

Isoscalar Giant Dipole Resonance and Nuclear Matter Incompressibility Coefficient

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Studies of compression modes of nuclei are of particular interest since their strength distributions, $S(E)$, are sensitive to the value of the nuclear matter incompressibility coefficient, K [1]. At present, Hartree-Fock (HF) based random-phase-approximation(RPA) calculations for the isoscalar giant monopole resonance (ISGMR) reproduce the experimental data for effective interactions associated with incompressibility $K = 210 \pm 20$ MeV.

The study of the isoscalar giant dipole resonance (ISGDR) is very important since this compression mode provides an independent source of information on K . Early experimental attempts to identify the ISGDR in ^{208}Pb resulted with a value of $E_1 \sim 21$ MeV for the centroid energy. Very recent and more accurate data on the ISGDR obtained at our Cyclotron Institute for a wide range of nuclei seems to indicate that the experimental values for E_1 are smaller than the corresponding HF-RPA results by 3–5 MeV. This discrepancy between theory and experiment raises doubts concerning the unambiguous extraction of K from energies of compression modes.

In this work we address this discrepancy between theory and experiment by examining the relation between the strength function $S(E)$ and the excitation cross section $\sigma(E)$ of the ISGDR, obtained by ∇ -scattering. We emphasize that it is quite common in theoretical work on giant resonance to calculate $S(E)$ for a certain scattering operator F whereas in the analysis of experimental data of $\sigma(E)$ one carries out distorted-wave-Born-approximation (DWBA) calculations with a certain transition

potential. Here we present results of accurate microscopic calculations for $S(E)$ and for $\sigma(E)$ with the folding model (FM) DWBA with transition densities $\Delta_i(\mathbf{r})$ obtained from HF-RPA calculations and suggest a simple explanation for the discrepancy between theory and experiment concerning the ISGDR.

In self-consistent HF-RPA calculation one starts by adopting specific effective nucleon-nucleon interaction, V_{12} , carries out the HF calculation for the ground state of the nucleus and then solves the RPA equation using the particle-hole (p-h) interaction V_{ph} which corresponds to V_{12} . The RPA Green's function G is obtained from

$$G = G_0(1 + V_{ph}G_0)^{-1}, \quad (1)$$

where G_0 is the free p-h Green's function. For

$$F = \sum_{i=1}^A f(\mathbf{r}_i), \quad (2)$$

the strength function and transition density are given by

$$S(E) = \sum_n |\langle 0|F|n \rangle|^2 \delta(E - E_n) = \frac{1}{\pi} \text{Im}[\text{Tr}(fGf)], \quad (3)$$

$$\rho_i(\mathbf{r}, E) = \frac{\Delta E}{\sqrt{S(E)\Delta E}} \times \int f(\mathbf{r}') \left[\frac{1}{\pi} \text{Im} G(\mathbf{r}', \mathbf{r}, E) \right] d\mathbf{r}' \quad (4)$$

Note that (4) is consistent with the strength in the region $E \pm)E/2$ and is consistent with

$$S(E) = \left| \int \rho_t(\mathbf{r}, E) f(\mathbf{r}) d\mathbf{r} \right|^2 / \Delta E. \quad (5)$$

In fully self-consistent HF-RPA calculations, the spurious state (associated with the center of mass motion) $T = 0, L = 1$ appears at $E = 0$ and no spurious state mixing (SSM) in the ISGDR occurs. However, although not always stated in the literature, actual implementations of HF-RPA (and relativistic RPA) are not fully self-consistent. One usually makes the following approximations: (i) neglecting the two-body Coulomb and spin-orbit interactions in V_{ph} , (ii) approximating momentum parts in V_{ph} , (iii) limiting the p-h space in a discretized calculation by a cut-off energy E_{ph}^{\max} , and (iv) introducing a smearing parameter (i.e., a Lorentzian with $\vartheta/2$). Although the effect of these approximations on the centroid energies of giant resonances is small (less than 1 MeV), the effect on the ISGDR is quite serious since each of these approximations introduces a SSM in the ISGDR.

Recently [1,2] we have shown that in order to correct for the effects of SSM on $S(E)$ and the transition density we use the projection operator

$$F_\eta = \sum_{i=1}^A f_\eta(\mathbf{r}_i) = F - \eta F_1, \quad (6)$$

With $f_0 = f - Of_1 = (r^3 - Or)Y_{1M}(\Sigma)$. The value of O associated with the coherent state transition density

$$\rho_{ss} = \frac{\partial \rho_0}{\partial r} Y_{1M}(\Omega), \quad (7)$$

where Δ_0 (is the ground state density of the nucleus, is given by

$$\eta = \langle f_3 \rho_{ss} \rangle / \langle f_1 \rho_{ss} \rangle = \frac{5}{3} \langle r^2 \rangle \quad (8)$$

To determine the transition density Δ_t for the ISGDR we use (4) with F_0 and obtain Δ_0 then project out the spurious term by making use of (7)

$$\rho_t(\mathbf{r}) = \rho_\eta(\mathbf{r}) - \alpha \rho_{ss},$$

$$\alpha = \langle f_1 \rho_n \rangle / \langle f_1 \rho_{ss} \rangle \quad (9)$$

We have carried out [1] numerical calculations for the $S(E)$, $\Delta_t(\mathbf{r})$ and $\Phi(E)$ within the FM-DWBA-HF-RPA theory. We used the SLI Skyrme interaction, which is associated with $K = 230$ MeV, and carried out HF calculations using a spherical box of $R \geq 25$ fm. For the RPA calculations we used the Green's function approach with mesh size $\Delta r = 0.3$ fm and p-h maximum energy of $E_{ph}^{\max} = 150$ MeV (we include particle states with principle quantum number up to 12), since it is well-known that in order to extract accurate $\Delta_t(\mathbf{r})$, E_{ph}^{\max} should be much larger than the value required ($E_{ph}^{\max} \sim 50$ MeV) to recover EWSR. Since in our calculation we also neglected the two-body coulomb and spin-orbit interactions, the spurious state energies differ from 0 by a few MeV. We therefore renormalized the strength of the V_{ph} by a factor (0.99 and 0.974 for ^{116}Sn and ^{208}Pb , respectively), to place the spurious state at $E=0.2$ MeV. We have included a Lorentzian smearing ($\vartheta/2 = 1$ MeV) and corrected for the SSM as described above. We carried out the

FM-DWBA calculation for $\Phi(E)$ using a density dependent Gaussian nucleon- ∇ interaction with parameters adjusted to reproduce the elastic cross section, with Δ_0 and Δ_t from HF-RPA.

Using the operator $f = r^2$ for the ISGMR we calculated the corresponding $S(E)$, for E up to 60 MeV. We recover 100% of the corresponding EWSR and obtained the values of 17.09 and 14.48 MeV for the centroid energy of the ISGMR in ^{116}Sn and ^{208}Pb , respectively. The corresponding recent experimental values obtained at our Institute are 16.07 ± 0.12 and 14.17 ± 0.28 MeV, respectively.

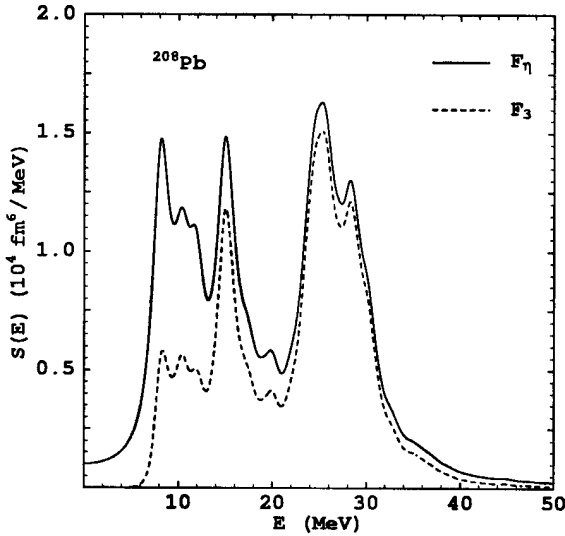


Figure 1: Strength functions for the ISGDR in ^{208}Pb obtained from Eqs. (4), (9) and (5), using f_3 (dashed line) and $f_\eta = f_3 - \eta f_1$ (solid line), with $\theta = 52.1 \text{ fm}^2$.

Figure 1 exhibits the strength functions for the ISGDR in ^{208}Pb obtained from Eqs. (5), (4) and (9). The solid line describes the result obtained using f_0 . Note that this result coincides with $S_\theta(E)$, which is free of SSM contribution. Similarly, the dashed line describes the erroneous result obtained using f_3 (it is also different from $S_3(E)$). We find that when using f_3 , the excitation strengths obtained for certain states are sensitive to the value of ε . The result obtained with f_3 coincides with that obtained

with f_0 for $\varepsilon \rightarrow 0$, as expected. Thus, in configuration RPA calculation of Δ_t , one may use f_3 and correct for the SSM contribution before the smearing process.

Our results for the ISGDR, $S_\theta(E)$, indicate two main components with the low energy component containing close to 30% of the EWSR (for E up to 23 and 19 MeV for ^{116}Sn and ^{208}Pb , respectively), in agreement with the experimental observation [3].

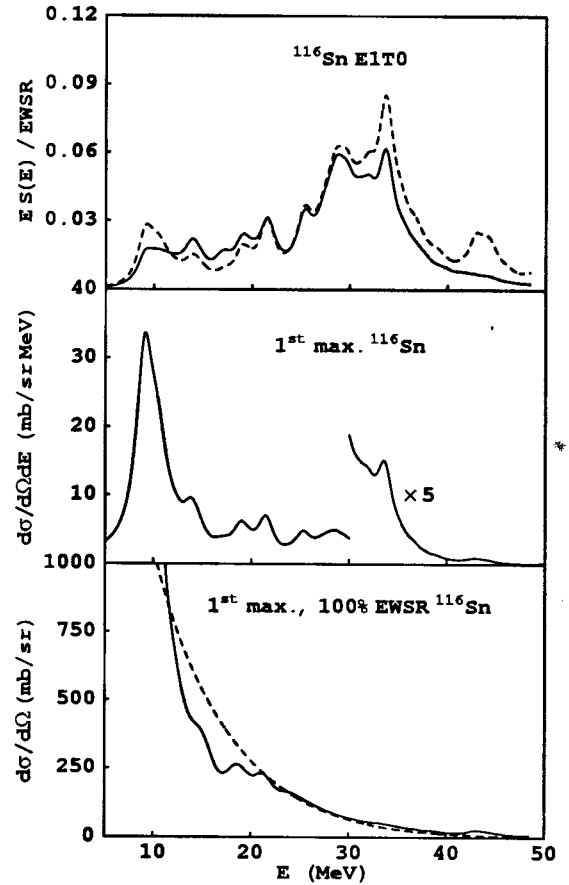


Figure 2: The ISDGR in ^{116}Sn . The middle panel: maximum double differential cross section obtained from Δ_t (RPA). The lower panel: maximum cross section obtained with Δ_{coll} (dashed line) and Δ_t (solid line) normalized to 100% of the EWSR. Upper panel: The solid and dashed lines are the ratios of the middle panel curve with the solid and dashed lines of the lower panel, respectively.

In Figure 2 we present results of microscopic calculations of the excitation cross

section of the ISGDR in ^{116}Sn by 240 MeV ∇ -particle, carried out within the FM-DWBA. The dashed lines are obtained using $\Delta_{coll}(r)$ of the ISGDR. It is seen from the upper panel that the use of Δ_{coll} increases the EWSR by at least 10% and may shift the centroid energy by a few percent. An important result of our calculation is that the maximum cross section for the ISGDR drops below the current experimental sensitivity of 2 mb/sr/MeV for excitation energy above 35 MeV (30 MeV for ^{208}Pb), which contains about 20% of the EWSR. This missing strength leads to a reduction of more than 2.5 MeV in the ISGDR energy and thus explains the discrepancy between theory and experiment. More sensitive experiments and/or with higher ∇ -particle energy are thus needed.

In summary, we developed and applied an accurate and general method to eliminate the

SSM contributions from $S(E)$ and Δ_i . Our results indicate: (i) Existence of non-negligible ISGDR strength at low energy and (ii) Accurate determination of the relation between $S(E)$ and $\Phi(E)$ resolves the long standing problem of the conflicting results obtained for K , deduced from experimental data $\Phi(E)$ for the ISGDR and data for the ISGMR.

References

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