}} = \frac{r_0 e^2 \upsilon^4}{32 \pi^2 c a^2} \left[ 1 - \sin^2 \theta \sin^2 \varphi \right]$  $=\frac{\mu_{0}e^{2}\upsilon^{4}}{32\pi^{2}ca^{2}}\left(2-\frac{1}{2}-\frac{1}{2}\omega^{2}\right)=\left(\frac{\mu_{0}e^{2}\upsilon^{4}}{32\pi^{2}ca^{2}}\left(1+\frac{1}{2}\omega^{2}\Theta\right)\right).$ A quick check shows Jazde = Prot, as it must.

Final Exam, Problem 4 Exi Eei Exf Eet Consofp BaitEei = Exf + Eet Exf Eet Pei - Pri = Pet + Prt  $= \sqrt{\frac{E_{ei}}{c^2} - \frac{2c^2}{mc}} - \frac{E_{st}}{c} = \sqrt{\frac{E_{ef}}{c^2} - \frac{2c^2}{mc^2}} + \frac{E_{sf}}{c}$  $\Rightarrow E_{ei}\left(1-\frac{1}{2}\frac{m^2c^4}{E_{ei}^2}\right) - E_{ri} \approx E_{ei}\left(1-\frac{1}{2}\frac{m^2c^4}{E_{ei}^2}\right) + E_{ri}$  $= \left( E_{ei} \right) - \frac{1}{2} \frac{m_{ei}^2 e^4}{E_{ei}} - E_{Yi} = \left( E_{ei} \right) - \frac{1}{2} \frac{m^2 e^4}{E_{ei}} + \left( E_{Yi} \right)$  $= \frac{1}{2} \frac{m^2 c^4}{E_{01}} - E_{01} = E_{01} - \frac{1}{2} \frac{m^2 c^4}{E_{01}}$  $\Rightarrow 4E_{s_1} + \frac{m^2 c^4}{E_{s_1}} = \frac{m^2 c^4}{E_{ef}} = \frac{m^2 c^4}{E_{s_1} + E_{e_1} - E_{s_f}}$ Exi <10 keV, so it can be discorded in the denominator. Then  $E_{e_i} - E_{Y_f} = \frac{m^2 e^4}{4E_{Y_i} + \frac{m^2 e^4}{E_{Y_i}}} = \frac{E_{e_i}}{1 + 4\frac{E_{x_i} E_{e_i}}{m^2 e^4}}$ > Exi+4 Esitei - Exp (1+4 Esitei) = Exp  $- = \frac{4E_{ri}E_{ei}^{2}}{\frac{1}{m_{e}^{2}+4E_{ri}E_{ei}}}$  $\Rightarrow \boxed{E_{x_{f}} = \frac{4 \frac{E_{ri}E_{ei}^{2}}{m^{2}c^{4}}}_{1 \neq 4} \frac{E_{ri}E_{ei}}{m^{2}c^{4}}}$