

Saturday Morning Physics 2008 at A&M: From the Micro-World to the Universe



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SMP-2008 Lecture 7

Texas A&M University, College Station, 29.03.07

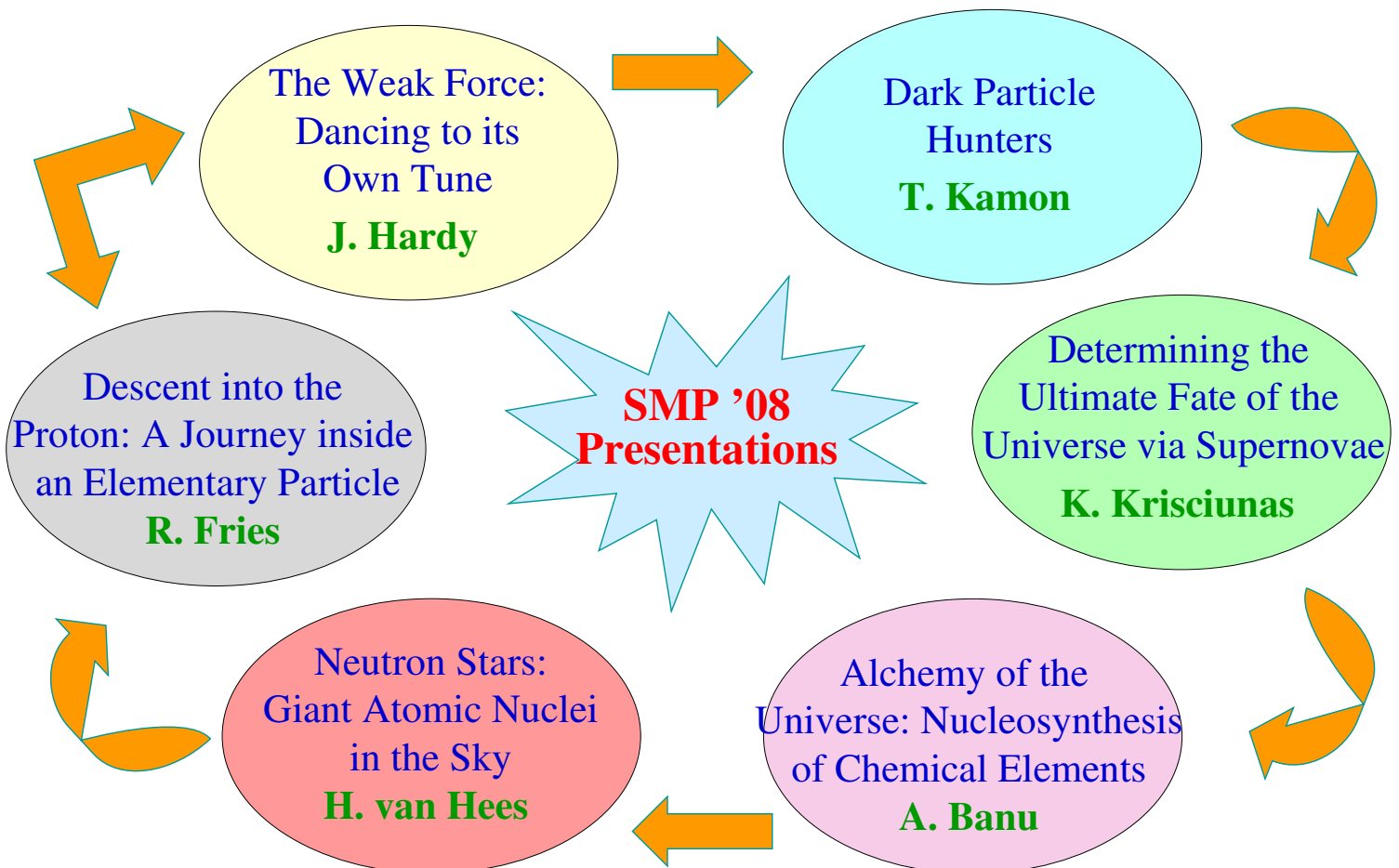
Outline

- 1.) Overall Structure of SMP08 Lectures
- 2.) From Particles+Forces to the Universe
 - Force Properties and Proton Structure
 - The Role of Fundamental Forces in the Cosmos
- 3.) Gravity and Dark Matter
- 4.) Weak Force and Universality
- 5.) Nucleosynthesis, Stars and the Universe's Fate
 - Supernovae: Nuclear Burning + Expansion of the Universe
 - Neutron Stars
- 6.) Concluding Remarks

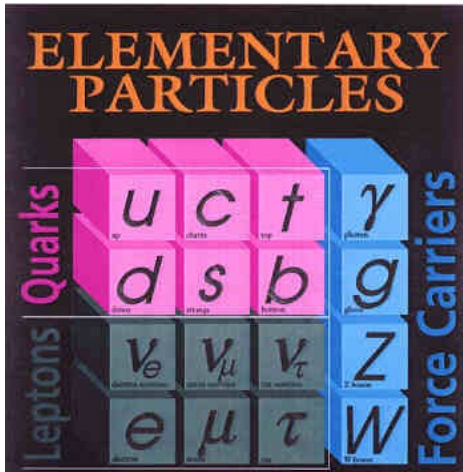
1.) From the Smallest to the Largest

- 20th (and 21st ?!) century: tremendous progress in our understanding of elementary particles + their interactions
- Also true for the origin and evolution of the Universe
- Intimate relations between the subatomic and the cosmic world unraveled
- early example (17th century): Newton discovered the connection between the falling apple and planetary motion (universality of gravity!)
- SMP2008 was an attempt to illuminate some of these fascinating discoveries in Nuclear/Particle/Astrophysics (and exhibit open problems ...)

1.2 Topical Structure of SMP08



2.) Known Particles + Forces in the Universe



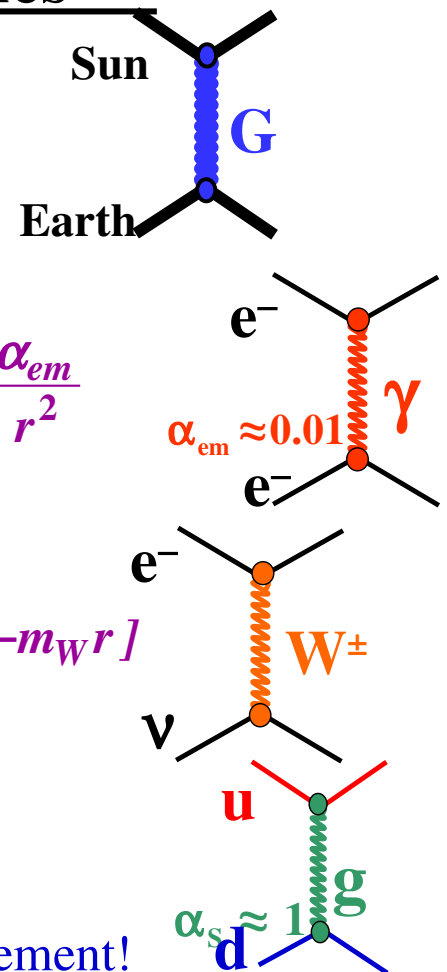
- “Standard Model of Particle Physics”
- based on symmetry principles:
matter particles interact via **force carriers**
 - stable matter: **u** , **d** , **e⁻** , **ν_e**
 - 2 more “generations” (heavier + short-lived)
 - bare masses: **m_{u,d} = 5 MeV ... m_t = 175 GeV**

	Gravity	Weak (Electroweak)	Electromagnetic	Strong
Carried By	Graviton (not yet observed)	W ⁺ W ⁻ Z ⁰	Photon	Gluon
Acts on	All	Quarks and Leptons	Quarks and Charged Leptons and W ⁺ W ⁻	Quarks and Gluons

- **forces differ by**
 - exchange particles
 - charge content
 - range
 - strength

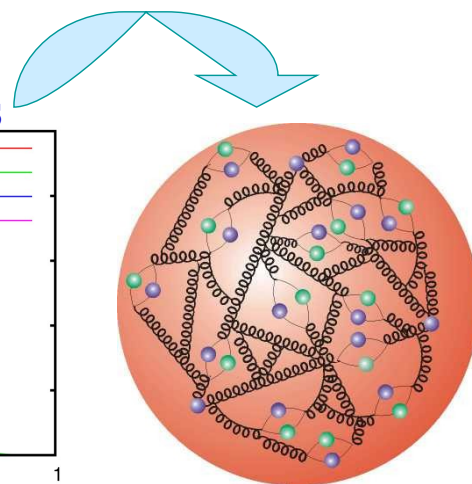
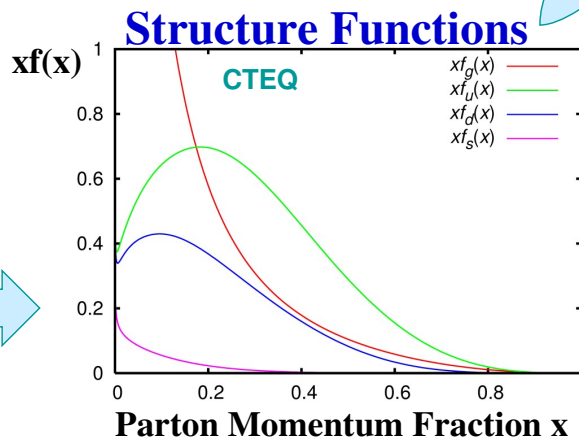
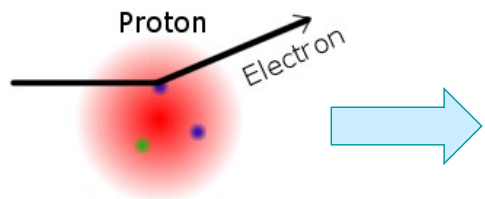
2.2 Fundamental Force Properties

- **Gravity (graviton exchange)**
 - static force law (Newton): $F_G(r) = -G_N \frac{m_1 m_2}{r^2}$
 - extremely small coupling
but only one charge \Rightarrow long range!
- **Electromagnetism (photon exchange)**
 - static force law (Coulomb): $F_G(r) = G_C \frac{q_1 q_2}{r^2} \sim \frac{\alpha_{em}}{r^2}$
 - small coupling, 1 anti-/charge
 \Rightarrow medium range
- **Weak Force (W⁻, Z-boson exchange)**
 - static force law (Fermi, GSW): $F_w(r) \sim \frac{\alpha_w}{r^2} \exp[-m_W r]$
 - small coupling, 2 anti-/charges
 - massive carriers \Rightarrow extremely short range
- **Strong Force (gluon exchange)**
 - static force law (QCD): $F_s(r) \sim -\alpha_s / r^2 - \sigma$
 - large coupl., 3 anti-/charges, short range, confinement!



2.3 Proton Structure

• Deep-Inelastic Electron Scattering

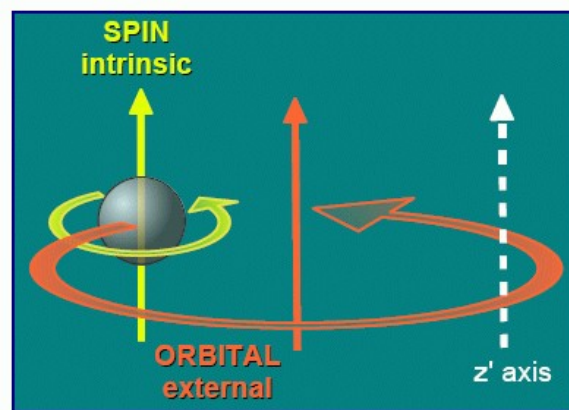


• Spin Structure of the Proton

$$\langle S_q \rangle + \langle S_g \rangle + \langle L_q \rangle + \langle L_g \rangle = \frac{1}{2}$$

quark spin $\langle S_q \rangle$
 gluon spin $\langle S_g \rangle$
 quark orbital angular momentum $\langle L_q \rangle$
 gluon orbital angular momentum $\langle L_g \rangle$

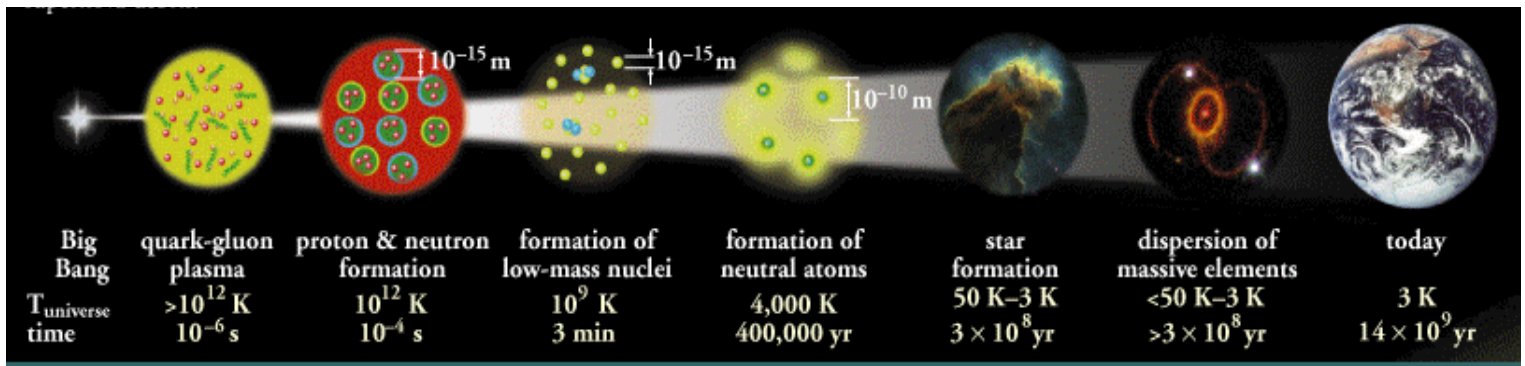
- quark- and gluon-spin contribution small
- quark orbital angular momentum?!



2.4 The Role of Particles + Forces in the Evolution + Structure of the Universe

- **Gravity** (long range)
 - large scale structure (galaxies, galaxy clusters)
 - star formation and collapse, black holes
- **Electromagnetism** (medium range)
 - neutralization of electrons/nuclei 400,000y after Big Bang
 - cosmic microwave background $T_{\text{Universe}} = 2.73^\circ\text{K}$
 - γ -ray bursters (the largest fire crackers in the Universe)
- **Weak Force** (very short range)
 - generation of bare masses 10^{-11}s after Big Bang
 - star cooling (neutrinos), element transmutation ($p \leftrightarrow n$)
- **Strong Force** (short range)
 - generation of visible mass 10^{-6}s after Big Bang
 - star burning and explosion, neutron-star structure

Nuclear Physics and the Universe



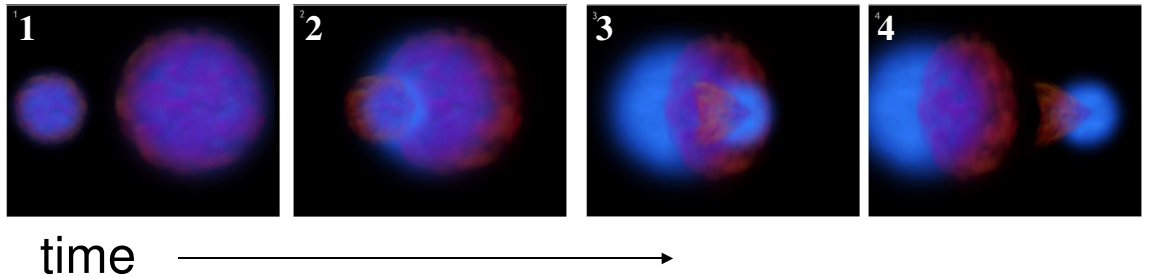
- Quark-Gluon Plasma: $T > 200 \text{ MeV}$ (< 0.000001 sec.)
- Phase transition to Hadronic Matter (Mass Generation, Quark Confinement), $T \approx 170 \text{ MeV}$ (0.00001 sec.)
- Low-mass nuclei: H (p), d (pn), ^3He , ^4He , ^7Li (3 min.)
- Heavy elements in star collapses: supernovae (still today)
- Exotic forms of (quark) matter in neutron stars (still today)

3.) Gravitational Force and Dark Matter

- Evidence for Dark Matter
- Supersymmetry?!
(Neutralino!)

3.1 Dark Matter Evidence and Properties

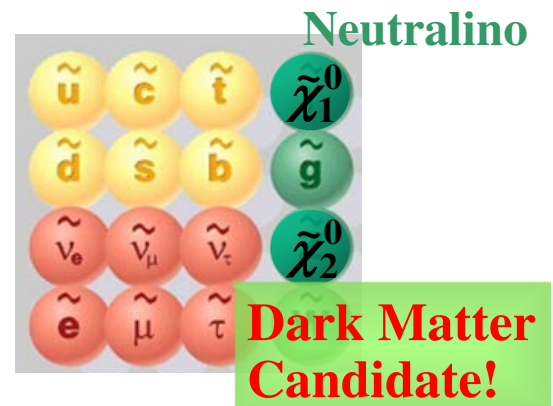
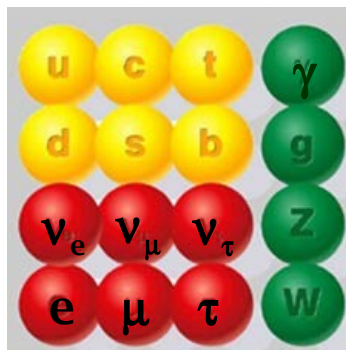
Cosmic
collision of 2
galaxy clusters:
DM unaffected!



Dark Matter Properties: - very weakly interacting, charge-neutral
- slowly moving (“cold”), stable+heavy particle
⇒ no such particle in the Standard Model, new idea needed!

Supersymmetry:

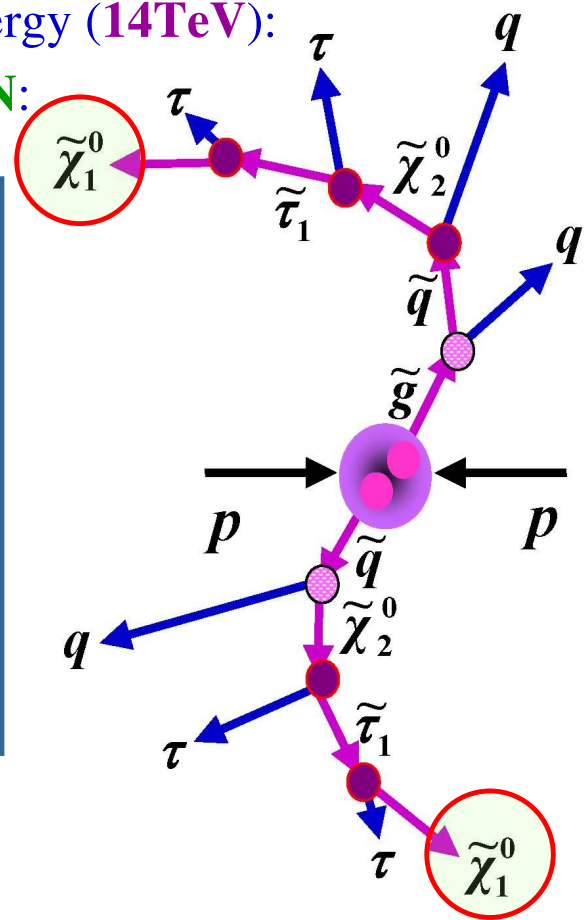
- fermion ↔ boson partners for all standard-model particles
- Supersymmetry “broken”:
 $m_{\text{stand}} \ll m_{\text{super}} \sim 1\text{TeV}/c^2$



3.2 How to Measure Dark Matter in the Lab?

- proton-proton collisions at the highest energy (14TeV):

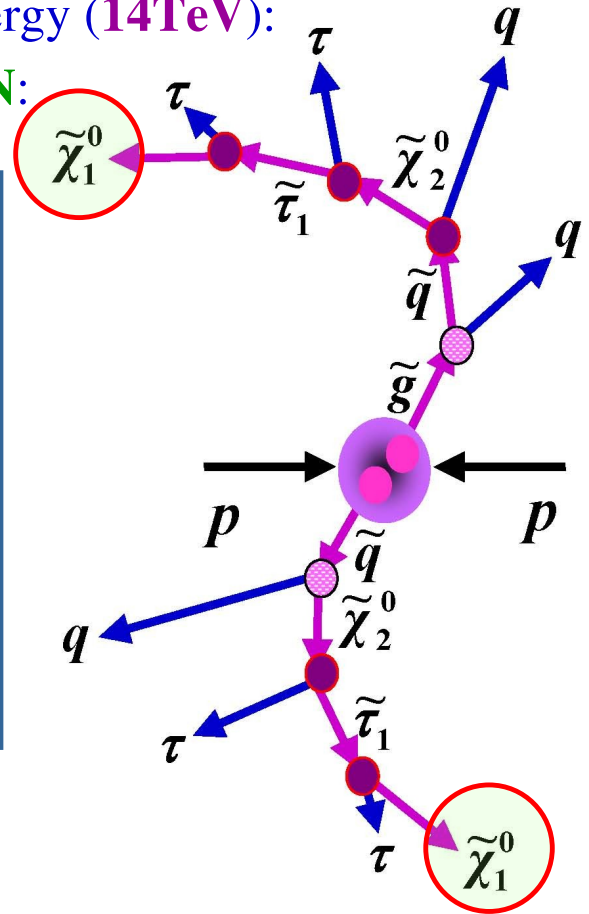
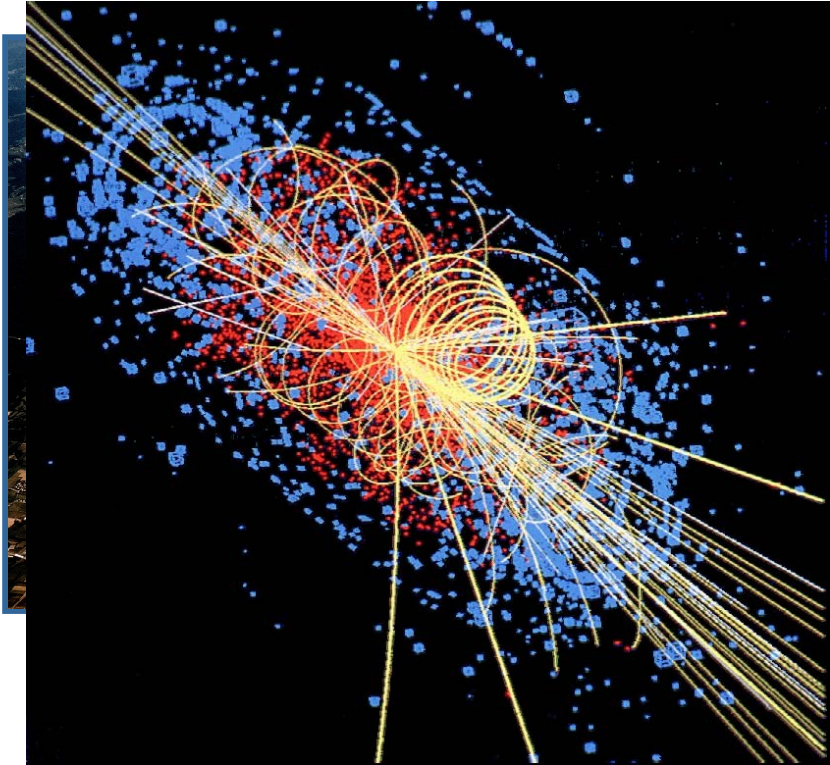
Large Hadron Collider (LHC) at CERN:



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4.) Precision β -Decay: Testing the Weak Force

Recall: around ~1700 Sir Isaac Newton realized the universality of gravity



Is the Weak Force universal, too?

Use precision measurements of nuclear beta-decay to check

We know now that ...

1. The weak force (vector component) is constant in nuclei to 0.026%.
2. We can also test full universality of the weak force -- including the decay of other particles like the kaon -- but the jury is still out on this one!
3. Nuclear physics is the source of key data for these tests, the most precise ones available.

5.) Nucleo-Synthesis, Stars + the Universe

- Nuclear Burning
- Supernovae
 - White-Dwarf Explosions (type-Ia)
 - Heavy-Star Explosions (type-II)
- Fate of the Universe
- Neutron Stars

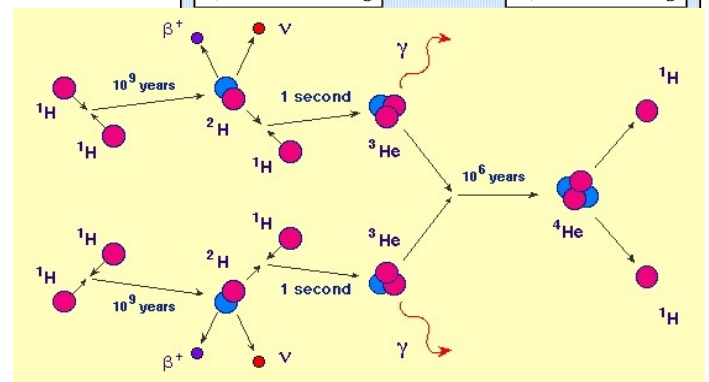
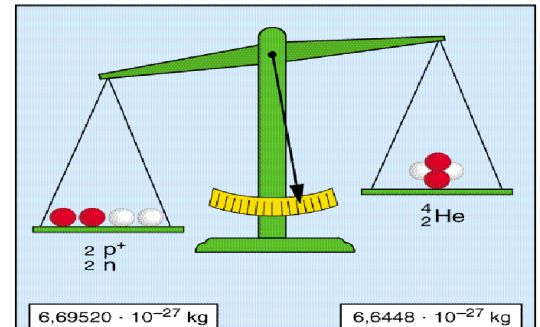
5.1 Nuclear Burning

- Principle: large energy gain if light nuclei

“fuse”: $A + B \rightarrow C + \text{binding energy}$

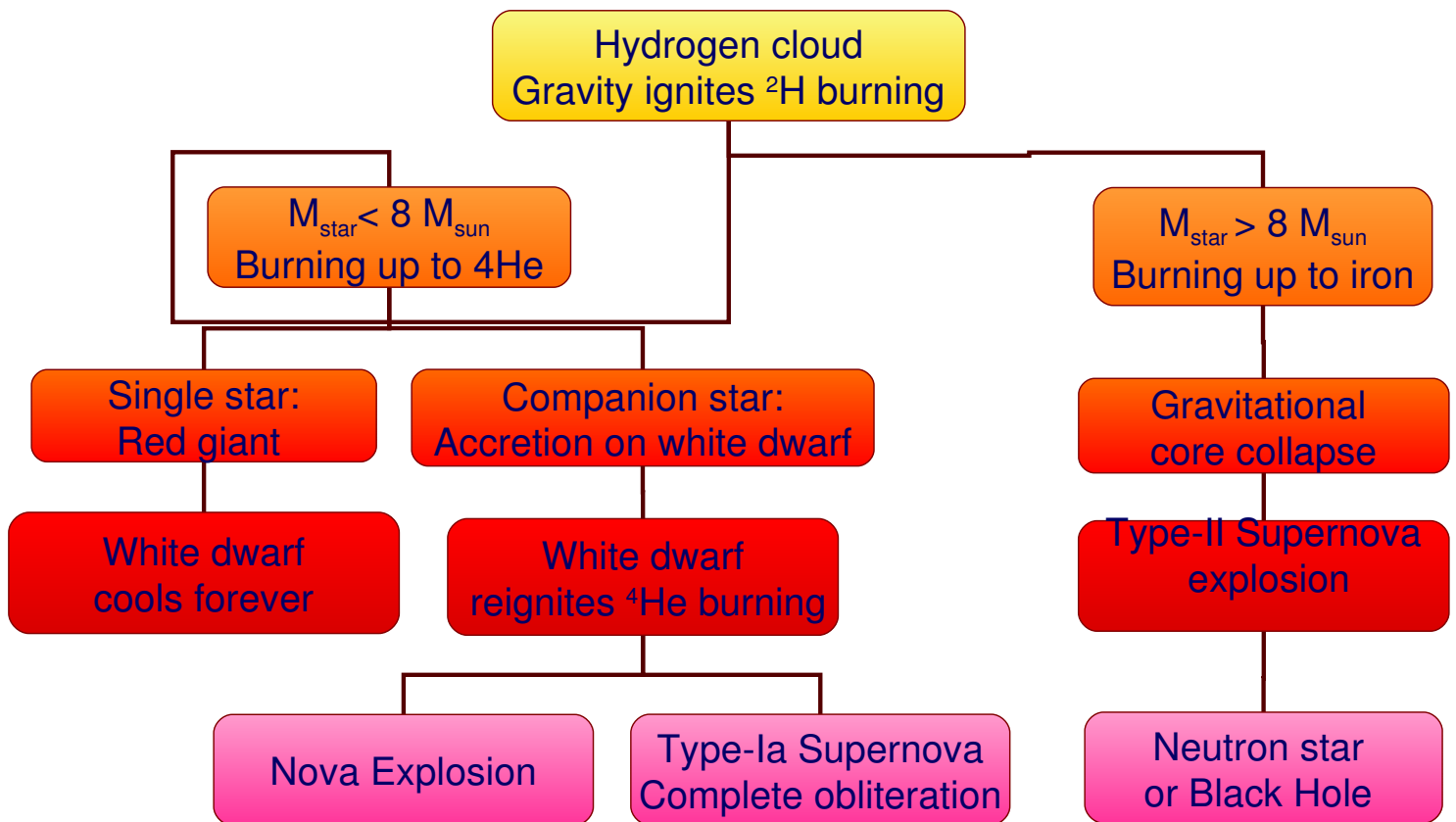
$$E_B = [(M_A + M_B) - M_C] c^2$$

- “problem”: Coulomb repulsion between **A** and **B**
- **A** and **B** need to “touch” to feel strong force and bind
- large temperature and density required for nuclear fusion (burning)!



- Big Bang: rapid expansion, only up to mass **A=7 (Li)**, gaps at **A=5,8!**
- much later: gravity-driven star formation
 - Sun ($M \sim M_{\odot}$) $p + p \rightarrow d + p \rightarrow {}^3\text{He} + {}^3\text{He} \rightarrow {}^4\text{He} + 2p$
 - heavy stars: $3 {}^4\text{He} \rightarrow {}^{12}\text{C} + {}^4\text{He} \rightarrow {}^{16}\text{O} + {}^4\text{He} \rightarrow {}^{28}\text{Si} + {}^4\text{He} \rightarrow {}^{56}\text{Co} \rightarrow {}^{56}\text{Fe}$
 - star collapse and explosion: all elements beyond **A=56** (no energy gain)

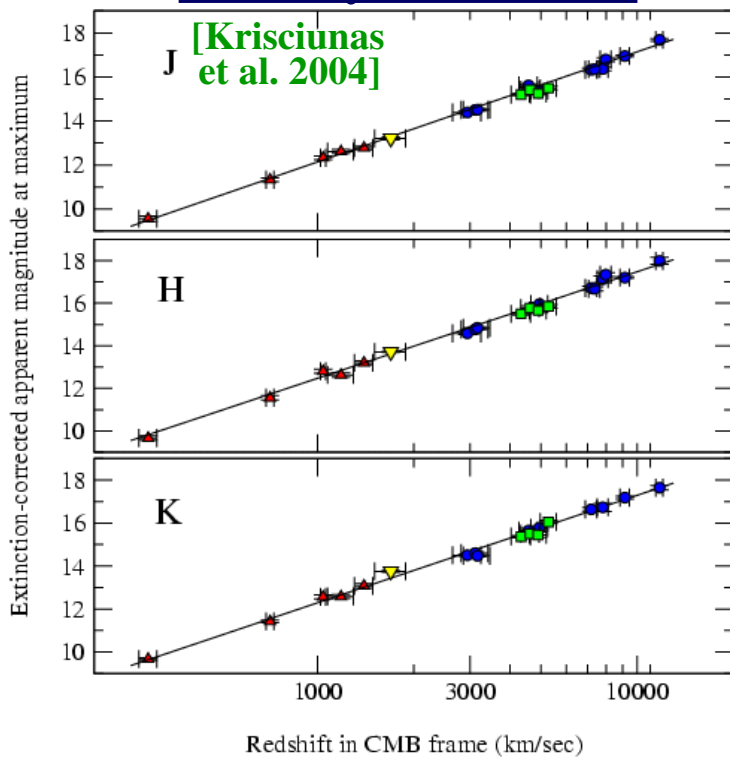
5.2 Star Evolution Chart



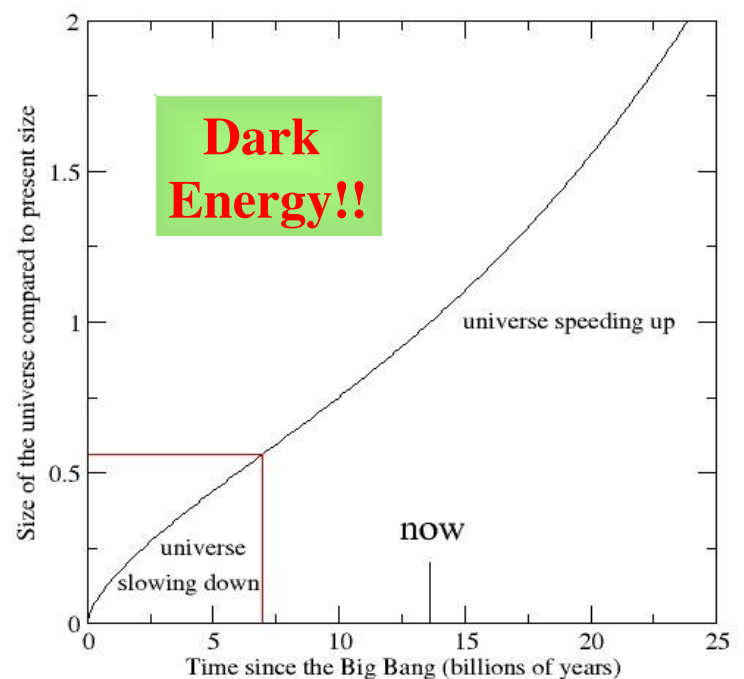
5.3 SN-Ia Candles + Expansion of the Universe

- accurate light output I_0 , intensity $I(r) = I_0/4\pi r^2 \Rightarrow$ precise distance $r(I)$
- Doppler (red-) shift of spectral lines \Rightarrow recession velocity, v_r , of source

Intensity vs. Redshift

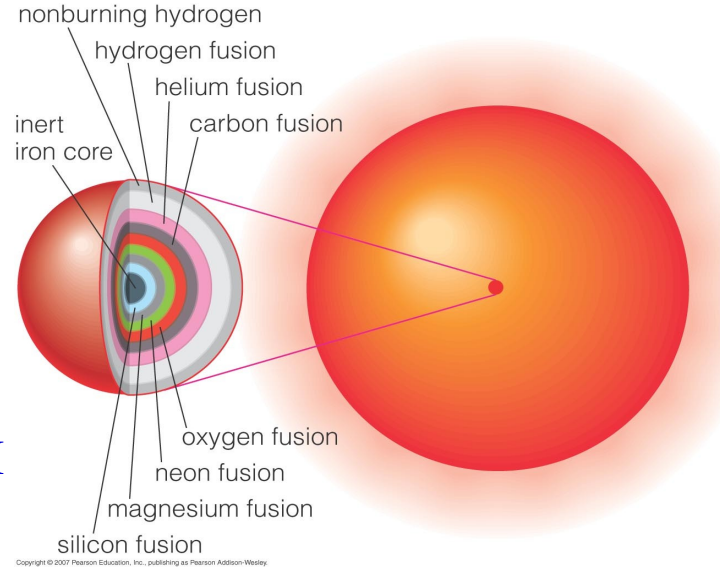
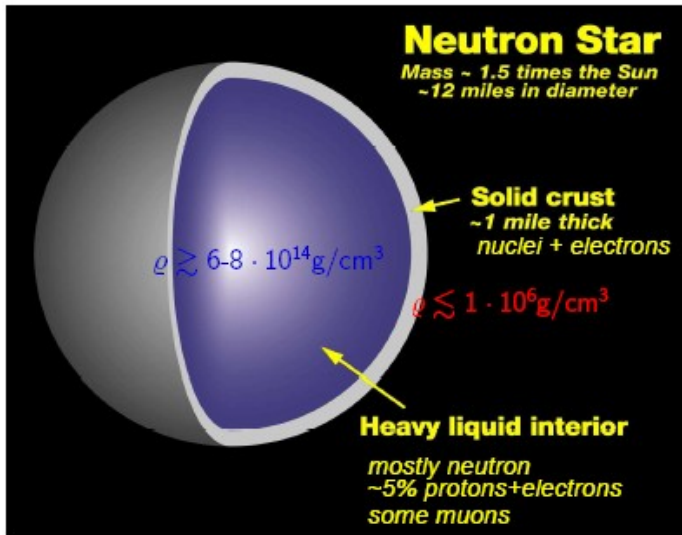


Accelerating Universe



5.4 Type II Supernovae

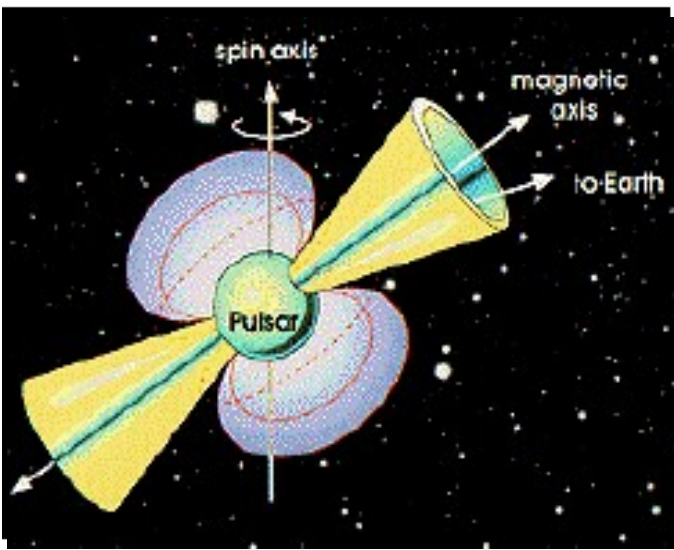
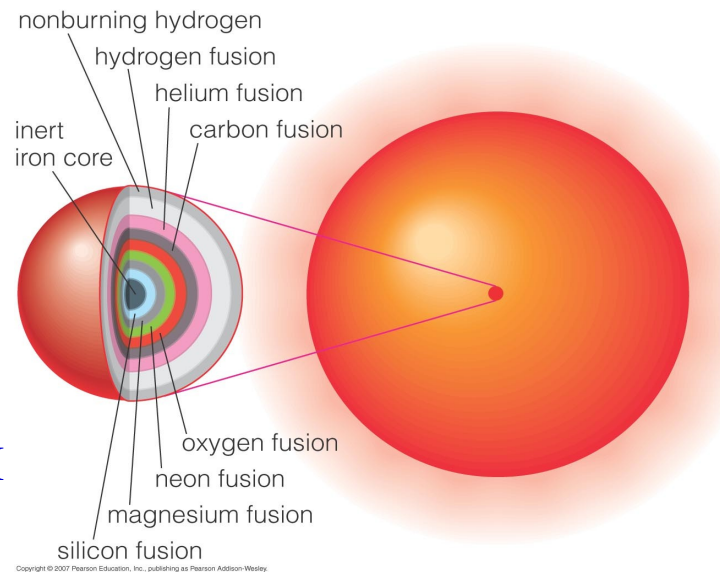
- High-mass star ($M_{\text{star}} > 8M_{\odot}$),
- burns fast ($\sim 50 \text{ My}$), up to ^{56}Fe
 \Rightarrow core collapse
 \Rightarrow type-II supernova explosion
- produces all known heavy elements
- leaves behind moving+rotating NS/BH



- $M_{\text{NS}} \approx 1.4 M_{\odot}$, but $R=15\text{km}$
- up to 5-10 times density of nuclei!!
- study cold nuclear equation of state (quark plasma, color-superconductor ...)
- rotation: lighthouse (mag. field), glitches
- γ -ray emission (bursters?)
- general relativity (grav. waves?)

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6.) Some Perspectives for You

If you

- Enjoy / are excited by Physics / Science
- Tend to be curious
- Like to try things out **AND/OR** like math, computers

then we recommend to:

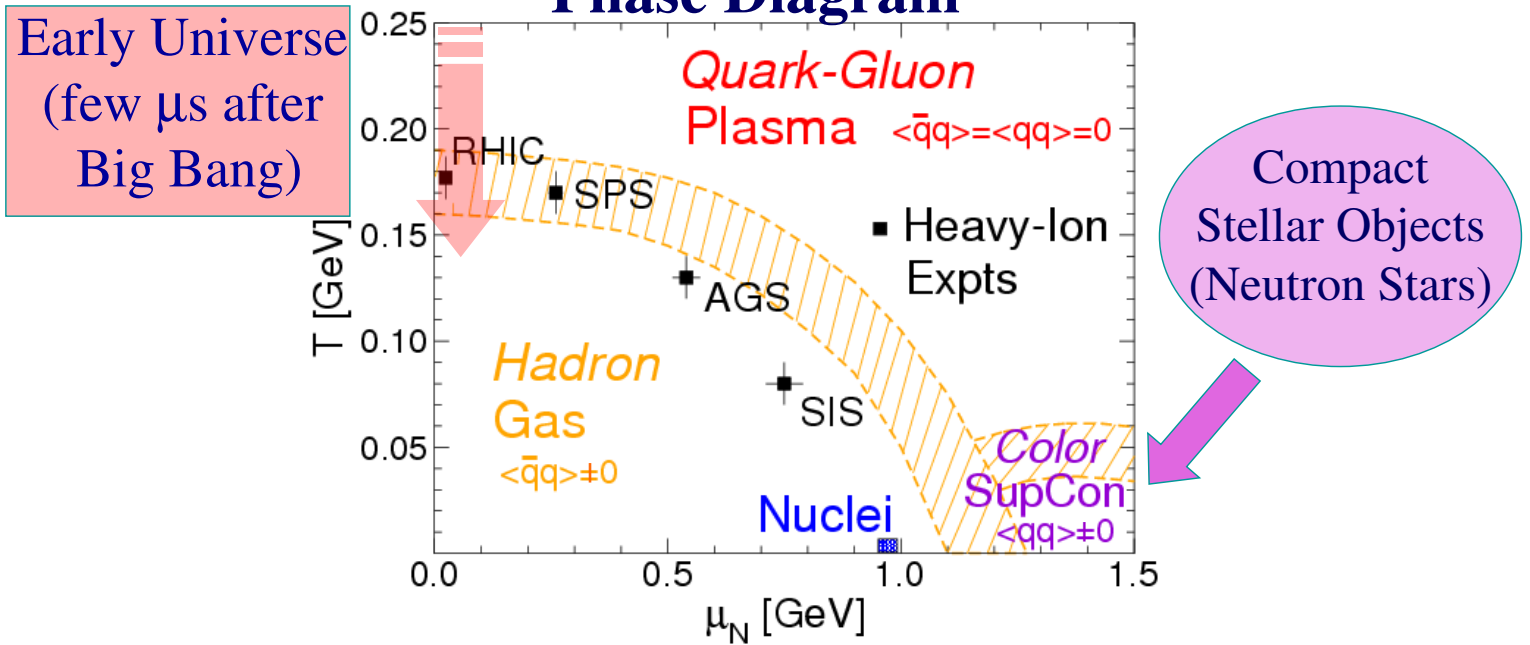
- Watch out for future SMP Series at A&M (2+ more)
- Consider enrolling in the Physics Undergraduate Program at A&M
- Inform yourself about future career paths in Physics

6.2 Thanks to:

- You! (students)
- Our high school teachers!
- Our lecturers: Profs. John Hardy, Teruki Kamon, Kevin Krisciunas, Hendrik van Hees, Rainer Fries, Dr. Adriana Banu
- The staff support team: Kendra Beasley, Shana Hutchins, Bruce Hyman, Leslie Spikes, Sharon Jeske, Tony Ramirez, Jerry Deason
- The SMP organizing team: Hendrik van Hees, Xingbo Zhao, Trent Strong, Saskia Mioduszewski, Rainer Fries + Adriana Banu
- Financial Support: U.S. National Science Foundation,
Texas A&M Cyclotron Institute + Physics Department

2.1 Hot+Dense QCD Matter in Nature

Phase Diagram



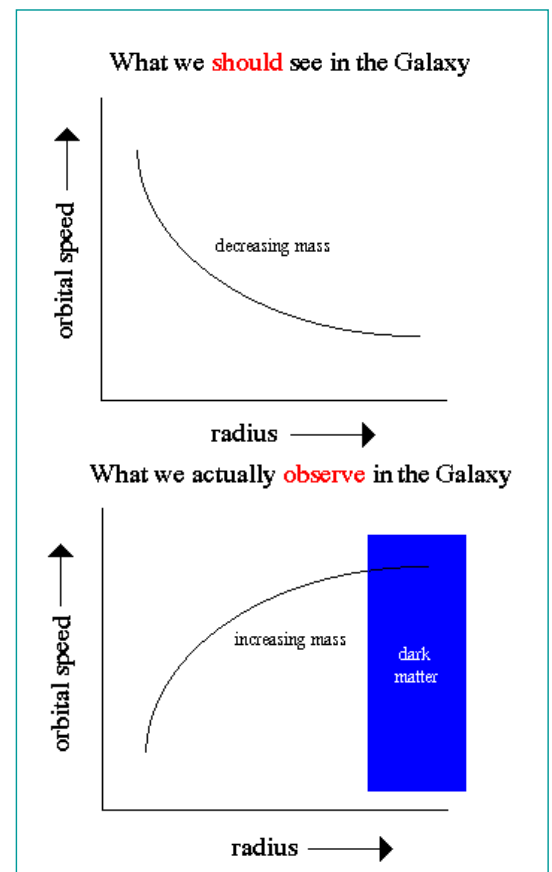
In the laboratory: high-energy collisions of heavy nuclei!

Objective: to create matter at temperatures $T > T_c \approx 170 \text{ MeV}$
and energy densities $\varepsilon > \varepsilon_c \approx 1 \text{ GeV fm}^{-3}$

3.1 Evidence for Dark Matter I

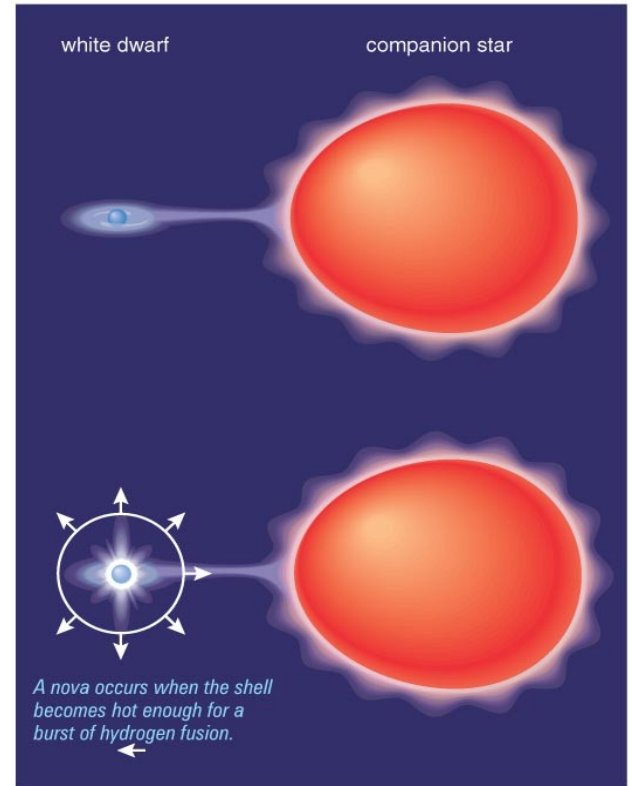
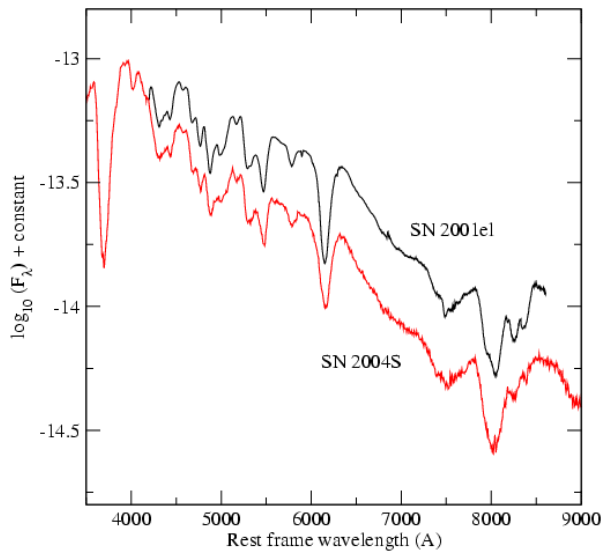


- motion of stars within galaxies:
there must be more matter than
we “see” (emits light)
⇒ **Dark Matter**:
 - “background”?
 - new particles?



5.3 Novae and Type-Ia Supernovae

- White dwarf accretes matter from red-giant companion (binary system)
 - (i) helium burning ignited on surface
⇒ **Nova** explosion
 - (ii) mass accretion up to $1.4 M_{\odot}$
⇒ **Type-Ia Supernova** explosion,
extremely regular light output:



a Diagram of the nova process.

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